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# Method-independent cusps for atomic orbitals in quantum Monte Carlo

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We present an approach for augmenting Gaussian atomic orbitals with correct nuclear cusps. Like the atomic orbital basis set itself, and unlike previous cusp corrections, this approach is independent of the many-body method used to prepare wave functions for quantum Monte Carlo. Once the basis set and molecular geometry are specified, the cusp-corrected atomic orbitals are uniquely specified, regardless of which density functionals, quantum chemistry methods, or subsequent variational Monte Carlo optimizations are employed. We analyze the statistical improvement offered by these cusps in a number of molecules and find them to offer similar advantages as molecular-orbital-based approaches while maintaining independence from the choice of many-body method.

# I. INTRODUCTION

The choice of functions to include in the one-electron basis set is critical in electronic structure theory. In general, a good basis set offers efficient numerical calculation, appropriate flexibility for the physics at hand, and a form that is straightforwardly transferrable between molecules. In molecular electronic structure, examples of important flexibility include the use of polarization functions to angularly resolve molecular orbital (MO) features<sup>1</sup> and the use of diffuse functions to capture the physics of anions<sup>2</sup> and Rydberg states.<sup>3</sup> To achieve efficient calculation, quantum chemistry has long relied on atom-centered Gaussian atomic orbital (AO) basis sets, which greatly simplify the evaluation of key integrals.<sup>4</sup> Thanks to their dependence on only the identities and positions of the nuclei, such Gaussian basis sets have proven readily transferrable between different molecules. However, the incorrect shape of Gaussian orbitals in the immediate vicinity of the nucleus makes them less desirable in all-electron real-space quantum Monte Carlo methods, whose accuracy and efficiency rely on the hydrogenic divergence of kinetic energy (KE) to counteract the diverging Coulomb energy at the nucleus.<sup>5</sup> Given the high accuracy offered by Monte Carlo methods, a way to correct orbital shapes at nuclei that retains the efficiency, flexibility, and transferability of the parent basis set is highly desirable.

Variational Monte Carlo (VMC), for example, has proven capable of delivering high accuracy in a number of contexts that challenge more traditional approaches like density functional theory (DFT) and ground state coupled cluster (CC). These include charge transfer and doubly excited states,<sup>6,7</sup> core excitations,<sup>8</sup> thiophene,<sup>9</sup> and the carbon dimer.<sup>10</sup> Success in these areas is largely due to VMC's ability to combine Jastrow factors with a configuration interaction expansion carefully selected from quantum chemistry methods,<sup>9,11–17</sup> although much recent work has also investigated neural network wave function forms that have proven particularly accurate in small molecules.<sup>18–20</sup> In both approaches, it can be useful and sometimes critical to employ a quantum-chemistry-derived initial guess for the molecular orbitals (MOs) and other wave function components, and so the ability to interface effectively with Gaussian basis sets is crucial.

The most widely used approach to make VMC compatible with Gaussian basis sets and to deal with energy divergences near nuclei more generally is to employ pseudopotentials, also known as effective core potentials (ECPs).<sup>21-24</sup> In these approaches, the core electrons are removed and the divergent Coulomb potential near the nucleus is replaced with an effective potential designed to replicate the effect of the nucleus and core electrons on the valence electrons. As pseudopotentials lack Coulomb divergences, they remove the need for the orbital basis to produce KE divergences, making VMC directly compatible with Gaussian basis sets. Furthermore, core electrons typically make the largest contributions to VMC's energy variance, and so pseudopotentials also tend to reduce statistical uncertainty. They can also account for relativistic effects,<sup>25</sup> which can improve the energy compared to all-electron non-relativistic calculations on heavier atoms.<sup>26</sup> However, despite these clear advantages, pseudopotentials require careful management of the localization  $error^{27-29}$  and can be less accurate than all-electron calculations. For example, Wang et al. found that all-electron calculations for first row atoms outperformed those using pseudopotentials when evaluating ionization potentials.<sup>26</sup> More directly, pseudopotentials cannot be used to study processes involving core electrons, such as X-ray absorption spectroscopy.<sup>8</sup> Thus, while pseudopotentials are highly effective in many applications, all-electron calculations remain preferred for some. In those settings, electron-nuclear Coulomb divergences are present, and Gaussian basis sets alone are no longer sufficient.

To avoid energy divergences in all-electron calculations, the wave function must satisfy cusp conditions that ensure the KE diverges in a way that cancels the Coulomb divergence as particles coalesce.<sup>30</sup> In practice, electron-electron cusps are typically satisfied using Jastrow factors.<sup>31</sup> Electron-nuclear cusps can also be satisfied with Jastrow factors, but doing so often makes wave function optimization more difficult.<sup>8,31</sup> Alternatively, electron-nuclear cusps can be incorporated directly into

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the orbitals. One widely used approach is to replace Gaussian basis functions with a Slater-type orbital (STO) basis,<sup>32</sup> which by construction provide the necessary non-analytic features for cusps. However, employing STOs makes communicating with Gaussian-based quantum chemistry packages more difficult, and most do not support STOs directly (although there are exceptions<sup>33</sup>). Furthermore, even if an individual STO satisfies a particular cusp condition, linear combinations of STOs with different exponents, necessary for constructing accurate molecular orbitals, may not, which can further complicate wave function optimization within VMC.

