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Detection of Kardar-Parisi-Zhang hydrodynamics in a quantum Heisenberg spin-1/2 chain

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Classical hydrodynamics is a remarkably versatile description of the coarse-grained behavior of many-particle systems once local equilibrium has been established¹. The form of the hydrodynamical equations is determined primarily by the conserved quantities present in a system. Quantum spin chains are known to possess, even in the simplest cases, a greatly expanded set of conservation laws, and recent work suggests that these laws strongly modify collective spin dynamics even at high temperature^{2,3}. Here, by probing the dynamical exponent of the one-dimensional Heisenberg antiferromagnet KCuF_3 with neutron scattering, we find evidence that the spin dynamics are well described by the dynamical exponent $z = 3/2$, which is consistent with the recent theoretical conjecture that the dynamics of this quantum system are described by the Kardar-Parisi-Zhang universality class^{4,5}. This observation shows that low-energy inelastic neutron scattering at moderate temperatures can reveal the details of emergent quantum fluid properties like those arising in non-Fermi liquids in higher dimensions.

Interacting magnetic moments (“spins”) governed by the laws of quantum mechanics can exhibit a vast set of complex phenomena such as Bose-Einstein condensation and superfluidity⁶, topological states of matter⁷, and exotic phase transitions^{8,9}. Understanding quantum magnets is therefore a challenging task, connecting experiments with intensive theoretical modeling and state-of-the-art numerical simulations. In that respect, linear arrangements of spins (“spin chains”) at temperatures close to absolute zero have been influential because the one-dimensional (1D) setting produces especially prominent quantum fluctuations¹⁰.

The most celebrated model magnetic system realized in nature is the Heisenberg spin-half chain, where isotropic magnetic moments are coupled by a nearest-neighbor antiferromagnetic exchange interaction of strength J . It is characterized by fractional quasiparticles excitations called spinons (Fig. 1b) with a dispersion relation given by $\hbar\omega(Q) = J\frac{\pi}{2}|\sin(Qa)|$, where a is the lattice spacing unit (following convention for 1D chains, we set $a = 1$ in this study). They are responsible for the physical properties of the system and can be identified by the dynamical spin response function, as measured in inelastic neutron spectroscopy. Spinons are created in pairs and interact with one another, leading to a continuum in the neutron scattering spectrum.

In fact, in the ground state, two-spinon states accounting for 71% of the total spectral weight have an upper bound $\hbar\omega(Q) = J\pi|\sin(Q/2)|$, which gives its distinctive shape to the spectrum (Fig. 2a), and including the four-spinon contribution on top of the two-spinon exhausts 98% of the weight¹¹, etc.

As temperature increases and many spins are excited, the spin dynamics at frequencies $\hbar\omega \ll k_B T$ is usually thought of in terms of collective thermal rather than quantum effects. This high-temperature regime has not been the focus of experimental study, but recent theoretical progress in 1D quantum systems suggests that it nevertheless holds precious information on the underlying quantum features^{2,3,12}. One can make an analogy with the phenomenological derivation of the equations of fluid dynamics, based on the continuity equations of conserved quantities (such as mass, energy, or momentum): depending on the intrinsic quantum conservation laws of the system, one expects the emergence of different kinds of coarse-grained hydrodynamic behaviors for the spins at high-temperature. Remarkably, some 1D quantum systems, known as integrable — including the Heisenberg spin-half chain — possess an infinite number of nontrivial conserved quantities. They strongly constrain the overall dynamics of integrable systems and can endow them with peculiar hydrodynamic properties, some of which have been observed experimentally in a 1D cloud of trapped ⁸⁷Rb atoms¹³. In the case of magnets, three universal regimes have been identified^{14,15} and are classified by the dynamical exponent z , governing the length-time scaling, i.e., $\text{LENGTH} \sim \text{TIME}^{1/z}$: $z = 2$ corresponds to diffusion, $z = 1$ to ballistic, and $z = 3/2$ to superdiffusive dynamics (Fig. 1c).

The presence of ballistic spin dynamics in integrable systems is theoretically established by showing that spin currents \hat{j}_s are dissipationless, resulting in a perfect spin conductor with no resistance (or equivalently an infinite spin dc conductivity). Quantitatively, this can be achieved by looking at the long-time asymptote of the spin current-current correlation $\lim_{t \rightarrow \infty} \langle \hat{j}_s(t) \hat{j}_s(0) \rangle$, where saturation to a nonzero value signals dissipationless spin transport and a nonzero Drude weight. Although challenging for many-body quantum systems, the Drude weight can be accessed numerically¹⁶ and a lower bound can often be obtained analytically^{17,18}. Diffusive behavior, one of the other universal regimes, is typically recovered for systems with zero Drude weight, which implies eventual relaxation of spin currents and finite transport coefficients. Unexpectedly, an intermediate scenario was recently unveiled^{5,19}: a zero Drude weight but a slowly

decaying (typically algebraically with time) spin current-current correlation, giving rise to superdiffusive dynamics with $z = 3/2$. The intermediate scenario was found numerically⁵ by calculation of the full scaling function to belong to the Kardar-Parisi-Zhang (KPZ) universality class in 1+1 dimension, reproduced in the Supplementary Material; a theoretical scenario for how KPZ dynamics emerges in the Heisenberg chain has been proposed²⁰.

This universality class originates from the classical non-linear stochastic partial differential equation of the same name⁴, initially introduced to describe the evolution in time of the profile of a growing interface. Generally speaking, a system is considered to be in the KPZ universality class if its long-time dynamics shows the same scaling laws as appear in the KPZ equation itself. Besides interface growth²¹, such scaling has been found in disordered conductors²², quantum fluids²³, quantum circuits²⁴, traffic flow²⁵, and was recently predicted also to appear in the high-temperature dynamics of some one-dimensional integrable quantum magnets^{5,19,14,15}, although its exact microscopic origin is still under active research in this case.

Here, using neutron scattering experiments on KCuF_3 , which realizes a nearly ideal quantum Heisenberg spin-half chain, we report on the observation of KPZ dynamics at various temperatures in the range $T = 75$ K to $T = 300$ K. Combining experimental measurements with extensive numerical simulations based on a microscopic description of the system, we identify a characteristic power-law $\propto Q^{-3/2}$ behavior in the neutron scattering spectrum, in agreement with the KPZ universality class predictions¹⁸.

Searching for Kardar-Parisi-Zhang hydrodynamics

KCuF_3 has long been studied as a model of 1D Heisenberg antiferromagnetism with $S = 1/2$ spins borne by Cu^{2+} ions^{27,28,29}. Due to the Cu^{2+} $d_{x^2-y^2}$ orbital order (Fig. 1a), the magnetic exchange interaction is limited to nearest-neighbor spins and is spatially anisotropic. It is dominant along the c axis ($J_c = 33.5$ meV) while the interchain coupling is much weaker ($J_{a,b} = -1.6$ meV), leading to effective one-dimensional c axis spin-half chains. Although the system magnetically orders at $T_N = 39$ K due to the inherent presence of a finite exchange interaction $J_{a,b}$, its behavior for $T \gtrsim T_N$ is a good approximation to an ideal 1D Heisenberg

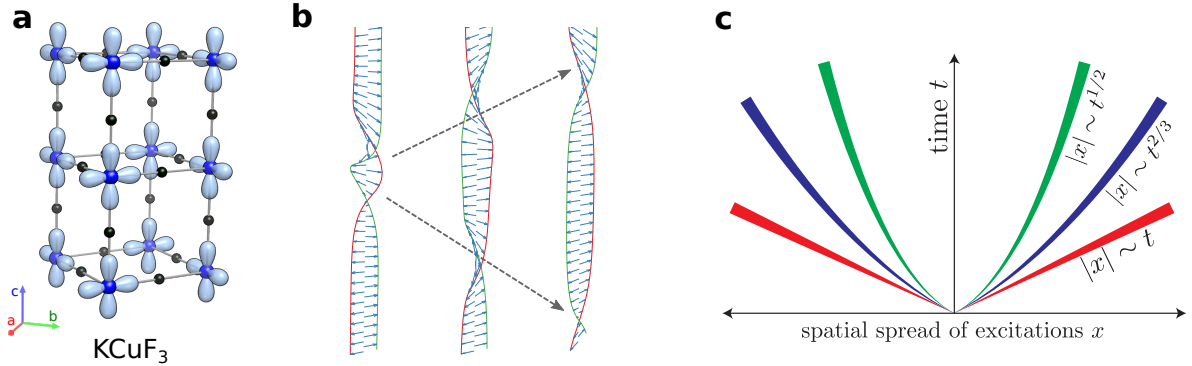


Figure 1: **a** Crystal structure of KCuF_3 , showing the orbital order of the Cu $x^2 - y^2$ orbitals. This order leads to strong magnetic exchange interactions along the c (vertical) axis and weak exchange interactions along a and b , such that the Cu^{2+} ions effectively make 1D chains. **b** Schematic illustration of spinon excitations in a 1D Heisenberg antiferromagnet (based on Ref. ⁷). **c** Schematic illustration of three possible length-time scaling behaviors $|x| \sim t^{1/z}$ observed at high temperature in 1D quantum magnets, classified by the dynamical exponent z : $z = 2$ corresponds to diffusion (green curve), $z = 3/2$ to superdiffusive (blue curve) and $z = 1$ to ballistic dynamics.

antiferromagnet which can be modeled by the following Hamiltonian,

$$\hat{\mathcal{H}} = J_c \sum_n \hat{\mathbf{S}}_n \cdot \hat{\mathbf{S}}_{n+1}, \quad (1)$$

with $\hat{\mathbf{S}}_n$ the spin-1/2 operator on the site index n . At equilibrium, the spin dynamics can be characterized through the correlation function between two spatially separated spins at different moments in time, and whose Fourier transform to momentum and frequency spaces is the dynamical spin structure factor $\mathcal{S}(Q, \omega)$. This quantity is directly proportional to the measured inelastic neutron scattering intensity, and can be computed numerically for the model (1) using matrix product state (MPS) techniques (See Methods), allowing for a direct comparison between theory and experiments. Especially, the universal dynamical exponent z is expected to manifest itself in the form ¹⁸,

$$\mathcal{S}(Q, \omega) \sim Q^{-z}, \quad (2)$$

in the limit of small momentum Q and fixed small energy $\hbar\omega$, with KPZ behavior identified by $z = 3/2$.

To search for effects of KPZ behavior, we measured the inelastic KCuF_3 neutron spectrum

using the SEQUOIA spectrometer at the Spallation Neutron Source at Oak Ridge National Laboratory. Our sample was a 6.86 g KCuF_3 single crystal mounted with the c axis perpendicular to the incident beam. To probe hydrodynamic signatures, we focused on the low-energy part of the spectrum. We measured with an incident energy $E_i = 8$ meV, which gives access to the very bottom of the spectrum (the total KCuF_3 bandwidth is 105 meV²⁹), as shown in Fig. 2, with a resolution full width at half maximum 0.25 meV. Elastic incoherent scattering prevents us from isolating the magnetic scattering at $\hbar\omega \rightarrow 0$, so we take the $0.7 < \hbar\omega < 2$ meV scattering to be an approximation to the $\hbar\omega \rightarrow 0$ spectrum. To evaluate the robustness of this approximation, we consider three different energy ranges with $\hbar\omega > 0.7$ meV (this is empirically where elastic incoherent scattering background is negligible), as indicated in Fig. 2. They all lead to similar results (see Supplementary Information for details).

