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IDENTIFICATION OF THE NEUTRON-RICH ISOTOPE [SUP]174 ER

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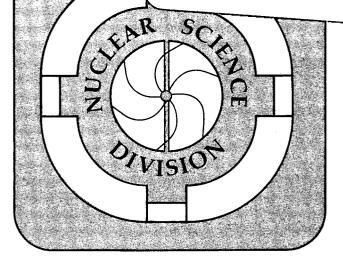
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R.M. Chasteler, J.M. Nitschke, R.B. Firestone, K.S. Vierinen, P.A. Wilmarth, and A.A. Shihab-Eldin

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Identification of the Neutron-Rich Isotope ¹⁷⁴Er

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<u>Abstract:</u>

The heaviest known erbium isotope, 3.3(2) min. ^{174}Er , has been identified using on-line isotope separation and by β and γ -ray spectroscopy. Twelve γ rays were placed in a preliminary decay scheme. The mass of the new isotope was established by the separator and the Z was deduced from Tm K x ray coincident β and γ rays. In addition, the heaviest known isotope of holmium 171 Ho was observed with a half-life of 49(5) s.

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There has been much progress recently in the search for new neutron-rich isotopes in the rare-earth mass region $^{1-5}$ The experiments can be classified according to the different production techniques. Several rare-earth isotopes have been produced using online isotope separation of $^{252}Cf^1$ and $^{235}U^2$ fission products; however, low fission yields limit this approach to the lighter neutron-rich isotopes. The heaviest known neutron-rich isotopes of Tm, Yb, and Lu have been produced at GSI using multinucleon transfer reactions followed by on-line mass separation. In this method, $^{nat}W/Ta$ targets are bombarded with $^{136}Xe^{3,4}$ or $^{186}W^5$ to produce target-like neutron-rich isotopes. Similar multinucleon transfer reactions have been explored at the OASIS mass-separation facility⁶ on-line at the Lawrence Berkeley Laboratory SuperHILAC. Since the targets are located in the high temperature region of our ion source, the only target choices are heavy refractory materials like W or Ta. To overcome this restriction, projectiles of the heaviest stable lanthanide isotopes, such as 176 Yb, were used to produce projectile-like neutron-rich rareearth isotopes.

Natural tungsten targets were bombarded with 8.5 MeV/u ¹⁷⁶Yb ions and multinucleon transfer reactions were used to produce projectile-like isotopes of Ho, Er, and Tm. Using the stable mass marker ¹⁷⁴Yb to calibrate the separator, the reaction products were mass analyzed and A=174 isobars were transported ionoptically to a shielded counting area, collected with a fast-cycling tape system, and transported (within 70 ms) to a detector array for β and γ -ray spectroscopy. A 718 μ m Si detector and a planar hyperpure Ge (HPGe) detector faced the radioactive layer to detect x rays, low-energy (\leq 400 keV) γ rays, and β particles (utilizing the two detectors as a β telescope). A 1-mm thick plastic scintillator, for electron detection, and a 52% Ge detector, for high efficiency x-ray and γ -ray intensities essentially free of summing effects. Coincidence events registered in the various detectors were recorded in an event-by-event mode while singles spectra were acquired from all three Ge detectors concurrently. The singles spectra from the 52% Ge

and the HPGe detectors were acquired in a time resolved multispectrum mode where the 512-s tape cycle was divided into eight equal time intervals for half-life determinations.

Twelve γ rays with energies, relative intensities, and coincidences listed in Table I are assigned to the decay of the new isotope ¹⁷⁴Er based on the following observations: the γ rays decayed with a 3.3(2) min. half-life which does not match any of the known half-lives in this isobaric chain, the γ rays were observed in coincidence with Tm K x rays, the γ rays were measured in coincidence with β particles (in the β telescope) having energies up to ~1.3 MeV, and the γ rays were not in coincidence with any of the known⁷ γ rays from the decay of ¹⁷⁴Tm. The first two points are consistent with a possible isomer of ¹⁷⁴Tm or the new isotope ¹⁷⁴Er; however, the ¹⁷⁴Tm isomer is ruled out by the last two points. The 58.5-keV γ transition was not seen in the singles data due to interfering x-ray activity at that energy, but it was clearly identified in the coincidence spectra. The low energy part of the HPGe detector singles spectrum is shown in figure 1. Based on the coincidence and intensity information, we propose the ¹⁷⁴Er decay scheme in figure 2. The two strong β transitions from the ¹⁷⁴Er 0⁺ ground state to the 766.9- and 772.9-keV levels are similar to the 0⁺ to 1⁺ and (1)⁻ β transitions in ¹⁷²Er decay.⁸ Assuming then a spin difference of one unit between the the ¹⁷⁴Er ground state and the 770-keV levels, the observed strong γ transitions to our proposed ground state in fig. 2 are inconsistent with a tentative 4⁻ assignment for the ¹⁷⁴Tm ground state from ref. 7.