To maintain easy compatibility with Gaussian and plane wave basis sets while achieving correct cusps at the nuclei, multiple groups have taken the approach of adding cusps to MOs that initially lack them.  $^{34\mathchar`-38}$  These approaches significantly lower the energy variance in VMC, and are directly compatible with Gaussian-type orbitals (GTOs). However, the precise outcomes of these MO-based approaches, along with similar methods that add cusps to the electron density,<sup>3</sup> depend on the shapes of the MOs and therefore on the manyelectron method (e.g. density functional) that was chosen that produced them. This method dependence stands in sharp contrast to the situation for AO basis sets, whose transferability stems in part from the fact that, once a basis is chosen, its properties depend only on the molecular geometry. In order to remove this method dependence while maintaining compatibility with the GTOs of quantum chemistry, we therefore explore in the present study the effectiveness of adding cusps directly to the AO basis functions themselves.

Specifically, we study an approach to constructing cusped Gaussian AOs in which each AO is cusped at each nuclei. In addition to making VMC's orbital basis set independent of the quantum chemistry method that supplies the initial wave function, it also ensures that the basis functions remain localized in space, a property that helps avoid unnecessary orbital evaluation costs in larger systems. Further, as with cusped MO approaches, any linear combination of these basis functions will also satisfy the nuclear cusps, simplifying the task of orbital optimization within VMC. To begin, we will describe our theoretical approach and its instantiation in a lightweight opensource software library, Cusping Gaussian Atomic Orbitals with Slaters (CGAOWS). We then test the cusped AOs in a collection of small molecules and analyze the results, finding that the good performance of previous MO-based methods can be maintained while simultaneously freeing the VMC orbital basis from any dependence on the many-electron method used to produce the initial MOs.

### **II. THEORY**

For a wave function to obey Kato's cusp condition<sup>30</sup> as an electron approaches nucleus I, it must satisfy

$$\left. \frac{\partial \langle \Psi \rangle}{\partial r} \right|_{r=0} = -Z_I \langle \Psi \rangle_{r=0} \tag{1}$$

where  $\langle \Psi \rangle$  is the average value of the wave function on the surface of the radius-*r* sphere centered on nucleus *I*.

If we wish the cusp condition in Eq. 1 be satisfied by the Slater determinant, then one may equivalently express the cusp condition in terms of the AOs  $\chi_i$  within the linear combination of atomic orbitals (LCAO) construction of each MO,  $\phi_a(r) = \sum_i c_{ia} \chi_i(r)$ .

$$\left. \frac{\partial \langle \chi_i \rangle}{\partial r} \right|_{r=0} = -Z_I \langle \chi_i \rangle_{r=0} \quad \forall \quad i, I \tag{2}$$

If this equation is satisfied for all AO-nucleus pairs, including in the tails of AOs centered on other nuclei, then the Slater determinant of MOs will satisfy the original cusp condition.

To address GTOs' failure to satisfy Eq. 2, we introduce a modification that smoothly transitions from the original GTO to a cusp-satisfying Slater-type function (cusped-STO) within a cusp radius  $r_c$  for each AO at each nuclear center. Any linear combination of cusp-satisfying AOs will thus result in a similarly cusp-corrected MO – as long as each AO is corrected over all nuclei in which it has a non-negligible magnitude. This approach (Fig. 1) (i) is MO-independent (ii) is independent of the method used to construct the initial trial wave function (iii) accommodates GTO basis sets from quantum chemistry, and (iv) can be applied before starting the QMC calculation.

Within a radius  $r_c$  sphere around each nucleus, we replace each GTO basis function  $\chi_n$  with a cusp-corrected basis function  $\tilde{\chi}_n$  that is defined within the  $r_c$  sphere as

$$\tilde{\chi}_n(r) = (1 - b(r)) \chi_n(r) + b(r) Q(r; \vec{q}_{nI})$$
(3)

where *r* is the distance to the nuclei in question and b(r) is the 5th-order polynomial

$$b(r) = c_1 r^5 + c_2 r^4 + c_3 r^3 + c_4 r^2 + c_5 r + c_6$$
(4)

that gradually switches from the GTO at  $r = r_c$  to the cusped-STO,  $Q(r; \vec{q})$ , at r = 0. Outside the cusp radius, the original GTO is left unchanged. For simplicity, and because we anyways want to disturb the original GTO as little as possible to maintain the character of the original basis set, we constrain the cusp radii so that the correction regions around different nuclei do not overlap. The precise coefficients within the switching function b(r) are chosen as  $c_1 = -6/r_c^5$ ,  $c_2 = 15/r_c^4$ ,  $c_3 = -10/r_c^3$ ,  $c_4 = 0$ ,  $c_5 = 0$ , and  $c_6 = 1$ , in order to ensure that b(0) = 1, that  $b(r_c) = 0$ , and that b'(0) = b''(0) = 0 $b'(r_c) = b''(r_c) = 0$ . This choice makes the updated orbital's first and second derivatives continuous at  $r = r_c$  and preserves the cusped-STO's satisfaction of the cusp condition at r = 0. It also ensures that, at the r = 0 and  $r = r_c$  boundaries, the KE contribution is coming entirely from the STO or GTO, respectively.