Results

A quantitative test to distinguish different kinds of hydrodynamics is the scattering intensity behavior at small energy versus Q near the ferromagnetic wavevector $Q = 0$, see Eq. (2). We first compare these data to MPS simulations integrated over the same energy range in Fig. 3. There is a very good agreement between the two, with deviations only appearing at the low temperature (mainly 75 K). This discrepancy is attributed to the inherent interchain couplings (not present in the pure 1D model), and which lead to antiferromagnetic ordering at $T_N = 39$ K. In fact, it was theoretically shown³⁰ that for dynamical quantities, the 1D temperature crossover in quasi-one-dimensional systems such as KCuF_3 is spoiled for $T \lesssim 3T_N$. Furthermore the spectral intensity at low-energy in a strictly 1D system away from $Q \rightarrow \pm 0$ and $\pm\pi$ is greatly suppressed with temperature, making an accurate estimation from numerical simulations difficult; hence the non-physical oscillatory behavior in the MPS data of Fig. 3 at 100 K and 75 K for intermediate Q values. The $Q \lesssim 0.1$ experimental data deviates from power law behavior, partly because of Q resolution broadening, and partly because of the dispersion peaking away from $Q = 0$ at finite energy. Moreover, the numerical simulations do not allow us to reliably access the $Q \lesssim 0.2$ regime: this is because simulations are performed on finite-length chains (typically a hundred spins on the lattice) which introduces an artificial cutoff at low Q as

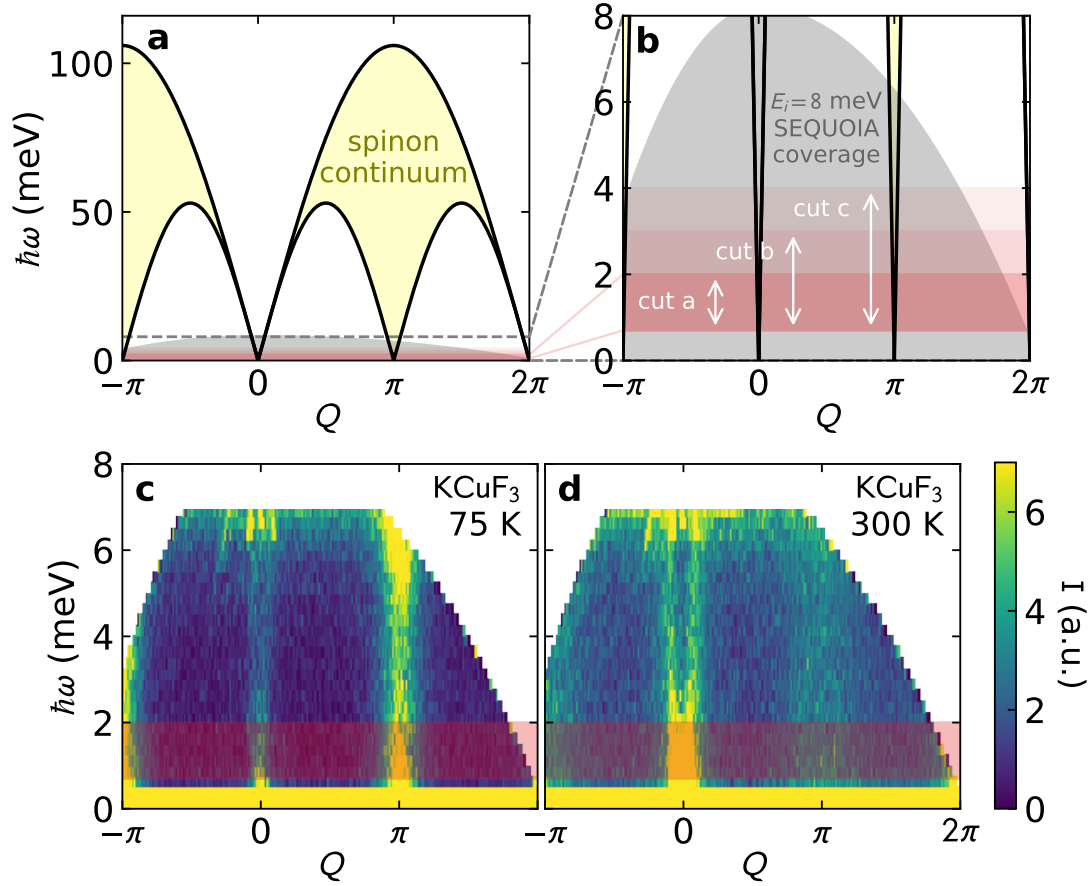


Figure 2: Measured neutron spectrum of KCuF_3 . **a** Cartoon of the KCuF_3 spinon spectrum. The gray region at the bottom shows the region measured. **b** Zoom in on the region measured in the SEQUOIA experiment, also showing three cuts (cut a, cut b, and cut c) used to approximate the $\hbar\omega \rightarrow 0$ scattering. **c** and **d** show measured spectra at 75 K and 100 K, respectively. Cut a is indicated by the horizontal red bar. It is not possible to directly measure the magnetic scattering at $\hbar\omega \rightarrow 0$ due to the strong elastic incoherent scattering. Therefore, we take the lowest energy cuts where magnetic scattering dominates, cut a, as shown in Fig. 3.

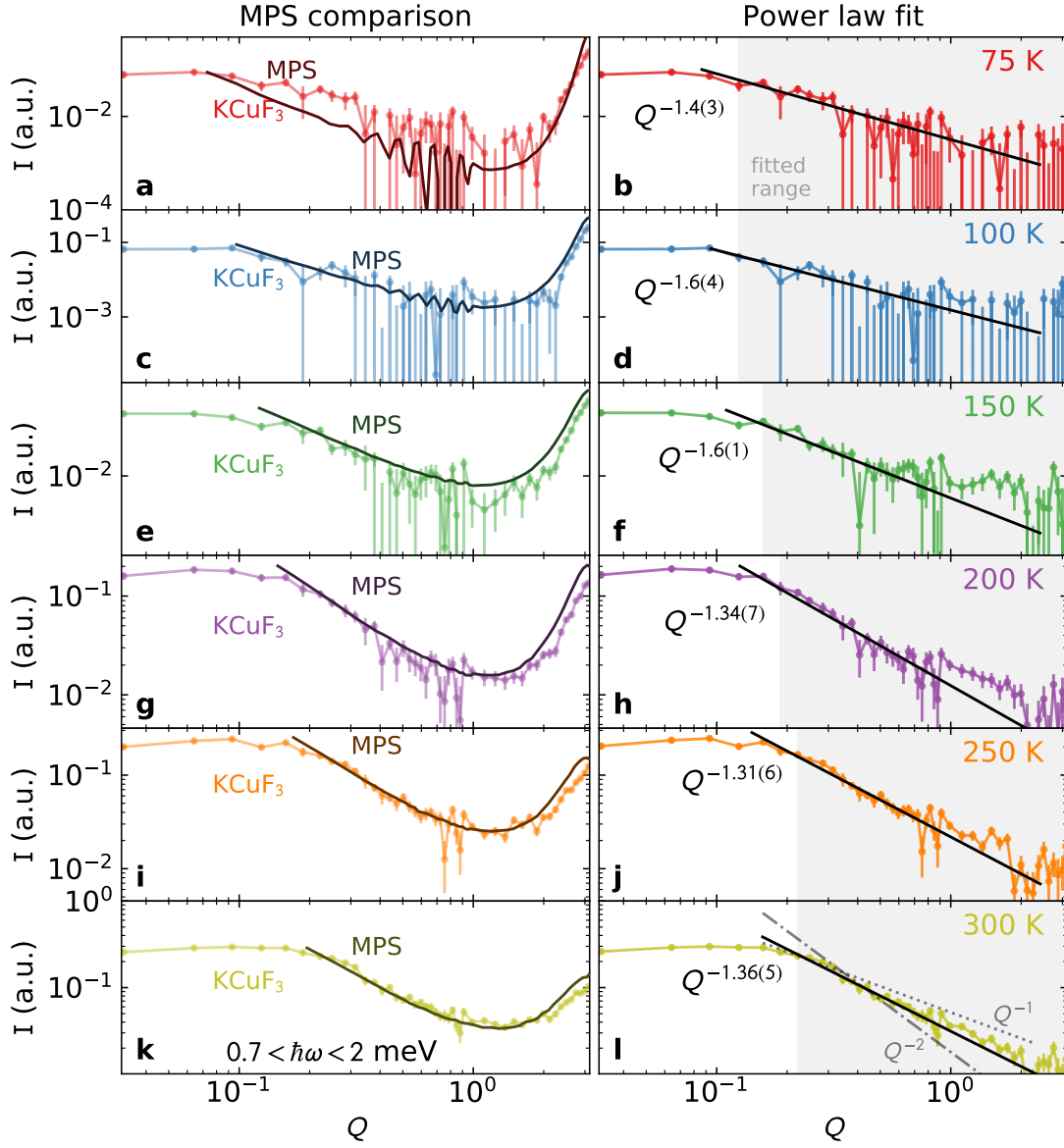


Figure 3: Power law behavior of KCuF_3 around $Q = 0$. The left column shows experimental data integrated over $0.7 < \hbar\omega < 2$ meV (cut a in Fig. 2) symmetrized about $Q = 0$ compared with the MPS simulations. The same multiplicative scaling factor is used for all temperatures, and the agreement is quite good above $Q \approx 0.2$, below which finite-size effects are significant for MPS (See the Supplementary Information). The right column shows the data fitted to a phenomenological power law. As a part of the fit, the $Q = \pi$ peak was also fitted to a power law and subtracted off as background. The fitted power is very close to $-3/2$ at all temperatures. Comparison to $z = 2$ and $z = 1$ exponents are given in panel l. (Note that Q is unitless $0 \rightarrow 2\pi$ as in Fig. 2.)

$\hbar\omega \rightarrow 0$ when it comes to the dynamics, as compared to a system in the thermodynamic limit (See the Supplementary Information).

Having identified the temperature window where 1D physics take place, we consider in Fig. 3 the same experimental data from which a phenomenological $Q = \pi$ power-law background is subtracted. This highlights the $Q \rightarrow 0$ regime, where power-law fits of the form $\propto Q^{-z}$ give an exponent close to $z \approx 1.5$ at all temperatures. Note that fits without this phenomenological background yield results which are equivalent to within uncertainty, and are shown in the Supplementary Information. The fitted exponent, while not $Q^{-3/2}$ exactly, is clearly closer to that value than to either Q^{-1} (ballistic) or Q^{-2} (diffusive). See Fig. 3I for a comparison between the three cases. At $T = 300$ K, the discrepancy between the expected $Q^{-3/2}$ behavior and the experimental fit $Q^{-1.36(5)}$ can be explained by the fact that we are not measuring at exactly $\hbar\omega = 0$. Indeed, we show in the Supplementary Information that finite-energy fits to the MPS data also yield powers less than $z = 3/2$, while fits in the low-energy limit $\hbar\omega \rightarrow 0$ — more amenable numerically than experimentally — precisely deliver the correct exponent. This effect is experimentally unavoidable, and we thus consider our experimental results, in conjunction with a comprehensive numerical study observing the same trend, to be clearly more consistent with the KPZ universality class behavior than either of the conventional possibilities, ballistic or diffusive behavior. A different experiment (e.g., with polarized neutrons or a spin-echo spectrometer) may be able to measure the magnetic scattering at the elastic line.

In measuring the low-energy KCuF_3 neutron spectrum, we also noticed a previously unreported feature, shown in Fig. 4: as temperature increases, the dispersion around $Q = 0$ softens. In other words, the split between $\pm Q$ modes increases, showing a decreased dispersion velocity: $190(20)$ $\text{meV}\cdot\text{\AA}$ at 100 K (within uncertainty of the theoretical $T = 0$ velocity 207 $\text{meV}\cdot\text{\AA}$) to $84(9)$ $\text{meV}\cdot\text{\AA}$ at 300 K. This feature is also captured by the MPS simulations, as shown in Fig. 4g-l. This temperature dependent mode softening has not been noticed before, and indicates that the excitation velocity is renormalized by interaction with other quasiparticles; a well-studied example of how low-energy quasiparticle velocities can be renormalized by interactions is the Dirac point in graphene.³¹ We leave it an open question how this mode softening may relate to KPZ physics.

