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1 Tropical weathering of the Taconic orogeny as a driver for
2 Ordovician cooling

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8 **ABSTRACT**

9 The Earth's climate cooled through the Ordovician Period leading up to the
10 Hirnantian glaciation. Increased weatherability of silicate rocks associated with
11 topography generated on the Appalachian margin during the Taconic orogeny has been
12 proposed as a mechanism for Ordovician cooling. However, paleogeographic
13 reconstructions typically place the Appalachian margin within the arid subtropics, outside
14 of the warm and wet tropics where chemical weathering rates are highest. In this study
15 we reanalyze the paleomagnetic database and conclude that Ordovician constraints from
16 cratonic Laurentia are not robust. Instead, we use paleomagnetic data from well-dated
17 volcanic rocks in the accreting terranes to constrain Laurentia's position given that the
18 Appalachian margin was at, or equatorward of, the paleolatitude of these terranes. To
19 satisfy these allochthonous data, Laurentia must have moved toward the equator during
20 the Ordovician such that the Appalachian margin was within 10° of the equator by 465
21 Ma. This movement into the tropics coincided with the collision and exhumation of the
22 Taconic arc system, recorded by a shift in neodymium isotope data from shale on the

23 Appalachian margin to more juvenile values. This inflection in detrital neodymium
24 isotope values precedes a major downturn in global seawater strontium isotopic values by
25 more than one million years, as would be predicted from a change in weathering input
26 and the relatively long residence time of strontium in the ocean. These data are consistent
27 with an increase in global weatherability associated with the tropical weathering of mafic
28 and ultramafic lithologies exhumed during the Taconic arc-continent collision. A Taconic
29 related increase in weatherability is a viable mechanism for lowering atmospheric CO₂
30 levels through silicate weathering contributing to long-term Ordovician cooling.

31 **INTRODUCTION**

32 Ordovician strata record the transition from an Early Ordovician ice-free world to
33 end-Ordovician glaciation and mass extinction (Cooper and Sadler, 2012). Several
34 hypotheses have been proposed to account for this cooling and the initiation of glaciation
35 including: increased carbon burial (Brenchley et al., 1994), aerosol release from
36 volcanism (Buggisch et al., 2010), decreased volcanic outgassing (McKenzie et al.,
37 2014), increased silicate weathering due to topography associated with the Taconic
38 orogeny (Kump et al., 1999), and increased weathering of fresh volcanic rocks (Young et
39 al., 2009). Oxygen isotope data from brachiopods and conodonts indicate that Hirnantian
40 glaciation is the culmination of longer term cooling from 480 to 445 Ma (Trotter et al.,
41 2008; Veizer and Prokoph, 2015). Although short-term perturbations such as increased
42 organic carbon burial inferred from positive carbon isotope excursions, changes in ocean
43 circulation, or sulfur aerosol release could account for transient cooling associated with
44 the Hirnantian glacial maximum, tectonic changes associated with long-term changes to
45 CO₂ sources or sinks are required to drive ~35 m.y. of cooling. An increase in

46 global weatherability can lead to CO₂ levels decreasing through increased silicate
47 weathering, associated delivery of alkalinity to the ocean, and sequestration of
48 bicarbonate in chemical sediments. The silicate weathering feedback can lead to
49 stabilization at a lower steady-state CO₂ level (Kump et al., 1999).

50 Arc-continent collision is a tectonic process that can combine the mechanisms for
51 cooling outlined here and lead to a decrease in volcanic outgassing through the death of
52 an arc, and an increase in silicate weathering through increased topography and the
53 exhumation of highly weatherable mafic and ultramafic rocks (Reusch and Maasch,
54 1998; Jagoutz et al., 2016). Arc-continent collision associated with the Taconic orogeny
55 has been suggested to be associated with Ordovician cooling (Reusch and Maasch, 1998),
56 but paleogeographic reconstructions typically place the Taconic arc system outside of the
57 tropic weathering belt and within the arid subtropics (e.g., Mac Niocaill et al., 1997;
58 Domeier, 2016; Torsvik and Cocks, 2017). Modern evaporite belts and the paleolatitude
59 of evaporites constrain the arid subtropics to be persistently between latitudes of 15° and 35°
60 (Evans, 2006). Given that weathering rates are strongly dependent on temperature and
61 precipitation, and that weathering rates within basaltic watersheds in the tropics are
62 approximately an order of magnitude higher than those in mid-latitudes (Dessert et al.,
63 2003), such a subtropical position would likely preclude major CO₂ drawdown associated
64 with arc-continent collision (Jagoutz et al., 2016). Consequently, the reconstruction of the
65 paleolatitude of the orogeny is critical to the hypothesis that an increase in silicate
66 weatherability associated with the Taconic orogeny drove a portion of Ordovician
67 cooling. Did the Taconic arc-continent collision occur in the arid subtropics or in the wet
68 tropics?