In Eq. 3, every AO is cusp-corrected around each nucleus at which it has an appreciable magnitude, meaning a value whose magnitude is at least  $10^{-15}$  times greater than the orbital's maximum magnitude. Thus, energy divergences are handled for AOs at their own nucleus and at other nuclei, as shown in Figs. 1c and 1d (note that the magnitude cutoff means that we do not add cusps to *p* or *d* orbitals at their own nuclei). An advantage to smoothly switching between



FIG. 1. Visualization of the cusp corrected orbitals  $(\tilde{\chi}(r))$  from a 6-31G(*d*) basis evaluated through the H-O bond in methanol, where H is the origin, see top right molecule. Cusp correction over H (O) correspond to purple (orange). (a) H 1s orbital, (b) H 2s orbital orthogonalized against the 1s, (c) O  $3p_x$  orbital, and (d) O  $3d_{z^2}$ . A zoomed in view of the cusps on the tails of (c,d) are to the right of their respective plots.

the GTO and a cusped-STO in the cusping region via a linear combination of the two is that the spherical asymmetry about a nucleus in the tail of a GTO (Fig.1c,d) can be maintained within  $\tilde{\chi}(r)$  such that the original GTO shape is not too strongly altered. Meanwhile, the spherical symmetry of the cusped-STO dominates near the nucleus and satisfies the necessary cusp condition.

To construct the cusped-STO function Q(r) for a specific GTO-nucleus pair, we begin with the overall goal that it satisfy the cusp condition so that the KE will cancel the electron-nuclear Coulomb divergence at the nucleus in question.

$$-\frac{1}{2}\frac{\nabla^2 Q(r)}{Q(r)} = \frac{Z}{r} \tag{5}$$

A natural starting point is the hydrogenic 1s orbital for that nucleus, and so we define an initial form  $Q(r; q_0)$  for Q as

$$Q(r; q_0) = q_0 e^{-Zr}$$
(6)

in which the constant  $q_0$  controls how much the cusp correction feature sticks out from the original GTO (see Fig. 1). We

determine  $q_0$  for each GTO-nucleus pair by minimizing the one-electron energy functional

$$\tilde{E} = \frac{\langle \chi_n | \hat{H} | \chi_n \rangle - \langle \chi_n | \hat{H} | \chi_n \rangle_{r < r_c} + \langle \tilde{\chi}_n | \hat{H} | \tilde{\chi}_n \rangle_{r < r_c}}{\langle \chi_n | \chi_n \rangle - \langle \chi_n | \chi_n \rangle_{r < r_c} + \langle \tilde{\chi}_n | \tilde{\chi}_n \rangle_{r < r_c}}$$
(7)

with the Broyden–Fletcher–Goldfarb–Shanno algorithm.<sup>40–43</sup> In Eq. 7, the denominator accounts for the AO's normalization,  $\hat{H}$  is the one-body Hamiltonian

$$\hat{H} = -\frac{1}{2}\nabla_{e}^{2} - \sum_{I} \frac{Z_{I}}{|\vec{r}_{e} - \vec{R}_{I}|}$$
(8)

for the electron e, and the  $r < r_c$  subscript indicates that the integrals within the indicated expectation values are truncated to the region of space within the radius  $r_c$  sphere about the nucleus in question. Essentially, these one-electron optimizations choose the  $q_0$  value for each GTO-nucleus pair that minimizes a simple approximation of the AO energy.



FIG. 2. Local energy (a.u.) (a) of the localized core carbon MO in CH<sub>3</sub>OH, 6-31G(*d*) basis, as an electron nears the carbon nucleus  $(r \rightarrow 0)$  of the non-cusp corrected GTO (black dashed) and the cusp corrected orbitals (solid colors) as a function of the indices included in the Slater-like basis expansion of Eq. 9 (*n*).  $r_c = 0.20$  Bohr. (b) Magnified plot of the cusp corrected AOs E<sub>L</sub> (a.u.) as a function of *n*. (c) The contribution of the cusped orbital to the variance of the local energy ( $\sigma_{E_r}^2$  a.u.<sup>2</sup>) in the sphere around the carbon.

While this initial choice of  $Q = Q(r; q_0)$  satisfies the cuspcondition at r = 0, we find that it does a poor job of managing KE fluctuations in the intermediate region in between r = 0and  $r = r_c$ . We therefore generalize this initial choice by including expanded radial flexibility in our final form for Q.

$$Q(r;\vec{q}) = q_0 e^{-Zr} + \sum_{i=2}^{M} q_i r^i e^{-Zr}$$
(9)

Note that, while we include terms up to M = 7, we omit the term linear in r, as its derivatives would interfere with satisfying the cusp condition. We determine the coefficients  $\vec{q}$  through variational optimization ( $\mathbf{H}\vec{q} = E \mathbf{S}\vec{q}$ ) within the radius  $r_c$  sphere around the nucleus. In this optimization, we simplify even further, neglecting the space outside  $r_c$  entirely and including only the kinetic energy and the attraction to the nucleus in question. The Hamiltonian and overlap matrices **H** and **S** are constructed in the basis  $\{(1-b(r))\chi_n(r), b(r)e^{-Zr}r^i; i = 0, 2, 3, ...\}$ , and the resulting lowest-energy eigenvector  $\vec{q}$  is scaled so that the coefficient on the  $(1-b(r))\chi_n(r)$  term of Eq. 3 equals one, which ensures that the boundary conditions at  $r_c$  remain satisfied.