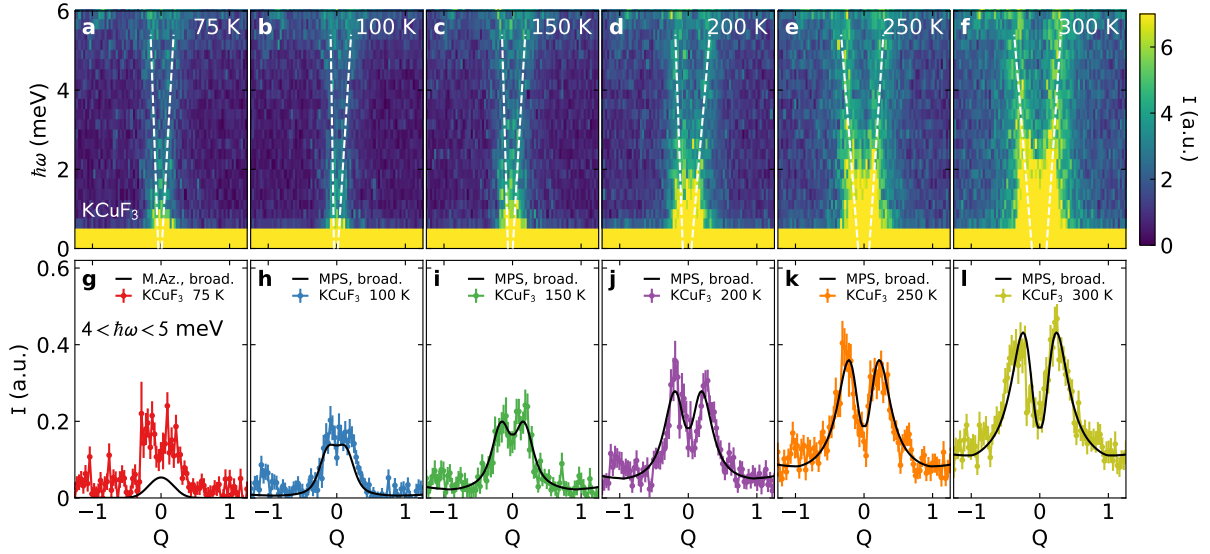


Figure 4: Temperature evolution of the KCuF_3 neutron spectra around $Q = 0$. (Q units are defined in Fig. 2.) Panels **a - f** show colormap plots of the spectra (white dashed lines are fitted linear dispersions), and panels **g - l** show constant energy cuts of $4 < \hbar\omega < 5$ meV. The lower panel also includes theoretical curves for comparison. (Panel **g** shows the zero-temperature Müller ansatz³² scaled to match the $Q = \pi$ intensities, the rest show the MPS calculations. Resolution broadening has been applied to the theoretical curves, see Supplementary Information for details.) These show the spinon modes splitting more as temperature increases, indicating a significant mode softening.

Conclusion

Our results experimentally show the presence of KPZ-like physics in KCuF_3 , characterized by the dynamical exponent $z = 3/2$. The observation of this superdiffusive scaling in the high-temperature dynamics of a Heisenberg chain demonstrates that inelastic neutron scattering can complement transport as a probe of quantum coherent collective phenomena, even when those phenomena do not have an interpretation in terms of a small number of quasiparticles. In some situations transport measurements are prohibitively difficult, like in many non-chiral 1D systems where even trace impurities or small defects simply interrupt macroscopic transport. In such cases, neutron scattering can probe subtle fluid properties where the dynamical exponent z around the low-energy dispersion reveals the nature of the collective quasiparticle flow.

In higher dimensions, this approach can provide insight into analogues of non-Fermi liquid scaling and other behaviors hypothesized to exist from transport, with the advantage of an increased degree of quantum coherence and a precise characterization of the model Hamiltonian. In one dimension, this work goes significantly beyond previous efforts with other techniques to see generalized hydrodynamics from integrability by revealing a different scaling regime. Theoretically, there remain important questions to be understood, such as how to characterize the crossover regime when integrability is broken by interchain couplings or other residual interactions. In particular, there are spin chains where the field-theory description at low temperatures shows emergent integrability but the lattice-scale physics is not integrable, unlike for the Heisenberg model, and it should be possible to apply our combined experimental and computational approach to this category of systems as well.

Methods

Full methods are available in the Supplementary Information.

Inelastic neutron scattering

Neutron scattering measurements were performed using the SEQUOIA spectrometer^{5,34} on a single crystal mounted in the hhl scattering plane, such that the c axis is perpendicular to the

beam. Although this does not measure a full reciprocal space volume, we integrate over the h and k directions because the magnetic signal only depends upon ℓ . At $E_i = 8$ meV, we measure below the lowest phonon dispersion, so the only contributions to the signal near $Q = 0$ are magnetic scattering and elastic background. We corrected the magnetic scattering for the form factor by calculating the magnitude of Q for each pixel and dividing by the Cu^{2+} form factor, as explained in the Supplementary Information. Data were plotted using Mslice³⁵.

Numerical simulations

The numerical simulations are based on a matrix product state (MPS) approach⁷ using the ITensor library⁸ to simulate the finite-temperature spin dynamics of the 1D quantum Heisenberg spin-half chain. We use the purification method to represent the finite-temperature quantum state (mixed state) as a pure state in an enlarged Hilbert space⁹. For the dynamics, we use a method expanding the spectral function $\mathcal{S}(Q, \omega)$ in terms of Chebyshev polynomials¹².

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Author contributions

A.S. and N.E.S. contributed equally to this work. The project was conceived by J.E.M. and D.A.T.. The experiments were performed by A.S., M.B.S., S.E.N., and D.A.T.. The numerical simulations and theoretical analysis were performed by N.E.S., M.D., and J.E.M.. The neutron data analysis was performed by A.S. with input from D.A.T., G.E.G., and S.E.N.. The results were discussed and paper written by A.S., N.E.S., M.D., S.E.N., J.E.M and D.A.T..

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Supplementary Information for Detection of Kardar-Parisi-Zhang hydrodynamics in a quantum Heisenberg spin-1/2 chain

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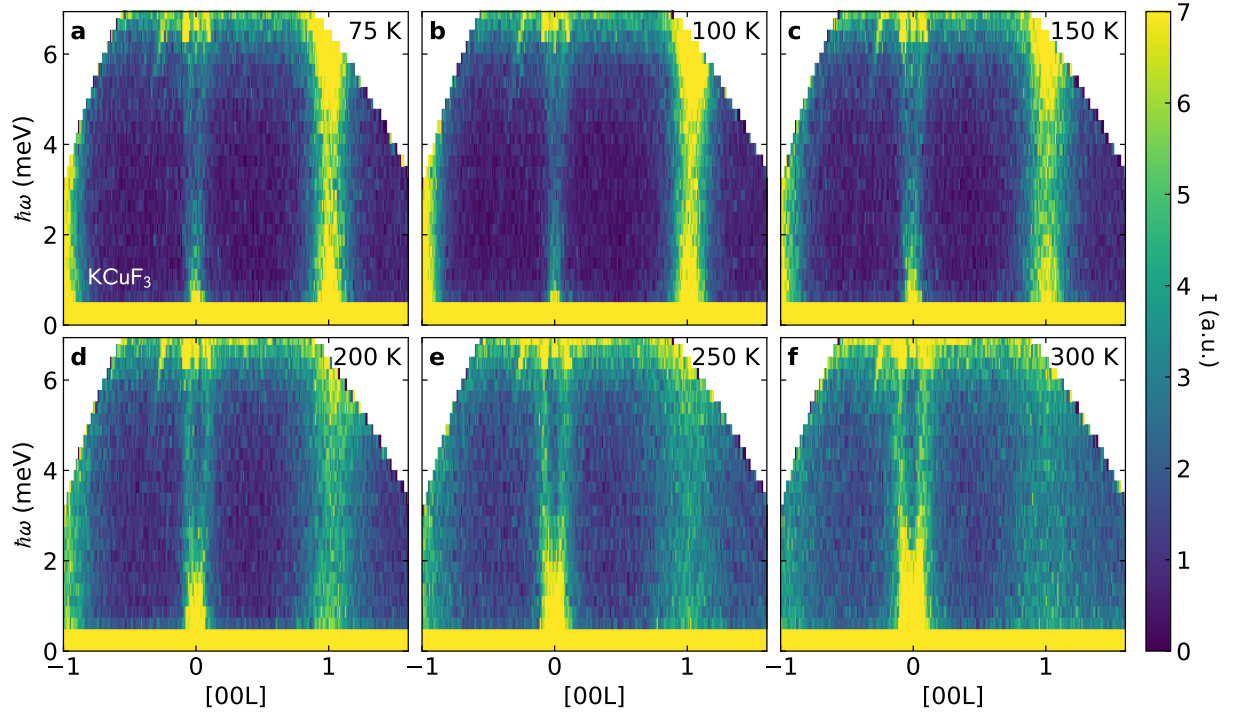


Figure S1: Temperature dependent neutron spectra of KCuF_3 .

1 Experiments

1.1 Data collection

We measured six temperatures of KCuF_3 with the c axis perpendicular to the incident neutron beam, and an incident energy of $E_i = 8.2$ meV, shown in Fig. S1. Data were integrated over all directions perpendicular to ℓ^1 . At each temperature, the spectra was measured for eight hours at nominal 1.4 Mw operation², with choppers set to 120 Hz (SEQ-100-2.0-AST chopper with 2 mm slit spacing). Data were corrected for the form factor by calculating the magnitude of Q for each pixel and dividing by the isotropic Cu^{2+} form factor (calculated following Ref.³). As shown in Fig. S2, each pixel at nonzero Q_ℓ (along the chain) or $\hbar\omega$ includes a nonzero Q_{hh} component

$$Q_{hh} = \sqrt{k_f^2 - Q_\ell^2} - k_i \quad (\text{S.1})$$

where k_f and k_i are the magnitudes of the incident and final neutron wavevectors. The final calculated form factor for this geometry is in Fig. S2b.

In order to estimate the experimental resolution of the SEQUOIA spectrometer for this

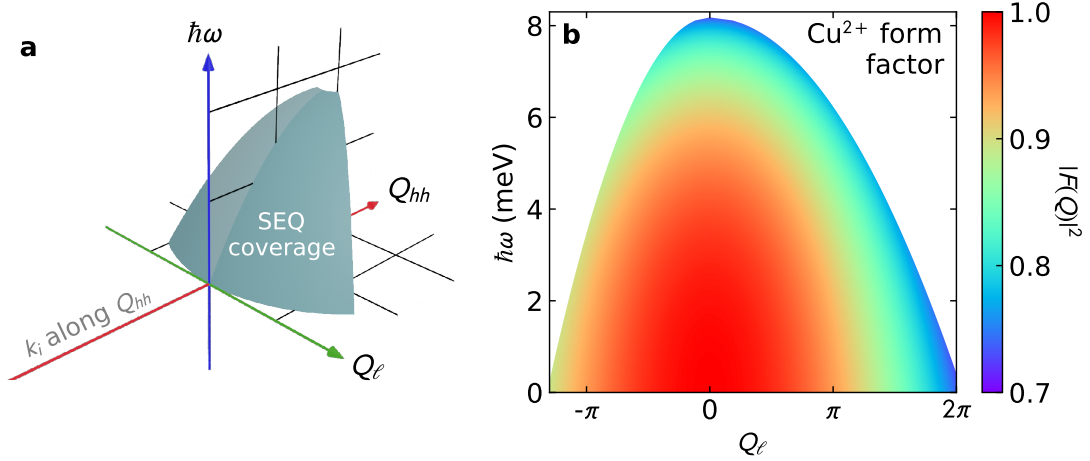


Figure S2: Form factor correction for KCuF_3 SEQUOIA experiment. Panel **a** shows the coverage in reciprocal space and energy. Because of the instrument geometry, any finite energy or momentum transfer also involves a finite momentum transfer along the $(h, h, 0)$ direction. This leads to the Cu^{2+} form factor correction shown in panel **b**, which is calculated based off the magnitude of Q .

experiment, we simulated an antiferromagnetic linear chain spin wave dispersion with a bandwidth 52.7 meV (the lower bound of the spinon continuum) using MCViNE virtual neutron experiment⁴. The Monte Carlo ray tracing simulations of SEQUOIA⁵ were run using McVine for 2×10^{10} incident neutron packets using the exact incident energy and chopper settings used during the experiment. The results are shown in Fig. S3, and indicate a Q resolution has a FWHM 0.08. Note that the finite-temperature mode softening and the power law in Q are absent from this simulated data, indicating that they are intrinsic to KCuF_3 and not resolution effects.