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An average half-life of 3.3(2) min. was measured for the γ rays and the Tm K x rays (excluding the 58.5-keV transition). This is longer than the theoretical prediction of 1.3 min. by Klapdor *et al.*,⁹ and shorter than the value of 15 min. calculated by Takahashi *et al.*¹⁰ However, recalculations with the gross theory of β decay¹⁰ using the more recent mass predictions of Jänecke and Masson¹¹ (Q_β=1.99 MeV) yield a half-life of about 3.3 min., in much better agreement with experiment.

During the same experiment, some preliminary data on the A=171 mass chain were obtained. We detected erbium K x rays with a half-life of 49(5) s. These x rays were seen in coincidence with β^- particles having energies up to ~2.5 MeV and in weak coincidence with the known¹² 198- and 280-keV γ transitions in ¹⁷¹Er. Previous authors^{13,14} reported a 45(5) s component in the electron spectrum recorded with a thin plastic scintillator and weak erbium K x rays in coincidence with electrons in the A=171 mass chain, and made a tentative assignment of this activity to a new isotope ¹⁷¹Ho or a new isomer ^{171m}Er decaying by a strongly converted transition. Based on our measurements of the Er x rays in coincidence with β^- particles and ¹⁷¹Er transitions, we can unambiguously assign the 49(5) s activity to the β^- decay of ¹⁷¹Ho and eliminate the ^{171m}Er alternative interpretation.

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Table L	Gamma-ray	venergies.	intensities.	and	coincidences	in the	e decay of 174	[‡] Er.
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Energy ^a (keV)	Ιγ ^b (%)	Coincident γ rays ^c (keV)
58.5 ^d		
71.7	15(1)	X,(58.5),(122),637,643
100.4	100(1)	X,152
121.8	39(6)	X,(58.5),(71.7),130
130.4	42(8)	122,637,643
151.8	91(15)	X,100.4
636.7	27(5)	X,(58.5),(71.7),130
642.7	25(2)	X,(58.5),(71.7),130
708.4	93(7)	(58.5)
714.4	56(6)	(58.5)
766.9	92(9)	
772.9	75(8)	

^a The uncertainty in γ -ray energies is ± 0.2 keV. ^b Intensities are relative to the 100.4-keV γ intensity. ^c X=Tm K x rays. () denotes a weak coincidence. ^d Not seen in the singles data. See text.

10⁶ 95.4 keV ¹⁷⁴Tm 174 Er -52.4 keV Yb K α_1 10⁴ -51.3 keV Yb K α_2 Pb Kβ₂ 00 COUNTS / 0.213 keV 50.7 keV Im Ka, 10⁵ 100 49.8 keV Tm K α_1 90 95 76.6 keV ¹⁷⁴Tm -61.0 keV Yb KB₂ -57.5 keV Tm Kβ, -- Pb Kα, 71.7 keV ¹⁷⁴Er -Pb Kα, AυKα -Au Ka, 10⁴ Au KB, 60 65 ENERGY (keV) 45 50 55 70 75 80

Fig. 1. Low-energy part of the γ -ray spectrum from the A=174 activities.

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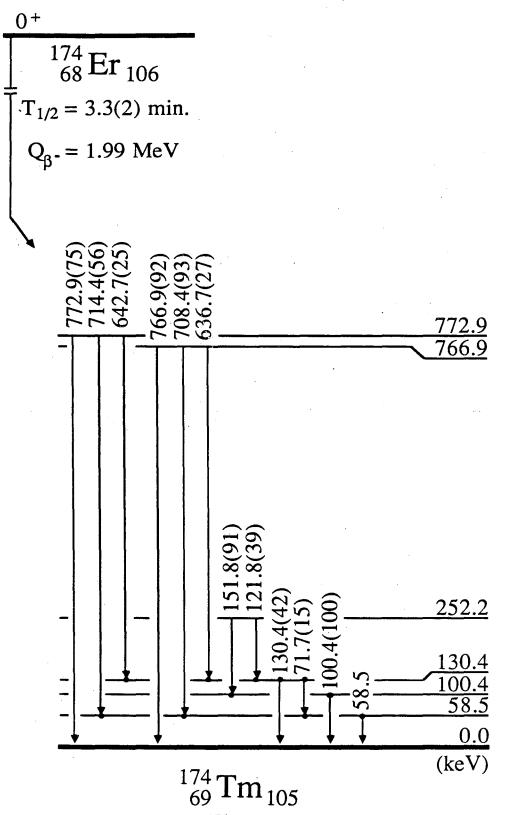


Fig. 2. Proposed decay scheme for 174 Er. Intensities are relative to the 100.4-keV transition. The beta decay energy is taken from reference 11.

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