In Fig. 2, we show how the local energy  $E_{L}$  of the localized 1s carbon MO in methanol behaves for different truncations in the order of M. A true energy eigenstate would have a constant E<sub>L</sub>, and therefore zero-variance.<sup>44</sup> The cusped orbital should ideally make E<sub>L</sub> finite everywhere and minimize the amplitudes of the "wiggles" seen in the GTO's EL in order to both satisfy the central limit theorem and reduce the variance within a VMC calculation. Note that, for  $M \ge 2$ , there are larger fluctuations of  $E_{L}$  near the origin but smaller fluctuations in the remainder of the  $r < r_c$  region. Since the probability of sampling very close to the nucleus (e.g. r < 0.02) is much smaller than sampling in the remainder of the  $r < r_c$ region, this behavior is an expected trade off made by the variational optimization that should reduce the overall VMC energy variance. Indeed, by plotting the net contributions to the variance at different radii (Fig. 2c), we see that this trade off is exactly what is happening, although we also note that the ability to reduce variance contributions near  $r = r_c$  ( $r_c = 0.2$ in this example) is limited by the constraint that we match the GTO at the boundary, limiting the final orbital quality there to that of the original GTO.

The expansion with M = 7 (red, Fig. 2a,b) was chosen for the final form of Q(r), as it most significantly decreases the cusp region's contribution to the energy variance  $\sigma^2$ , as demonstrated in Fig. 2c. Beyond M = 7, the basis functions rapidly become linearly dependent, limiting the quality of the optimization. In principle, one could adopt a different basis in which the functions are orthogonal by construction (e.g. Chebyshev polynomials), but, as we will see in the general results, stopping at M = 7 is sufficient to produce results comparable to previous MO-based cusping schemes.

In our cusping package, CGAOWS, we have chosen specific cusp radii for all nuclei up to Z = 10 so that the cusp corrections do not overlap and do not disturb the original GTO shapes more than necessary, so as to retain as closely as possible the character of the original basis. Our choices were guided by using methanol as a test case, in which we sought to find a balance between variance and average  $E_L$  reduction while maintaining generalizability across the various AO and Z combinations. The default cusp radius for s orbitals is  $r_c = 0.2$  Bohr, except in the case of a hydrogen s orbital being cusped at a hydrogen nucleus, in which case  $r_c = 0.1$  Bohr. For p or d orbitals, the default is  $r_c = 0.075$  Bohr. These parameters were chosen given that they produce a satisfactory reduction in the variance, however, further optimization is possible. They can be readily tuned within the software for specific molecular systems and basis sets if desired.

Finally, in order to increase the efficacy of the chosen cusp radii, we redefine s orbital GTOs with principle quantum number 2 or higher by Gram-Schmidt orthogonalizing them against their corresponding 1s orbital. Note that this does not change the span of the basis set, but it does produce more uniformity in the shapes of GTO s orbitals near their centers, making a one-size-fits-all cusp radius for s orbitals more effective. An example of this approach is seen in Fig. 1b, where the r = 0 behavior of the 2s basis function has been made similar to that of the 1s function via orthogonalization. Note that, as seen in Figs. 1c and 1d, we do not orthogonalize against other nuclei's 1s orbitals when adding cusps for those nuclei, as this would decrease the locality of the basis functions. Happily, we find that the default cusp radius, along with the optimization of Q(r), produces good results for those cusps. For an s orbital's own nucleus, however, we perform the orthogonalization as we find that it significantly improves the quality of the results. Of course, when importing initial MOs from quantum chemistry into VMC, the LCAO coefficient matrix must be updated to reflect this redefinition of higher principle quantum number s basis functions.

### III. RESULTS

The performance of the presented cusping scheme is tested using a restricted single Slater determinant trial wave function constructed with Foster-Boys<sup>45</sup> localized Hartree-Fock orbitals from  $PYSCF^{46-48}$  and a 6-31G(*d*) Pople basis set.<sup>49</sup> By using a single Slater wave function without a Jastrow factor, all benefit in variance reduction may be attributed to the cusp-correction. All VMC calculations from this work were performed on molecular geometries optimized at the MP2/6-31G(d) level of theory.<sup>50</sup> As tests cases, neon (Fig. 3) and methanol (Fig. 4) are used to demonstrate the impact of this cusp correction scheme on atomic and molecular systems, respectively, by evaluating  $E_L$  as an electron "walks" across the nucleus. Following this, results on a test set of 16 small gas phase molecules are presented in Tables I and II; the former containing statistical analysis assuming normally-distributed data (i.e. mean energy and variance) and the latter assuming non-normally distributed data (*i.e.* median energy, interquartile range, and range). Four of the molecules in this test set including LiH, C<sub>2</sub>H<sub>6</sub>, N<sub>2</sub>, and CO<sub>2</sub>, chosen based on data availability, are compared against other relevant cusping schemes in Table I. All VMC calculations from this work were computed using our own VMC software, augmented by the newly developed CGAOWS package. For details about the software, see the SI.