1.2 Data fitting

In Fig. 3 of the main text, we subtract a phenomenological power law at $Q = \pi$ in order to isolate the power law at $Q = 0$. We also chose the lowest energy window (“cut a” in Fig. 2 of the main text) to approximate the $\hbar\omega = 0$ scattering. If we do not do this, but instead define a constant background based off the lowest temperature data, the fitted exponents are shown in Fig. S4. We also show the results of fits to cuts b and c. Forgoing the phenomenological background means that the power law is visible over a narrower range in Q , but in nearly every case the fitted powers agree to within uncertainty. Meanwhile, an increase in the energy window

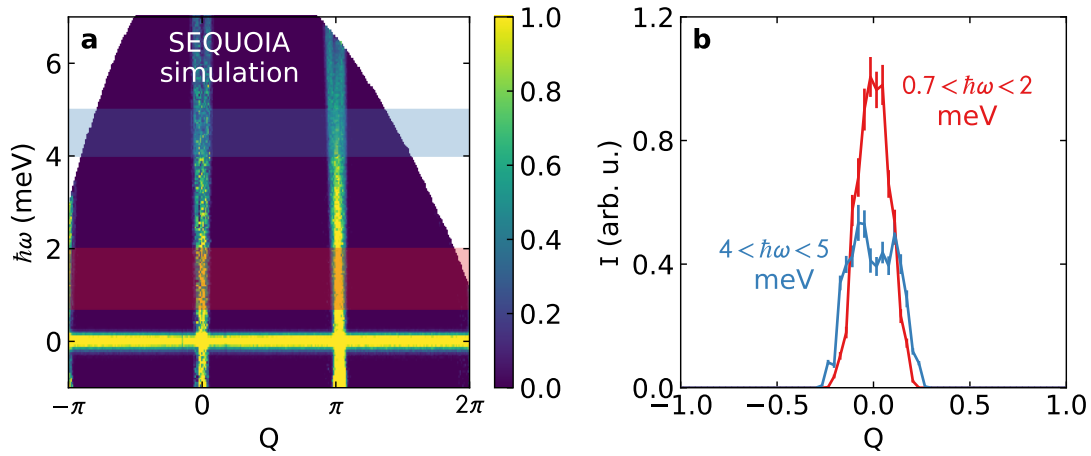


Figure S3: Simulated neutron spectrum of a 1D antiferromagnetic spin chain in the SEQUOIA experiment, in order to show resolution width. The inelastic spectra is assumed to be the semi-classical linear spin wave theory dispersion, and the Q resolution has a FWHM 0.08. Panel **a** shows the spectrum, and panel **b** shows constant-energy cuts. At 4.5 meV, the two modes of the dispersion are just barely distinguishable.

yields slightly different fitted powers, generally decreasing as the window increases. Note that all fits were performed with the scipy least squares routine⁶.

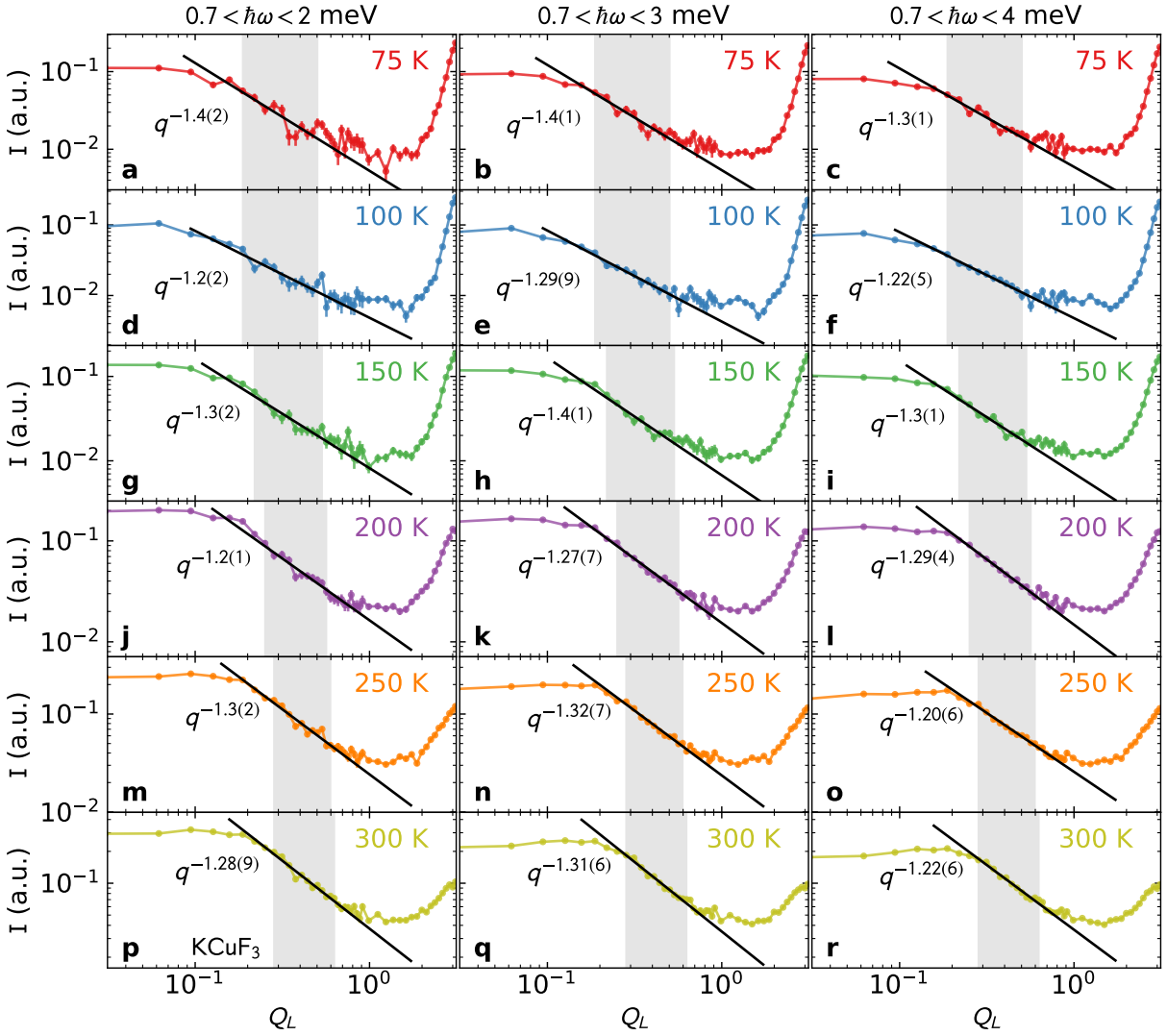


Figure S4: Power law fits for low- Q KCuF_3 scattering for three different energy windows: $0.7 < \hbar\omega < 2$ meV (cut a, left column), $0.7 < \hbar\omega < 3$ meV (cut b, middle column), and $0.7 < \hbar\omega < 4$ meV (cut c, right column). No background has been subtracted, and the window where data was fitted is indicated in gray. Note that these powers, in most cases, agree to within uncertainty with those in Fig. 4 in the main text.

2 Numerical simulations

2.1 Method

2.1.1 Finite temperature

The numerical simulations are based on a matrix product state (MPS) method⁷ using the ITensor library⁸ to simulate the finite-temperature spin dynamics of the 1D quantum Heisenberg spin-half chain, see Eq. (1) of the main text (the antiferromagnetic exchange coupling is $J_c = 33.5$ meV). We use the purification method to represent the finite-temperature quantum state $\hat{\rho}_T$ (mixed state) as a pure state $|\psi_T\rangle$ in an enlarged Hilbert space⁹. In practice, this enlarged Hilbert space corresponds to doubling the system size (from L to $2L$ degrees of freedom) with half physical “P” and half auxiliary “Q” degrees of freedom and the property that $\hat{\rho}_T = \text{tr}_Q |\psi_T\rangle\langle\psi_T|$, where tr_Q is the partial trace over the auxiliary spins (also called ancilla). In this approach, the thermal expectation value of an observable \hat{O} at temperature T is expressed as,

$$\langle\hat{O}\rangle = \langle\psi_T|\hat{O}|\psi_T\rangle / \langle\psi_T|\psi_T\rangle, \quad (\text{S.2})$$

with

$$|\psi_T\rangle = \exp\left(-\frac{\hat{\mathcal{H}}}{2k_B T}\right) |\psi_\infty\rangle. \quad (\text{S.3})$$

Here, $\hat{\mathcal{H}} = \hat{\mathcal{H}}_P \otimes \hat{I}_Q$ with $\hat{\mathcal{H}}_P$ the Hamiltonian of the physical system (and which only acts on the physical spins). The state $|\psi_\infty\rangle$ has an exact MPS representation and can be constructed exactly: it is a product state of L pairs, each involving one physical and one auxiliary spin. Each pair corresponds to a maximally entangled state (typically a singlet). In the following, in order to lighten notations, we will omit the tensor product with \hat{I}_Q when referring to operators acting on physical degrees of freedom, unless specified otherwise. The Heisenberg Hamiltonian of Eq. (1) of the main text being local with only nearest neighbor interaction terms, we perform the imaginary-time evolution of Eq. (S.3) using the time-evolving block decimation algorithm¹⁰ along with a fourth order Suzuki-Trotter decomposition¹¹. Typically, we use time step $\tau = 0.1$, or a close value commensurate with the desired inverse temperature and the number of Trotter steps N_τ such that $\tau N_\tau = 1/2k_B T$.

2.1.2 Dynamics and spectral function

When the state $|\psi_T\rangle$ is obtained and normalized such that $\langle\psi_T|\psi_T\rangle = 1$ (this corresponds to setting the partition function of the system to 1), we use a method expanding the desired spectral

function $\mathcal{S}(Q, \omega)$ in terms of Chebyshev polynomials^{12,13}. The dynamics is generated by the Liouvillian operator $\hat{\mathcal{L}} = \hat{\mathcal{H}}_P \otimes \hat{I}_Q - \hat{I}_P \otimes \hat{\mathcal{H}}_Q$, i.e., the dynamical structure factor is expressed directly in frequency space as^{14,15},

$$\mathcal{S}(Q, \omega) = \langle \hat{\mathbf{S}}_{-Q} \delta(\hbar\omega - \hat{\mathcal{L}}) \cdot \hat{\mathbf{S}}_Q \rangle, \quad (\text{S.4})$$

with the momentum space spin operators defined by,

$$\hat{\mathbf{S}}_Q = \sqrt{\frac{2}{L+1}} \sum_{r=1}^L \sin(Qr) \hat{\mathbf{S}}_r, \quad (\text{S.5})$$

with L the total number of spins in the system with lattice spacing taken equal to unity; r labels the spin index on the chain. Because MPS are more efficient at simulating systems with open boundary conditions, we have used a slightly different version of the Fourier transform in the above equation compared to the usual definition¹⁶; both are equivalent in the thermodynamic limit $L \rightarrow +\infty$. Here, the allowed momentum by the finite-length geometry are $Q = k\pi/(L+1)$ with $k = 1, 2, \dots, L$. Because the Heisenberg model is isotropic with respect to the different spin components, we can make the substitution $\langle \hat{\mathbf{S}}_{-Q} \delta(\omega - \hat{\mathcal{L}}) \cdot \hat{\mathbf{S}}_Q \rangle \rightarrow 3 \langle \hat{S}_{-Q}^\alpha \delta(\omega - \hat{\mathcal{L}}) \hat{S}_Q^\alpha \rangle$ and only compute the dynamics associated with the $\alpha \in [x, y, z]$ spin component. In practice we choose $\alpha \equiv z$, and ignore the factor of 3 as the overall scale is adjusted to compare with experiments whose spectral intensity is in arbitrary units (a.u.).