FIG. 3. Local energy (a.u.) of single Slater determinant wave function, no Jastrow factor, as an electron "walks" across the nucleus of a neon atom, in a 6-31G(*d*) basis.  $r_c = 0.2$  Bohr for all AOs, the shaded area outside of  $|r| > r_c$  indicates region of no cusp-correction. (a) Uncusped GTO basis (black dashed) and cusp-corrected basis (black solid). (b) Magnified E<sub>L</sub> of the cusp-corrected basis.

## a. Neon and Methanol

As an initial demonstration of the presented cusping scheme, Fig. 3 depicts E<sub>L</sub> of the Ne atom as an electron is moved through the nucleus, holding all other electrons fixed. An initial VMC calculation, with appropriate burn-in, was performed to produce an electron configuration distributed with respect to the wave functions probability distribution. The alpha spin electron closest to the nucleus was then repositioned to move through the nucleus along the x-axis. Fig. 3 compares both the uncusped- (dashed line) and cusped- (solid line) GTO bases. While the details of the fluctuations in  $E_L$ for both cases depend on the positions of the other electrons, the general behavior is consistent when other configurations are put through the same test. Note, outside of neon's cusp radius of 0.2 Bohr the cusped-GTO and GTO appear to return to evaluating the same orbital, however, in actuality there is a slight, unobservable difference between the two. This difference is expected, as it is caused by other electrons residing within the cusp radii of the atom in this particular configuration of the electrons, as would be expected in atomic or molecular systems.

As expected, the GTO basis yields significant oscillations and a divergence of  $E_L$  as the electron nears the nucleus. In the cusped-GTO basis, these oscillations are dramatically reduced and  $E_L$  remains finite everywhere. This result closely resembles the behavior of the MO-based cusping scheme shown in Fig. 4 of Ma et al.<sup>34</sup> The remaining subtle "wiggles" visible in the cusped-GTO case (Fig. 3b) are a byproduct of the choice of polynomial switching function in Eq. 4 that controls the interpolation between the pure GTO and the cusped-STO of Eq. 9. Two paths could be taken to further reduce the remaining fluctuations in  $E_L$  near the nucleus. First, a higherorder polynomial expansion, using an orthogonal construc-



FIG. 4. Local energy (a.u.) of single Slater determinant wave function, no Jastrow factor, as an electron "walks" across the carbon nucleus of methanol, see slice through molecule for the trajectory, in a 6-31G(*d*) basis. Light gray shading outside of |r| = 0.075 indicates the cusp radii around the carbon nucleus for the oxygen atom's *p* and *d* orbitals. The dark gray shading outside of |r| = 0.2 indicates the cusp radii around the carbon nucleus for all other basis functions.

tion to avoid linear dependency issues, could be employed. Second, adding similar corrections to non-*s* basis functions in the  $r < r_c$  region, although not necessary to cancel any divergences, would improve their kinetic energy behavior near the nucleus. We forgo these steps in the present study, because the M = 7 expansion applied only to *s* orbitals already achieves reductions comparable to previous MO-based schemes.<sup>34</sup>

To verify that similar behavior is achieved in a molecular system, where adding cusps for other nuclei in the tails of each AO now matters, we perform a similar test in methanol. Using the same procedure as in neon, Fig. 4 depicts  $E_L$  as an electron moves through the carbon nucleus. As in neon, we see a large reduction in local energy fluctuations compared to the original GTO basis. More significantly than neon, we see for methanol a difference in the local energy values of the cusped and uncusped results outside of the carbon atom's outermost cusp radius of 0.2 Bohr as well. Indeed, as one looks at larger and larger molecules, the chances of sampling a configuration in which none of the electrons lie within any of the nuclei's cusp radii becomes vanishingly small.

### b. Standard test set of small molecules

We performed a series of calculations on 16 small organic molecules taken from the dataset of Ma *et al*<sup>34</sup> in order to compare the performance of our AO-based scheme with existing MO-based approaches. As with all previous work to which we compare here, calculations were performed without a Jastrow factor in order to isolate the impact of the electronnuclear cusp correction on the calculations. In Table I, we report the average energy and its variance for both the uncusped and cusped cases, evaluating the uncertainty in each by as-