To compute the spectral function numerically, we expand the delta function of Eq. (S.4) in a Chebyshev series^{12,13}. An expansion in Chebyshev polynomials, $T_n(x) = \cos[n \arccos(x)]$, is only permitted for $x \in [-1, 1]$. Thus, we rescale the spectrum of $\hat{\mathcal{L}}$ and ω by an amount W to ensure that the spectral function $\mathcal{S}(Q, \omega)$ is only non-zero in the range $\omega \in [-1 + \epsilon, 1 - \epsilon]$. Using a small $\epsilon > 0$ helps with numerical stability. The rescaled quantities are defined by,

$$\hat{\mathcal{L}}' = \hat{\mathcal{L}}/W, \quad \omega' = \omega/W. \quad (\text{S.6})$$

Then, the dynamical structure factor is written

$$\mathcal{S}(Q, \omega) = \frac{1}{W\pi\sqrt{1-\omega'^2}} \sum_{n=0}^{\infty} (2 - \delta_{n0}) \mu_n(Q) T_n(\omega'), \quad (\text{S.7})$$

with $\mu_n(Q) = \langle t_0 | t_n \rangle$ the Chebyshev moments and $|t_n\rangle = T_n(\hat{\mathcal{L}}') \hat{S}_Q^z |\psi_T\rangle$ the Chebyshev vectors. They can be calculated iteratively by

$$|t_n\rangle = 2\hat{\mathcal{L}}'|t_{n-1}\rangle - |t_{n-2}\rangle, \quad (\text{S.8})$$

with $|t_0\rangle = \hat{S}_Q^z |\psi_T\rangle$ and $|t_1\rangle = \hat{\mathcal{L}}'|t_0\rangle$ as a starting point. For comparison between experiments and numerical simulations, we use a system size of $L = 128$ spins, a bond dimension $\chi = 256$, and the order of the Chebyshev expansion is $N = 3000$. We chose $W = L/2(1 - \epsilon)$ together with $\epsilon = 0.0125$ for this work.

Gaussian broadening.— Extracting an analytic spectral function from simulations is difficult due to the discrete spectrum for any finite system. A common technique is to broaden the delta functions with a smooth distribution, such as a Gaussian. Here, this is achieved by scaling the moments, $\mu_n(Q)$, by a damping factor g_n that also smooths out the Gibbs oscillations that occur from truncating the series in Eq. (S.7) to a finite value N . In this work, we occasionally use Jackson damping¹⁷, with

$$g_n = \frac{N - n + 1}{N + 1} \cos\left(\frac{\pi n}{N + 1}\right) + \frac{1}{N + 1} \sin\left(\frac{\pi n}{N + 1}\right) \cot\left(\frac{\pi}{N + 1}\right). \quad (\text{S.9})$$

Note that g_n is a monotonic function of n that decays from 1 to 0, and in the limit $N \rightarrow +\infty$, $g_n \rightarrow 1$ for all n . Jackson damping has the effect of smoothing out the finite system spectral peaks with a Gaussian with an ω dependent width. This Gaussian broadening procedure is used in Figs. S7 and S6, as well as in Fig. 4 of the main text.

Chebyshev moments.— The Chebyshev moments, $\mu_n(Q)$, are the expansion coefficients in Eq. (S.7), and dictate the convergence of the numerical simulations. The moments have an envelope that decays exponentially¹³, and if $\mu_n(Q)$ is not “sufficiently close to zero”, then our simulations are unreliable. We fit the envelope of the moments to a decaying exponential of the form $\propto e^{-n/\xi}$, and if $\xi > 500$, we say the moments are not converged. This procedure defines a minimum Q value, Q_{\min} , and we don’t show $Q < Q_{\min}$ when comparing with experiments in the main text. We will discuss this more in Sec. 2.3.3, and show the moments in Fig. S16 and Fig. S17. Numerical error can spoil the iterative process of Eq. (S.8), and can cause μ_n to rapidly drop to near zero, and then diverge for values of $n > N^*$. When this happens, we zero out the moments for all $n > N^*$ when computing $\mathcal{S}(Q, \omega)$ in the sum of Eq.(S.7), effectively truncating the series at N^* . This occurs only at $T = 100$ K, and $T = 75$ K, and for Q values larger than the power law region. The lowest value of N^* used is 2600, and we show in Fig. S5 the moments for this case.

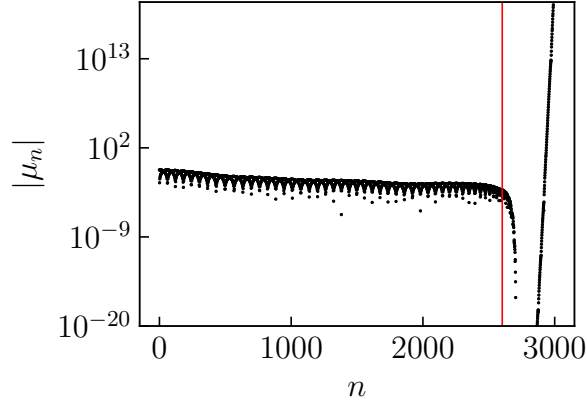


Figure S5: The absolute value of the Chebyshev moments, $|\mu_n|$, as a function of n for the case of $T = 75$ K, and $Q = 1.680$. The vertical line in red illustrates N^* , the value of n for which we terminate the series in Eq. (S.7).

2.2 Additional numerical data

2.2.1 Spectral function

The full ω dependence is computed, and shown in Fig. S6 and Fig. S7 for various temperature values. Jackson damping is used in these figures. As temperature is lowered, agreement with the low-energy spectrum is observed. The region near $(Q, \omega) = (0, 0)$ relevant for the neutron scattering experiments is also shown. We see strong intensity at $(Q, \omega) = (0, 0)$ with bifurcating dispersion lines emerging as we move away from this point. We note that the bifurcation point appears to occur at finite ω from the numerical data, but this is a finite size effect, and does not seem to occur in the experimental data. To verify this, we show the frequency bifurcation, ω_c as a function of the Chebyshev expansion order N in Fig. S8. As $N \rightarrow +\infty$, we're approaching the thermodynamic limit, and we see this bifurcation point tends towards $\omega_c = 0$.

2.2.2 Kardar-Parisi-Zhang scaling function

The scaling function for the $1 + 1$ KPZ universality class $f(\cdot)$ is known numerically exactly¹⁸ in real space x and time t . In particular, Ref.¹⁸ provides raw data for the scaling function $f(y)$ with $y \propto xt^{-2/3}$, and following this work, we Fourier transform this data to arrive at the scaling function $\hat{f}(\cdot)$ in momentum Q and frequency ω space. The relation to the spectral function, in

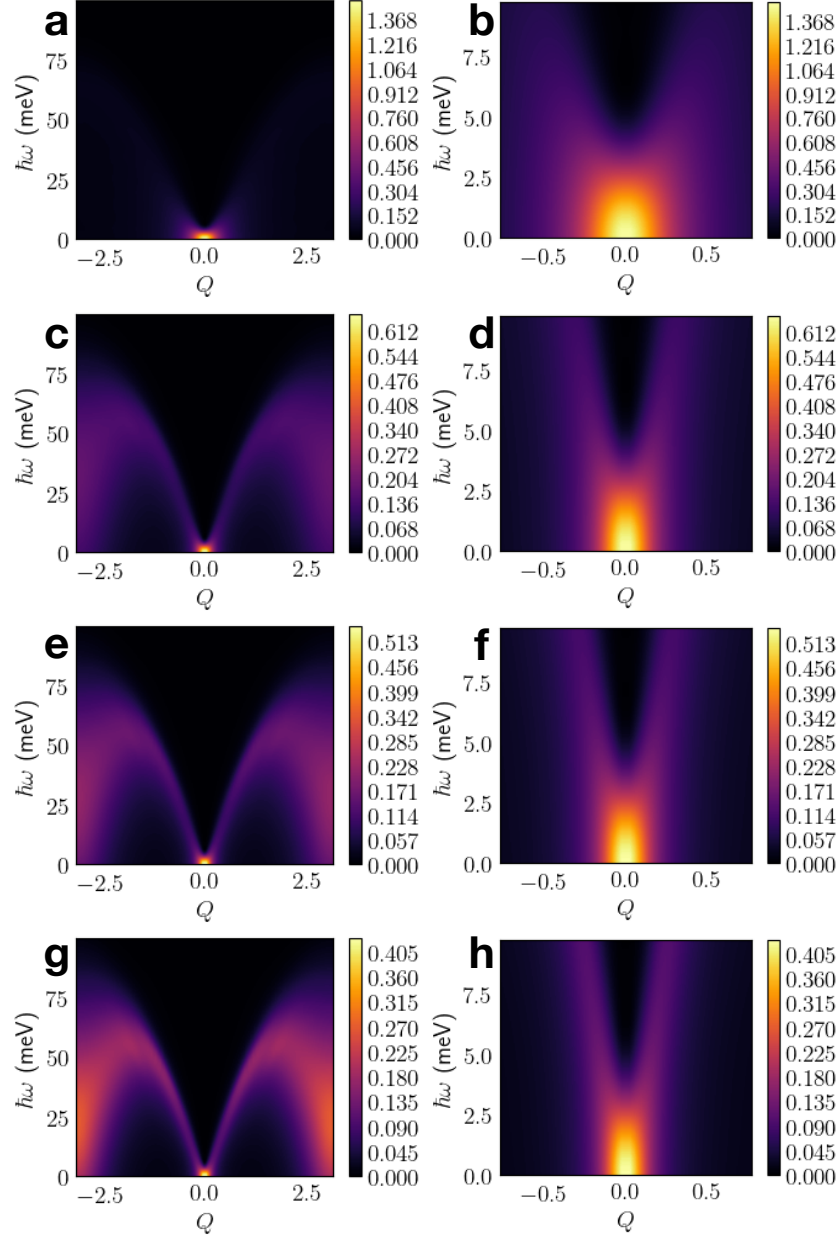


Figure S6: The spectral function, $\mathcal{S}(Q, \omega)$ at **a, b** $T = +\infty$, **c, d** 300 K, **e, f** 250 K, and **g, h** 200 K. The left column illustrates the full spectrum, and the right column zooms in around $(Q, \omega) = (0, 0)$, the region relevant for comparison with experiments.

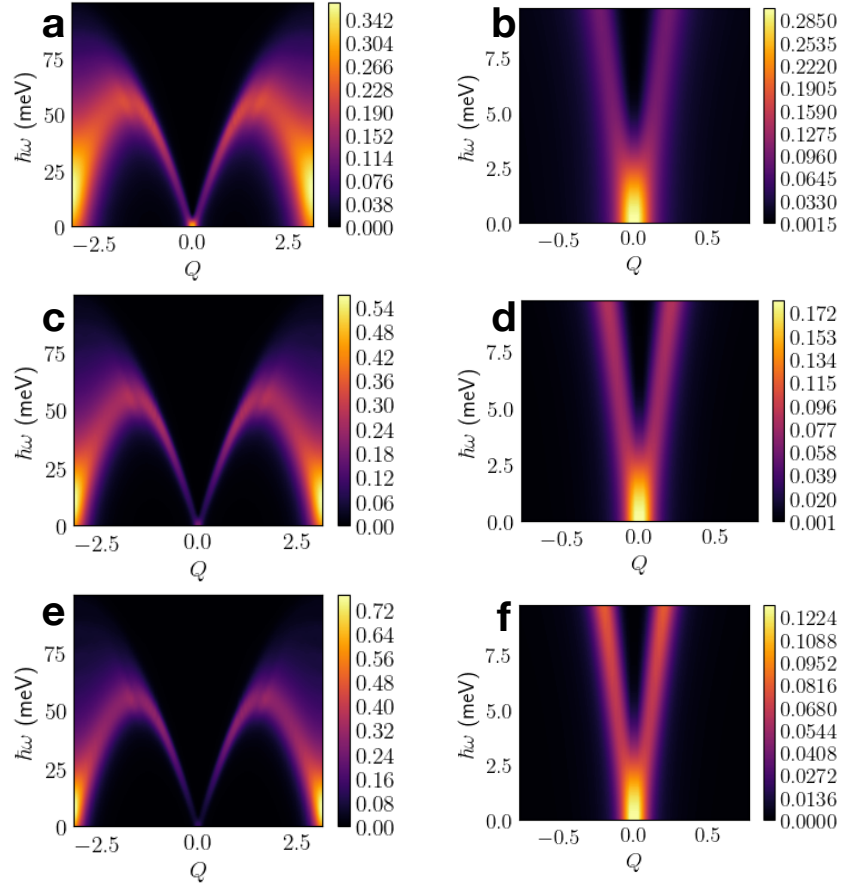


Figure S7: The spectral function, $\mathcal{S}(Q, \omega)$ at **a, b** $T = 150$ K, **c, d** 100 K, and **e, f** 75 K. The left column illustrates the full spectrum, and the right column zooms in around $(Q, \omega) = (0, 0)$, the region relevant for comparison with experiments.

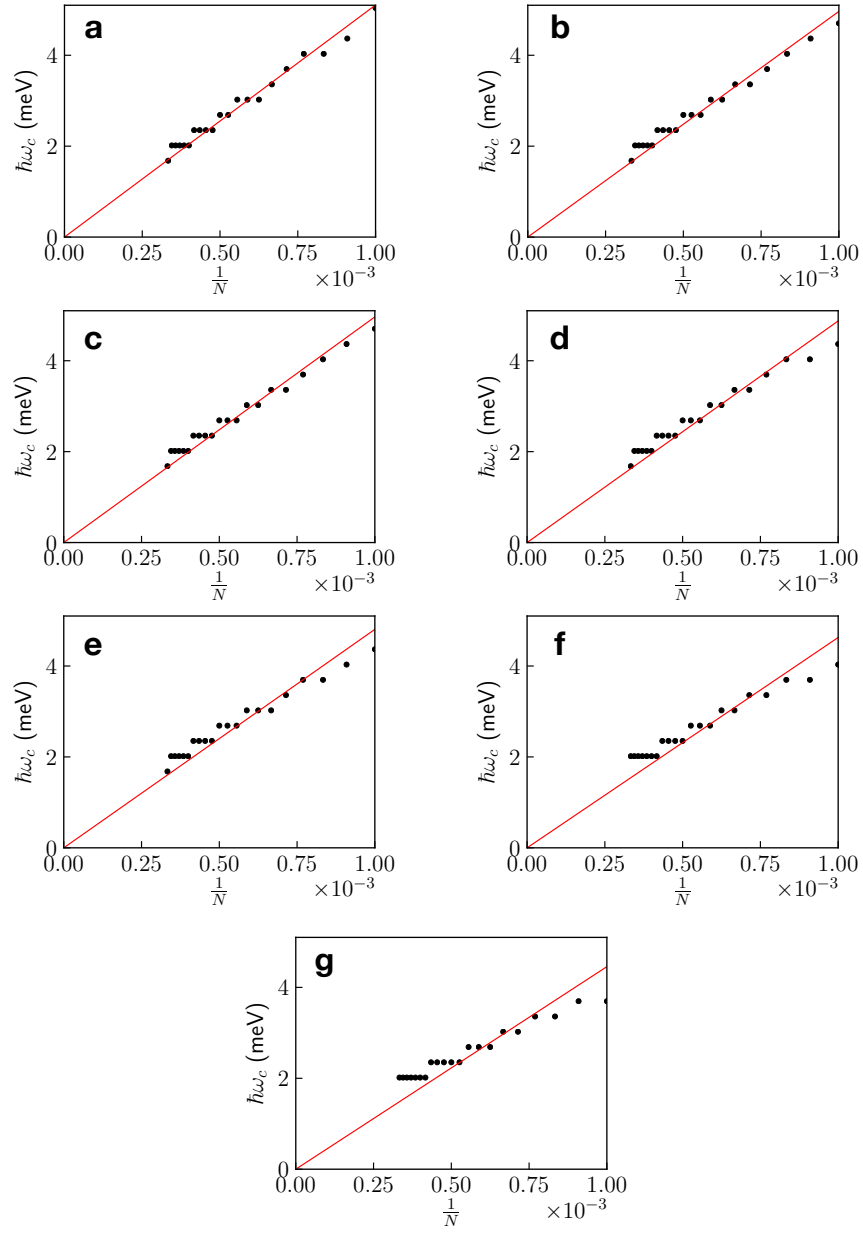


Figure S8: The value of $\hbar\omega_c$ at which the spinon mode splitting occurs at $Q = 0$ versus the inverse of the number of terms used in the sum of Eq. (S.7). The different subplots correspond to **a** $T = +\infty$, **b** 300 K, **c** 250 K, **d** 200 K, **e** 150 K, **f** 100 K, and **g** 75 K.

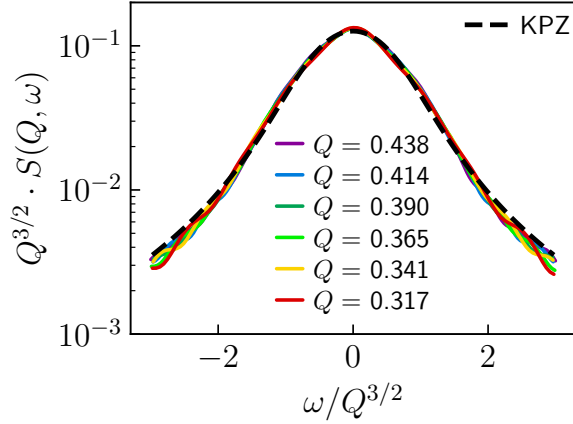


Figure S9: The KPZ scaling function plotted against the numerical data for various values of Q at $T = +\infty$.

the hydrodynamic regime is given by

$$S(Q, \omega) = c_1 Q^{-3/2} \mathring{f}(c_2 \omega Q^{-3/2}), \quad (\text{S.10})$$

where c_1 and c_2 are system-dependent constants that we use as fitting parameters to compare with the numerical simulations. We found $c_1 \simeq 0.026$ and $c_2 \simeq 0.639$ by fitting \mathring{f} to the numerical data at $T = +\infty$ and $Q = Q_{\min}$ for the optimum simulation parameters used in this work. Comparison with our numerical simulations and the scaling functions for KPZ are shown in Fig. S9.

2.2.3 Robustness of the power law fit with the energy window

We show the power law fit for different choices of the energy window at the temperatures relevant for experiments in Fig. S10, analogous to Fig. S4. In this case, the trends are more clear: at lower temperatures, the fitted power is larger, and larger energy windows suppresses the fitted power at high temperatures, and enhances the fitted power at the lowest temperatures.

These effects are shown more systematically in Fig. S11, which show deviations from the ideal KPZ $z = 3/2$ exponent as temperature becomes finite and the fitted energy increases. At infinite temperature and zero energy, the fitted exponent is almost exactly $-3/2$. At 300 K and at low energy, the fitted exponent is still quite close. However, as temperature drops the $\hbar\omega$ fitted exponent drifts toward -2 , while the finite-energy fitted value ranges from below 1 to above 2, depending on the temperature. These effects are important to take into account,

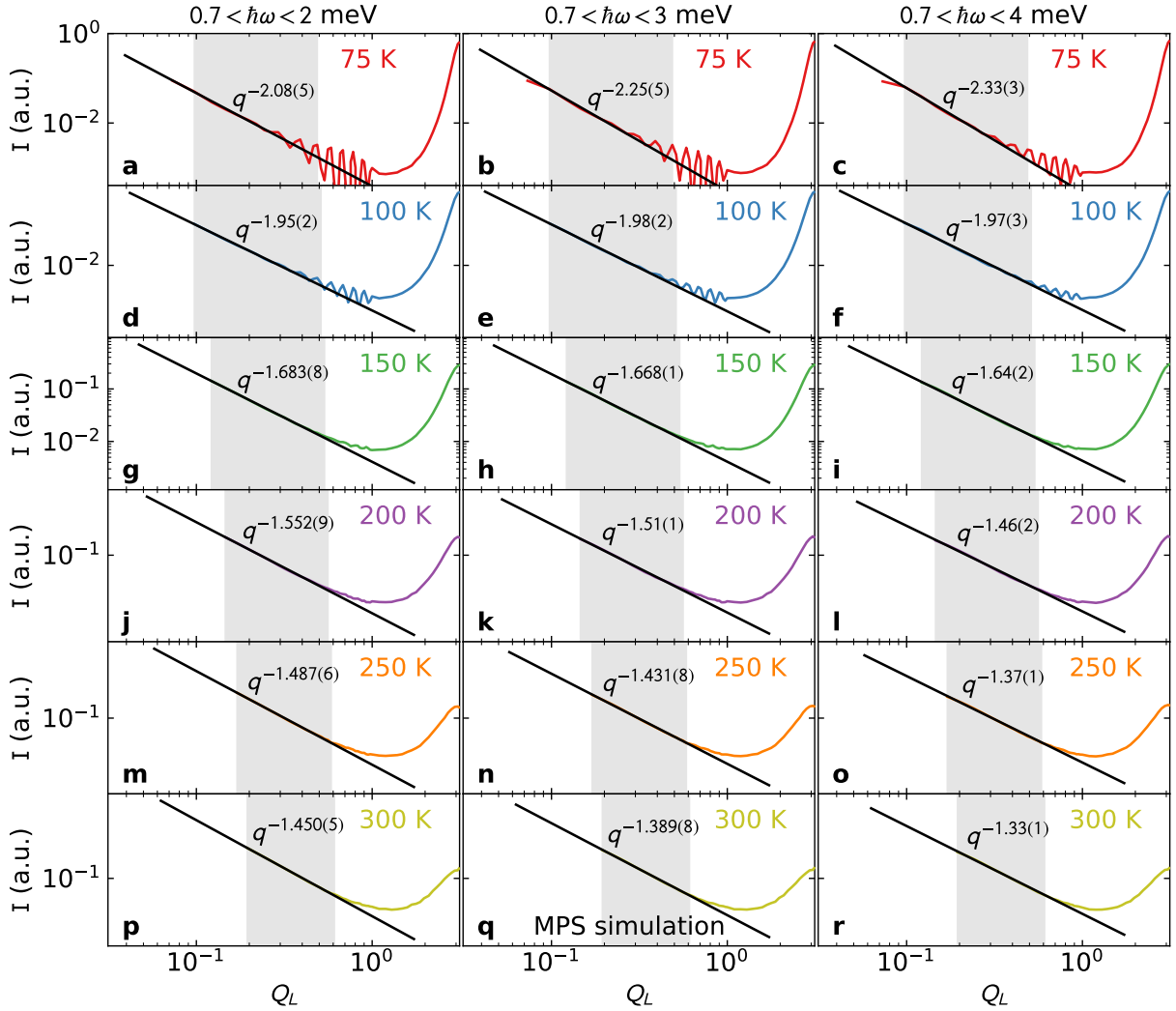


Figure S10: Power law fits to the low- Q MPS simulated spectrum for three different energy windows: $0.7 < \hbar\omega < 2$ meV (cut a, left column), $0.7 < \hbar\omega < 3$ meV (cut b, middle column), and $0.7 < \hbar\omega < 4$ meV (cut c, right column). The window where data was fitted is indicated in gray. Note that the finite energy transfer causes the exponent to deviate from the $\hbar\omega = 0$ value of $z = 3/2$: too large at low temperatures, and too high at high temperatures. Also note that the exponent magnitude is suppressed as the energy transfer window increases.

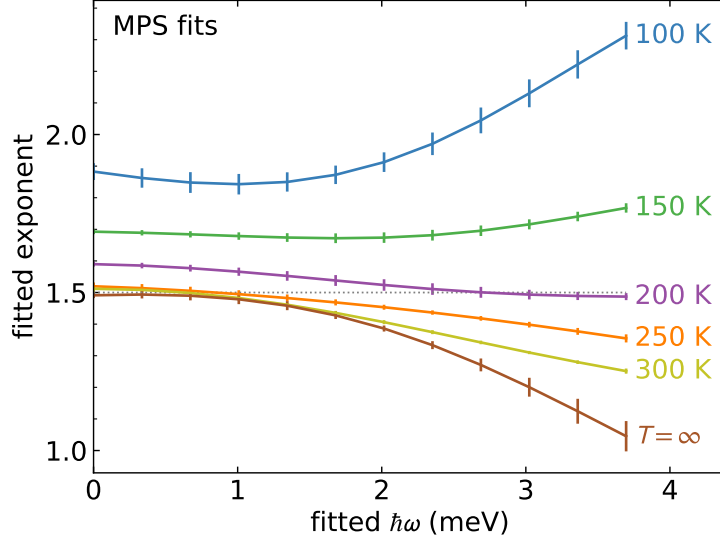


Figure S11: Deviations from $Q^{-3/2}$ KPZ behavior at finite temperature and finite energy transfer. Each curve shows the fitted exponent for the MPS simulated low- Q scattering as a function of energy. At infinite temperature and $\hbar\omega = 0$, the exponent is $-3/2$. As temperature decreases, the exponent generally increases. However, above 200 K, the KPZ behavior is still dominant.

because any real experiment will measure the exponent with a finite temperature and non-zero energy transfer.