			Not cusp corrected		Cusp corrected		
Molecule		$E_{ m HF}$	Mean <i>E</i> <sub>VMC</sub>	$\sigma^2$	Mean <i>E</i> <sub>VMC</sub>	$\sigma^2$	Variance reduction (%)
LiH	This work	-7.9808	$-7.982 \pm 0.001$	$8.3\pm0.8$	$-7.9817 \pm 0.0006$	$1.68\pm0.02$	79.81
	Kussman	-7.9809	$\textbf{-7.980}\pm0.001$	$7.14\pm0.02$	$-7.9826 \pm 0.0006$	$1.76\pm0.01$	75.35
	Ma	-7.9859	$\textbf{-7.984} \pm 0.002$	$6.7\pm0.9$	$-7.9864 \pm 0.0005$	$1.64\pm0.03$	75.52
	Per	-7.9855	$-7.9845 \pm 0.0015$	$8\pm 2$	$-7.98536 \pm 0.00049$	$1.607\pm0.016$	79.91
$C_2H_6$	This work	-79.2280	$-79.24\pm0.02$	$275\pm50$	$-79.231 \pm 0.002$	$19.9\pm0.3$	92.78
	Kussman	-79.2285	$\textbf{-79.22}\pm0.01$	$214\pm3$	$-79.237 \pm 0.005$	$16.1\pm0.4$	92.48
	Ma	-79.2567	$\textbf{-79.24} \pm 0.02$	$156\pm31$	$-79.259 \pm 0.002$	$18.0\pm0.2$	88.46
	Per	-	-	-	-	-	-
$N_2$	This work	-108.9343	$-108.95 \pm 0.03$	$453\pm90$	$-108.936 \pm 0.004$	$28.4\pm0.2$	93.72
	Kussman	-108.9354	$\textbf{-108.9}\pm0.1$	$390\pm8$	$-108.927 \pm 0.006$	$25\pm1$	93.59
	Ma	-108.9710	$-108.96\pm0.02$	$308\pm76$	$-108.973 \pm 0.003$	$25.1\pm0.3$	91.85
	Per	-108.961	$-108.912 \pm 0.030$	$276\pm76$	$-108.9677 \pm 0.0033$	$24.76\pm0.46$	91.03
$CO_2$	This work	-187.6275	$-187.62 \pm 0.04$	$774\pm231$	$-187.614 \pm 0.005$	$53.0\pm0.6$	93.16
	Kussman	-187.6284	$\textbf{-187.60}\pm0.01$	$598\pm9$	$-187.635 \pm 0.009$	$45\pm2$	92.47
	Ma	-187.6880	$\textbf{-187.62}\pm0.02$	$348\pm39$	$-187.691 \pm 0.003$	$44.9\pm0.5$	87.10
	Per	-	-	-	-	-	-
	This work						
Li <sub>2</sub>	-	-14.8664	$-14.869 \pm 0.003$	$21 \pm 4$	$-14.8676 \pm 0.0009$	$3.18\pm0.05$	84.79
$CH_4$	-	-40.1947	$-40.182 \pm 0.008$	$82\pm13$	$-40.1984 \pm 0.0009$	$9.9\pm0.1$	87.85
$C_2H_2$	-	-76.8149	$\textbf{-76.80} \pm 0.01$	$198\pm27$	$-76.820 \pm 0.002$	$19.4\pm0.7$	90.20
$C_2H_4$	-	-78.0306	$\textbf{-78.03}\pm0.02$	$227\pm44$	$-78.030 \pm 0.002$	$19.1\pm0.2$	91.60
HCN	-	-92.8693	$\textbf{-92.89} \pm 0.01$	$398\pm 63$	$-92.870 \pm 0.003$	$25\pm1$	93.82
NH <sub>3</sub>	-	-56.1832	$-56.17\pm0.01$	$145\pm19$	$-56.189 \pm 0.003$	$15.5\pm0.7$	89.30
H <sub>2</sub> CO	-	-113.8629	$-113.78\pm0.02$	$248\pm24$	$-113.858 \pm 0.004$	$31.0\pm0.4$	87.51
CH <sub>3</sub> OH	-	-115.0331	$-115.06\pm0.03$	$646 \pm 147$	$-115.038 \pm 0.003$	$31.4\pm0.7$	95.14
H <sub>2</sub> O	-	-76.0084	$-76.01\pm0.02$	$204 \pm 24$	$-76.006 \pm 0.004$	$22.3\pm0.5$	89.09
$H_2O_2$	-	-150.7577	$-150.74 \pm 0.06$	$581 \pm 133$	$-150.744 \pm 0.004$	$43.8\pm0.5$	92.46
O <sub>2</sub>	-	-149.5187	$-149.7\pm0.2$	$8732\pm8148$	$-149.512 \pm 0.004$	$43.6\pm0.8$	99.50
HF	-	-100.0002	$\textbf{-99.97} \pm 0.04$	$532\pm200$	$-99.992 \pm 0.003$	$32.5\pm0.2$	93.89

TABLE I. Comparison of HF energies ( $E_{\text{HF}}$ ), VMC energies ( $E_{\text{VMC}}$ ) with and without electron-nuclear cusps, variances ( $\sigma^2$ ), and standard errors ( $\pm$  following VMC results), all in a.u., for this work and work by Kussman *et al.*<sup>39</sup>, Ma *et al.*<sup>34</sup>, and Per *et al.*<sup>35</sup> All results employed a Hartree-Fock-based single Slater determinant with no Jastrow factor. See text for geometry, basis set, and sampling details.

suming normal statistics. We should stress, however, that the absence of a Jastrow factor means that electron-electron cusps are missing, and so the local energy diverges when electrons approach each other closely. This issue makes the quantities in Table I difficult to interpret, as these divergences violate the central limit theorem and lead to data that is not normally distributed. Indeed, a standard quantile-quantile analysis of our data revealed it to be non-normal with a significant rightward skew, an effect that was even more pronounced in the uncusped data where electron-nuclear divergences are also present. Despite this issue, we begin with an analysis based on the assumption of normal statistics in order to make an apples-to-apples comparison to previous work. After this comparison, we will present an analysis in which we make no assumptions about normality.

In the first half of Table I, we compare our results for LiH,

 $C_2H_6$ ,  $N_2$  and  $CO_2$  to Kussman *et al.*<sup>39</sup>, Ma *et al.*<sup>34</sup>, and Per *et al.*<sup>35</sup> (only for LiH and  $N_2$ ). Kussman *et al.* use the same 6-31G(d) basis while Per *et al.* use a 6-311G(d) basis. Ma *et al.* do not specify which basis they use for their small-molecule analysis, but their reported HF energies suggest it is a larger basis than either 6-31G(d) or 6-311G(d). These works also employed varying amounts of statistical sampling. In this work, energy and variance uncertainties are evaluated using 5 million total samples by treating the results of ten independent 500,000-sample calculations as ten independent draws from an unknown normal distribution. Kussman *et al.* used 500,000, and Per *et al.* left precise sample numbers unspecified.

Overall, all four approaches provide similar variance reductions compared to uncusped results. However, due to the use of different basis sets and sampling efforts, and the fact



FIG. 5. Local energies for 500,000 sample excerpts from (a) LiH (b)  $C_2H_6$  (c)  $N_2$  (d)  $CO_2$  (6-31G(*d*)). Left plot without cusps, right with cusps. As there is no Jastrow factor, the remaining outliers after nuclear cusp corrections are applied are explained by the lack of electron-electron cusps.

that none of these tests are expected to produce normally distributed data, it is not possible to say with precision which schemes offer larger variance reductions in which molecules. For example, Kussman *et al.* used the same basis set and MP2/6-31G(d) reference geometries, yet the uncusped variance uncertainties they report are at least an order of mag-

TABLE II. Comparison of the median, interquartile range (IQR), and total range of the local energy in a.u. in uncusped and cusped VMC calculations for the 16 molecules studied in this work. The sample size in each calculation is 5 million.

	Not cusp corrected			Cusp corrected		
Molecule	Median	IQR	Range	Median	IQR	Range
LiH	-8.159	0.984	1509	-8.192	0.923	256
$C_2H_6$	-79.663	4.997	7653	-79.913	3.982	1290
$N_2$	-109.410	6.197	6999	-109.778	4.719	942
$CO_2$	-188.180	9.308	11038	-188.717	6.770	1191
Li <sub>2</sub>	-15.126	1.508	4164	-15.183	1.345	313
$CH_4$	-40.500	3.060	4306	-40.691	2.606	681
$C_2H_2$	-77.252	4.825	5837	-77.498	3.812	1522
$C_2H_4$	-78.466	4.916	7295	-78.712	3.898	1146
HCN	-93.329	5.526	7254	-93.634	4.276	2543
NH <sub>3</sub>	-56.510	3.892	4385	-56.789	3.176	2042
H <sub>2</sub> CO	-114.326	6.475	4981	-114.731	4.924	942
CH <sub>3</sub> OH	-115.498	6.565	8653	-115.902	5.004	1416
H <sub>2</sub> O	-76.353	4.893	3671	-76.749	3.829	1120
$H_2O_2$	-151.260	8.013	5467	-151.762	5.921	1367
O <sub>2</sub>	-150.028	7.912	45765	-150.543	5.815	2219
HF	-100.376	6.073	12371	-100.846	4.687	1155

nitude smaller than ours. This difference appears to be due to our larger sample size, which makes our tests much more likely to encounter samples with near-divergent local energies arising from electron-nuclear coalescence. Although this issue makes uncertainties hard to interpret, the AO-based cusping scheme nonetheless shows very similar levels of variance reduction as the MO-based approaches for the molecules that are shared in common, and similarly large variance reductions in the remaining molecules shown at the bottom of Table I. In Fig. 5, we show detailed sampling data for the same four molecules compared to other work in Table I in order to highlight both the non-normality of the data and the degree to which non-normality is reduced when electron-nuclear cusps are added to the AOs. First, we note that the vertical axis scale is in thousands of Hartree, which immediately makes clear that, as expected, the data contains near-divergences due to particle coalescence that are not consistent with a normal distribution. Second, we see that the addition of nuclear cusps in the AOs dramatically reduces these near-divergences, but that some still remain due to electron-electron coalescence. Thus, on the one hand, we are reassured that the addition of electronnuclear cusps is helping a great deal, but, on the other hand, we would prefer to augment the discussion above with measures of improvement that do not assume normally distributed data.

To this purpose, Table II reports the median, interquartile range, and total range of the local energy across all 5 million samples for each molecule in our test set. In all cases, the local energy median and interquartile range decrease when cusps are added to the AOs, indicating both a variational improvement in the energy and a reduction in the local energy's statistical spread that is consistent with the variance analysis above. As for the near near-divergences seen in Fig. 5, these can be quantified via the total range measure. By eliminating electron-nuclear divergences, the AO cusp correction removes the most negative divergences, but it has little to no impact on the positive divergences related to electron-electron coalescence. Thus, the total range is significantly reduced by the AO cusps but remains large, as seen in Table II. Nonetheless, these more general metrics all indicate that the expected statistical improvements, which were tentatively indicated by an analysis assuming normal statistics, are indeed real.