2.3 Convergence checks with simulation parameters

2.3.1 Bond dimension of the matrix product state

For comparison with experiments, we used a bond dimension $\chi = 256$, and here we show the effects of changing χ . When examining the Q dependence for low ω , the effect of χ is insignificant as illustrated in Fig. S12, suggesting convergence, at least in the Q and ω regime of interest. In this figure, we integrate over the window $0.7 < \hbar\omega < 2$ meV, which corresponds with the energy window used in comparison with experiments. Lastly, we show the effect of χ on the scaling function in Fig. S13, where we see the oscillations decay with increasing χ , but overall good agreement with the KPZ scaling function is observed.

2.3.2 System size and finite size effects

The system size L is the simulation parameter that has the largest effect on the simulations. Determining the ω dependence in the thermodynamic limit from a finite system is inherently

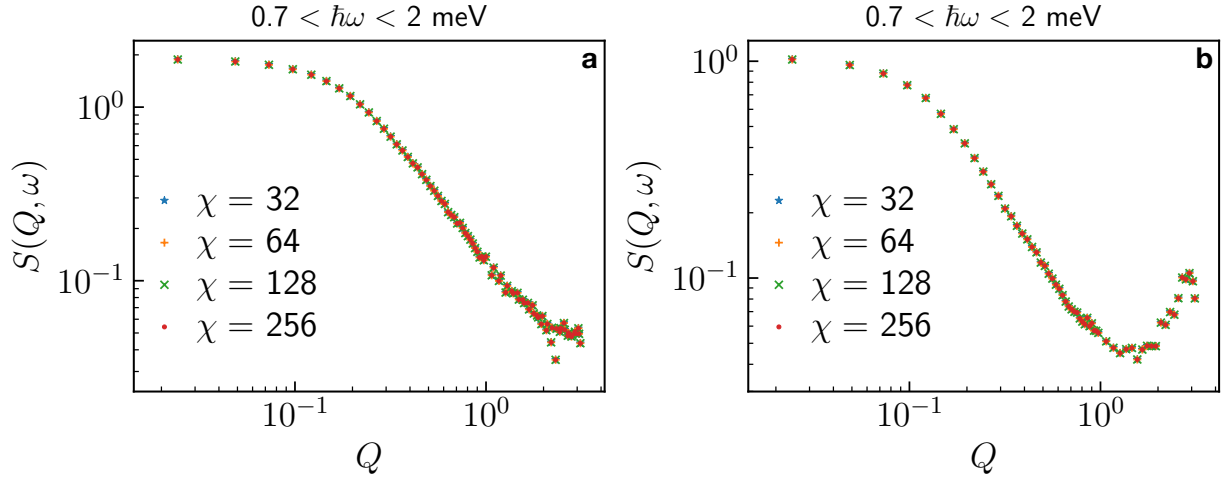


Figure S12: The Q dependence of $\mathcal{S}(Q, \omega)$ for $0.7 < \hbar\omega < 2$ meV for different choice of the MPS bond dimension χ . Subplot a corresponds to $T = +\infty$, and b corresponds to $T = J_c \simeq 390$ K.

difficult as for finite L , $\mathcal{S}(Q, \omega)$ is a sum of delta functions, yet in the thermodynamic limit, $\mathcal{S}(Q, \omega)$ is an analytic function of ω . We show the Q dependence $\mathcal{S}(Q, \omega)$, integrated over ω in the window $0.7 < \hbar\omega < 2$ meV, in Fig. S14. We see that the major difference between different system sizes is an overall multiplicative factor. Since the experimental data is collected with arbitrary units (a.u.), such an overall multiplicative factor is irrelevant when making a comparison between experiments and numerical simulations. Moreover, the curves at different system size exhibit nearly identical power law behavior for intermediate values of Q and thus the dynamical exponent z is quite robust to system size, and this is where the KPZ signature lies. We examine how the scaling function changes with system size in Fig. S15. We notice the major difference between system size occurs for low $\omega/Q^{3/2}$, which is due to an inability to access small frequencies for small system sizes (the system size provides an artificial cutoff).

2.3.3 Chebyshev expansion order

In the main text, we include $N = 3000$ terms in Eq. (S.7) for comparison with experiments. N/W is a pseudo-time parameter in this method, and larger values produce greater ω resolution. The Q dependence at small ω is illustrated in Fig. S17 and Fig. S16, where we see the simulations are well converged in the region where comparison with experiments is made. Convergence is not found for very small $Q < Q_{\min}$, which is understood when making the analogy between N and time t . Exactly at $Q = 0$, then $\hat{S}_{Q=0}^z$ is a conserved quantity, and so for very

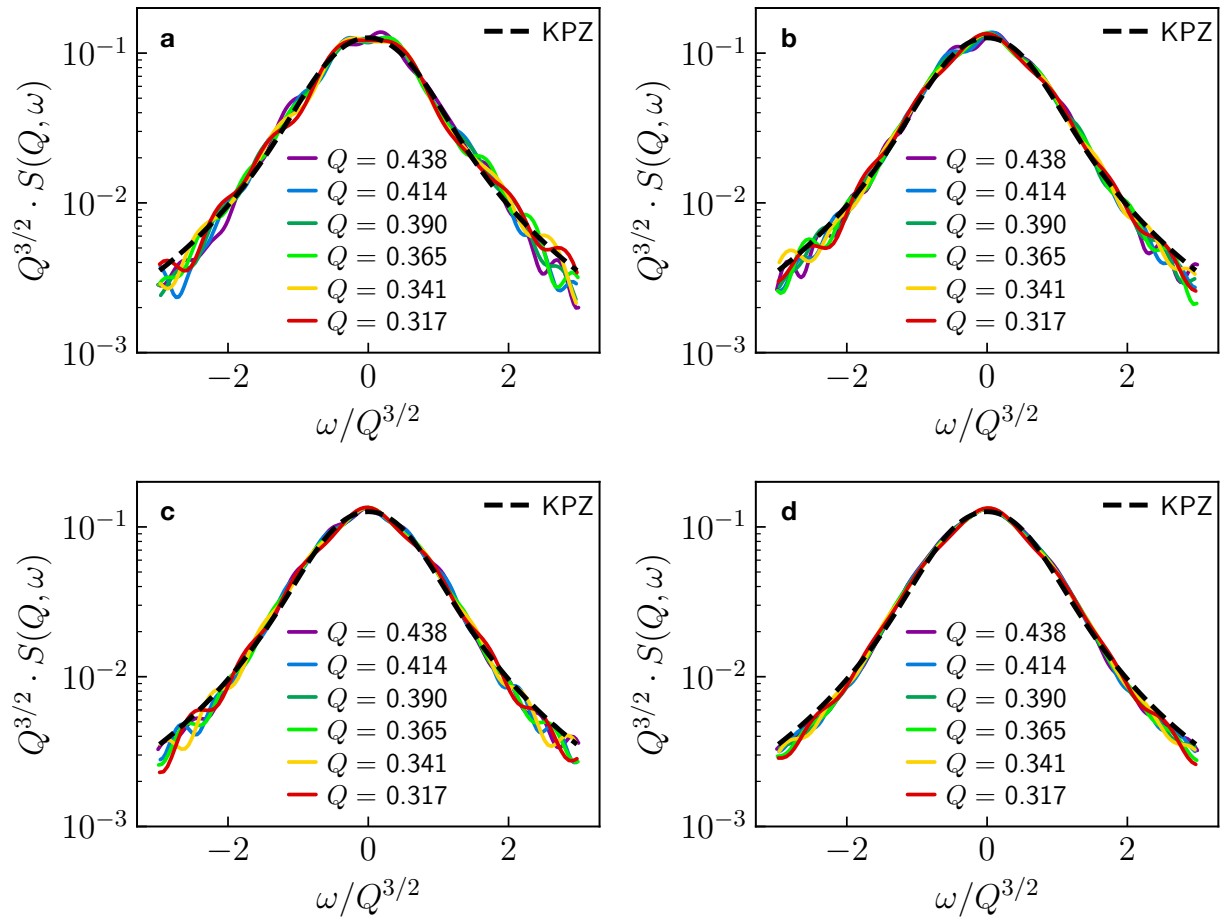


Figure S13: The KPZ scaling function plotted against the numerical data for various values of Q at $T = +\infty$. The different subplots correspond to **a** $\chi = 32$, **b** $\chi = 64$, **c** $\chi = 128$, and **d** $\chi = 256$.

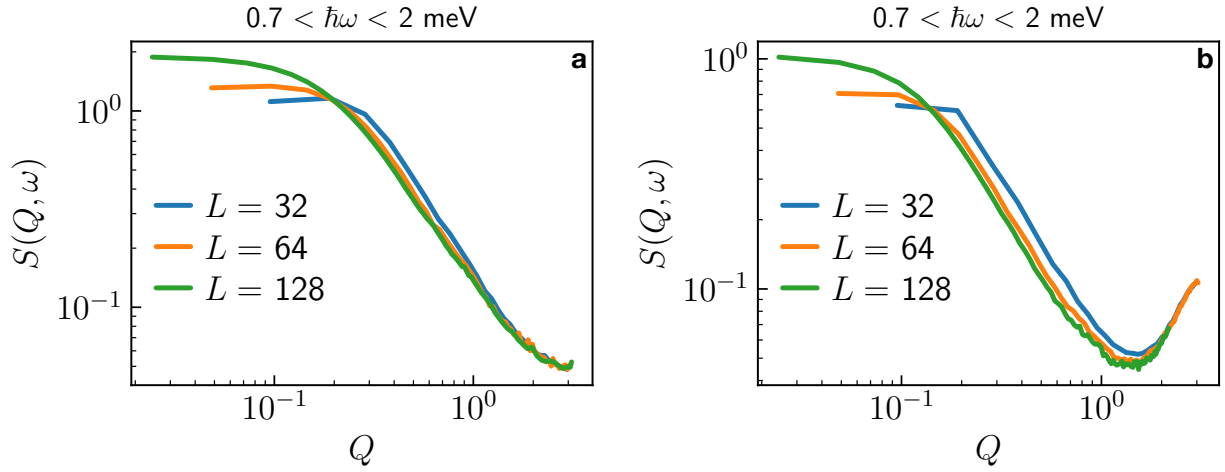


Figure S14: The Q dependence of $\mathcal{S}(Q, \omega)$ for $0.7\hbar\omega < 2$ meV with different choice of the system size L . Subplot **a** corresponds to $T = +\infty$, and **b** corresponds to $T = J_c \simeq 390$ K.

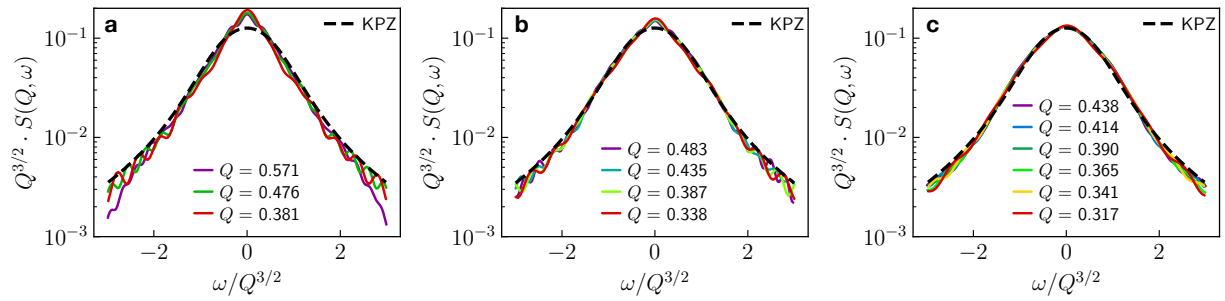


Figure S15: The KPZ scaling function plotted against the numerical data for various values of Q at $T = +\infty$. The different subplots correspond to **a** $L = 32$, **b** $L = 64$, and **c** $L = 128$.

low Q , very large t (and equivalently very large N) is needed to accurately resolve the ω dependence. Nevertheless, the capturing of the KPZ scaling function at $T = +\infty$ is quite robust to N , and is illustrated in Fig. S18.