# IV. CONCLUSION

In conclusion, adding nuclear cusps directly to Gaussian atomic orbitals offers similar advantages as molecular-orbitalbased cusping methods while maintaining independence from the many-body method that generates the molecular orbitals and other trial function parameters. In particular, we have shown that an interpolation between the original Gaussian type orbital and a Slater-based function with a low-order polynomial radial function produces electron-nuclear cusps that substantially improve the statistics in VMC evaluations. As with other cusping schemes, the update to the atomic orbitals can be applied cheaply before starting any QMC sampling. For convenience, we have provided a lightweight standalone library, CGAOWS, that provides code both for determining the cusps and evaluating the resulting cusped atomic orbitals. All told, this approach provides all-electron VMC with local and transferrable atomic orbital basis sets that, being based on those from quantum chemistry, interface easily with quantum chemistry methods.

#### SUPPLEMENTARY INFORMATION

See Supplementary Information for electron positions used in the case studies from Section III a and a description of the CGAOWS software library.

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#### DATA AVAILABILITY STATEMENT

The data that support the findings of this study are available within the article and its supplementary material.

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### V. SUPPLEMENTARY INFORMATION

#### a. Electron Positions and Geometries for Nuclear Walk-Through Tests

TABLE S1. Coordinates of the frozen electrons in the neon walk-through test (see Figure 3 in the paper). The nucleus is centered at the origin. A and B denote alpha and beta spin, respectively. A1 corresponds to the electron being walked through the nucleus. Coordinates are in Bohr.

Electron	Х	у	Z
A1	-	_	_
A2	0.213553	-0.444243	0.278260
A3	-0.494929	0.584347	-0.265310
A4	0.372814	0.868481	-1.627833
A5	-0.900304	-1.472222	-1.127845
B1	-0.919444	-0.636217	-0.295044
B2	1.001112	0.185600	-0.238352
B3	-0.598423	-0.200966	0.797388
B4	-0.274741	0.465994	-0.980761
B5	0.062476	-0.088619	-0.033220

TABLE S2. Molecular geometry of methanol for walk-through test (see Figure 4 in the paper). Coordinates are in Bohr.

$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Ζ	Х	у	Z
O         -0.088460         -1.434354         0.00000           H         -2.060765         1.850028         0.00000           H         0.829924         2.030361         1.68626	С	-0.088460	1.259762	0.000000
H -2.060765 1.850028 0.00000 H 0.829924 2.030361 1.68626	0	-0.088460	-1.434354	0.000000
Н 0.829924 2.030361 1.68626	Η	-2.060765	1.850028	0.000000
	Н	0.829924	2.030361	1.686264
Н 0.829924 2.030361 -1.68626	Н	0.829924	2.030361	-1.686264
Н 1.639359 -1.994492 0.00000	Н	1.639359	-1.994492	0.000000

TABLE S3. Coordinates of the frozen electrons in the methanol walk-through test (see Figure 4 in the paper). A and B denote alpha and beta spin, respectively. A1 corresponds to the electron being walked through the carbon nucleus. Coordinates are in Bohr.

		-	
Electron	Х	У	Z
A1	-	-	-
A2	0.106702	0.914076	0.381221
A3	0.763427	1.029360	0.300633
A4	-1.214422	1.824894	-0.535917
A5	-1.091921	1.072275	1.074244
A6	-0.398179	-1.133658	-0.590961
A7	-0.095003	-1.563129	-0.045628
A8	-0.126748	-1.720242	0.257205
A9	1.161947	-2.956023	-0.126195
B1	-0.222989	-0.861524	-0.514105
B2	1.083038	-0.238063	1.820019
B3	-0.179891	-1.328974	0.116196
B4	0.653737	-2.122696	-0.284307
B5	-1.353428	2.401788	1.105960
B6	-1.211755	-2.277452	-0.687876
B7	0.697330	1.929689	1.448539
B8	0.063754	1.337925	-0.142632
B9	-0.084558	0.904812	-1.026934

### b. Cusping Gaussian Atomic Orbitals with Slaters (CGAOWS) Package

We provide an open-source C++ library that produces cusped atomic orbitals for a chosen molecular system and basis set (currently limited to STO-3G, 6-31G, and 6-31G(d)). It can be found here: https://github.com/eneuscamman/cgaows. After the

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user specifies the molecular details in the input file, Python code is used to calculate the cusp parameters and molecular orbital coefficient matrix that takes into account s-type orbital orthogonalization. These parameters are printed into txt files that are subsequently read into the C++ code. We provide functions that evaluate the AOs for a given electron configuration as well as the first and second orbital derivatives (cartesian) required for the evaluation of the kinetic energy. We include an example (Ne atom/6-31G(d) basis set) of how the package can be integrated into a variational Monte Carlo code. See the code for further details.