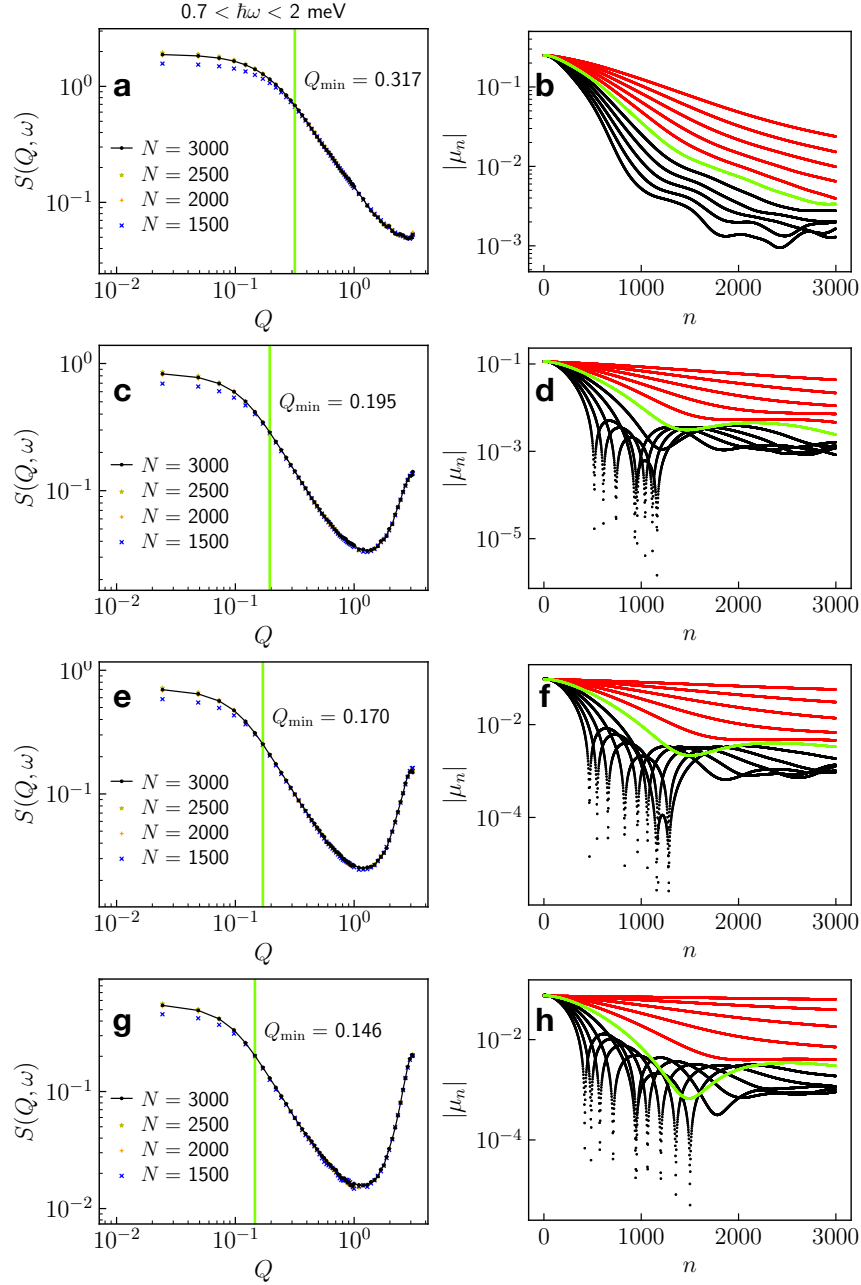


Figure S16: The left column shows the Q dependence of $S(Q, \omega)$ for $0.7\hbar\omega < 2 \text{ meV}$ with different choice of the number of terms, N , used in Eq. (S.7). The vertical green line corresponds to the value $Q = Q_{\min}$. The right column is the absolute value of the Chebyshev moments appearing in Eq. (S.7) as a function of the number of iterations n for several values of Q . Only the moments with even values of n are shown for clarity. The green curve in each plot depicts $Q = Q_{\min}$, the red curves are $Q < Q_{\min}$, and the black curves are $Q > Q_{\min}$. The temperatures shown are **a, b** $T = +\infty$, **c, d** 300 K, **e, f** 250 K, and **g, h** 200 K.

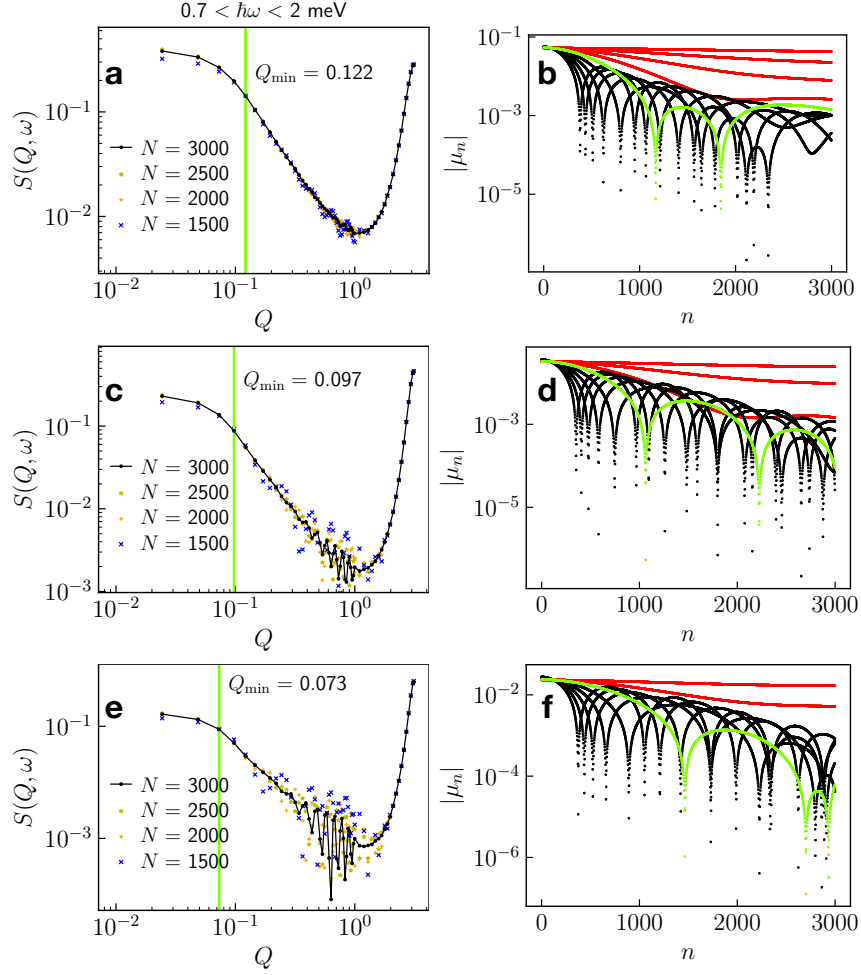


Figure S17: The left columns shows the Q dependence of $S(Q, \omega)$ for $0.7 < \hbar\omega < 2$ meV with different choice of the number of terms, N , used in Eq. (S.7). The vertical green line corresponds to the value $Q = Q_{\min}$. The right column is the absolute value of the Chebyshev moments appearing in Eq. (S.7) as a function of the number of iterations n for several values of Q . Only the moments with even values of n are shown for clarity. The green curve in each plot depicts $Q = Q_{\min}$, the red curves are $Q < Q_{\min}$, and the black curves are $Q > Q_{\min}$. The temperatures shown are **a, b** $T = 150$ K, **c, d** 100 K, and **e, f** 75 K.

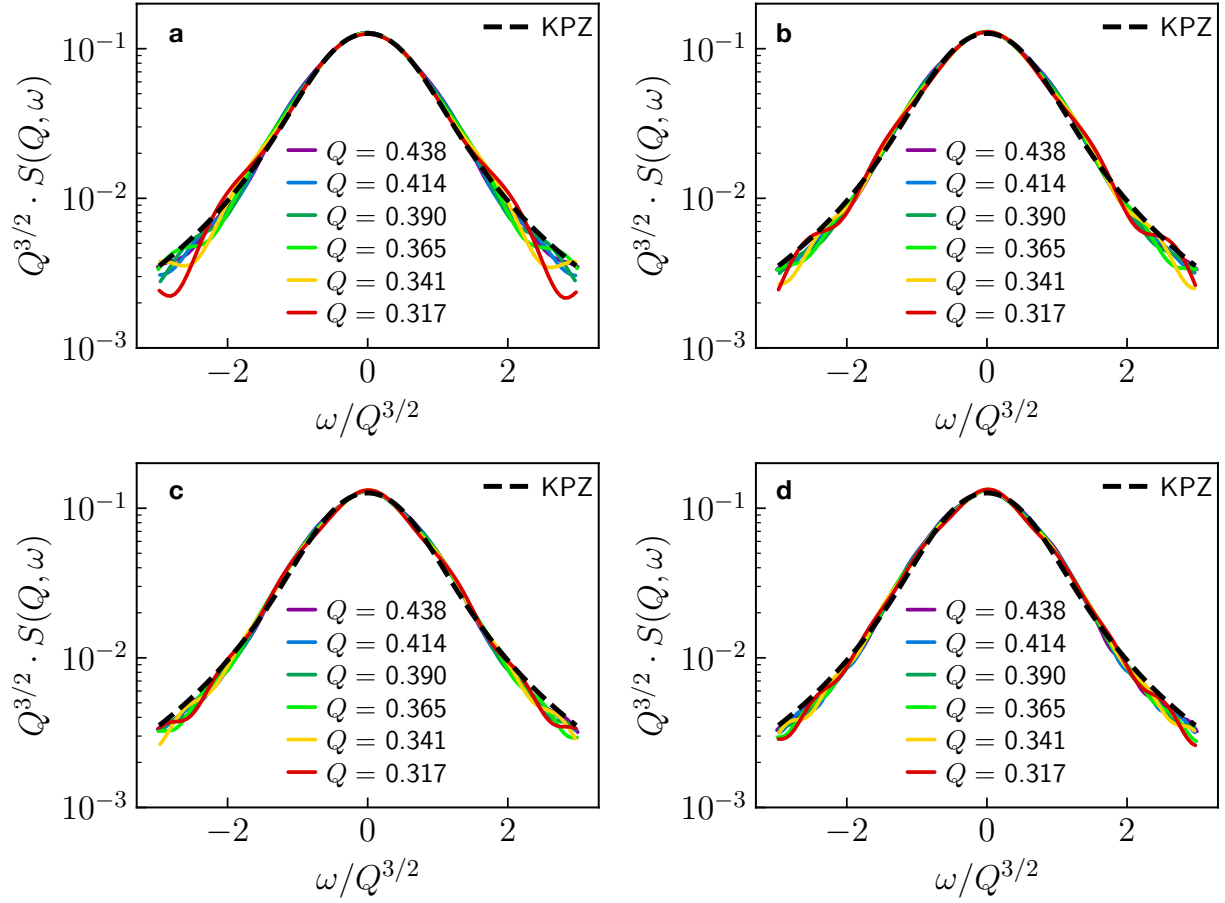


Figure S18: The KPZ scaling function plotted against the numerical data for various values of Q at $T = +\infty$. The different subplots correspond to **a** $N = 1500$, **b** $N = 2000$, **c** $N = 2500$, and **d** $N = 3000$.

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