69 **TECTONICS OF THE TACONIC OROGENY**

70 The Taconic orogeny encompasses Ordovician collisional and accretionary events
71 between volcanic arcs that formed within the Iapetus Ocean and the Appalachian margin
72 of Laurentia. The Taconic orogeny has been separated into three broad phases (van Staal
73 and Barr, 2012): Taconic 1 (495–488 Ma) includes local amphibolite-grade
74 metamorphism in the arc terranes; Taconic 2 (488–461 Ma) spans the collision of the
75 leading edge of the Taconic arc system with distended fragments and promontories of the
76 Laurentian margin and the initiation of north-directed subduction (Fig. 1); and Taconic 3
77 (461–445 Ma) comprises later arc accretion events. By ca. 465 Ma, amalgamated arc
78 terranes and fragments of the margin were thrust onto Laurentia, and delivered arc
79 detritus, including detrital chromite, into marginal basins (e.g., Hiscott, 1978; Macdonald
80 et al., 2017).

81

82 The colliding Taconic arc system extended west (paleocoordinates in Fig. 1) into
83 the southern Appalachians as far as Alabama (Hibbard, 2000), and east along the
84 Greenland margin to Ellesmere Island (Trettin, 1987). This elongate east-west exposure
85 of the arc system was all within a similar latitude band (Fig. 1).

86 **PALEOGEOGRAPHY**

87 Concerted efforts over decades of integrating geologic and paleomagnetic data
88 have led to an understanding that from the Cambrian into the Ordovician, Laurentia's
89 Appalachian margin was oriented east-west as the northern boundary of the Iapetus
90 Ocean (Mac Niocaill et al., 1997). Although paleogeographic models typically place this
91 margin south of 20°S in the relatively arid subtropics, this position in the Ordovician is

92 poorly constrained due to a lack of reliable paleomagnetic poles from cratonic Laurentia.
93 In the comprehensive apparent polar wander path compilation of Torsvik et al. (2012),
94 only two poles are included for the Ordovician: the St. George Group
95 and Table Head Group limestones of
96 Newfoundland. However, the Table Head Group limestones fail a conglomerate test
97 (Hodych, 1989). Therefore, their remanence, and the similar remanence of the underlying
98 St. George Group, must be the result of remagnetization. The Table Head Group rocks
99 pass a fold test, indicating that remagnetization occurred prior to Devonian folding.
100 Exclusion of these poles exacerbates an already large temporal gap between Laurentia
101 poles in the Torsvik et al. (2012) compilation, such that there are no robust poles from the
102 craton between the ca. 490 Ma Oneota Dolomite and the ca. 438 Ma Ringgold Gap poles
103 (Fig. 2). The paleolatitudes implied by Cambrian and Silurian poles for Laurentia's distal
104 margin (e.g., the New York and St. Lawrence promontories) are both in the subtropics
105 (Fig. 2), and extrapolation between these poles (such as a spline fit; Torsvik et al., 2012)
106 keeps Laurentia at a similar position through the Ordovician.

107 Given that there are no robust Ordovician paleomagnetic data from the Laurentian
108 craton, we take the approach of using paleomagnetic data from well-dated volcanic rocks
109 on the accreting terranes with magnetizations that are interpreted to be primary. Because
110 the Appalachian margin must have been at or equatorward of these terranes, these data
111 provide the best existing constraints on the Ordovician paleolatitude of Laurentia and
112 have been interpreted to indicate the presence of peri-Laurentian, intra-Iapetan, and peri-
113 Avalonian arc volcanism (Mac Niocaill et al., 1997). Open source reconstructions
114 developed in GPlates software (<https://www.gplates.org/>) for the

115 evolution of the Iapetus Ocean (Domeier, 2016; Torsvik and Cocks, 2017) provide an
116 excellent framework that can be modified with this approach.

117 In contrast to the Laurentian craton, eight robust Ordovician paleomagnetic data
118 sets have been reported from accreted Taconic arc terranes through extensive efforts of
119 the Rob Van der Voo research group at the University of
120
121 Michigan (USA) (see the GSA Data Repository¹). The interpretation of primary
122 remanence in these volcanic rocks is variably based on dual polarities, positive fold tests,
123 and interpretation of magnetic mineralogy. The oldest such locality is within the Notre
124 Dame arc of Newfoundland, where ca. 477 Ma mafic volcanics of the Moreton's Harbour
125 Group yielded a paleolatitude of $\sim 11^\circ\text{S}$
126 ($8^\circ\text{--}15^\circ\text{S}$ at 95% confidence) and were therefore interpreted to have formed in close
127 proximity to Laurentia (Johnson et al., 1991). Four paleomagnetic localities from ca.
128 470–465 Ma volcanic rocks of the Victoria arc of Newfoundland provide paleolatitude
129 constraints; the lowest latitude results are from the Lawrence Head volcanics,
130 which were at $\sim 12^\circ\text{S}$ ($2^\circ\text{--}24^\circ\text{S}$ at 95% confidence) (see
131 the Data Repository). Similar aged pillow lavas from arc terranes in Newfoundland (the
132 Annieopsquotch arcs in Fig. 2) give paleolatitudes of $\sim 30^\circ\text{S}$ that have been interpreted to
133 indicate that they formed some distance from the margin within the Iapetus Ocean (Van
134 der Voo et al., 1991). In New England (northeastern United States), ca. 467 Ma volcanics
135 of the Bronson Hill arc yield a paleolatitude of $\sim 20^\circ\text{S}$ ($12^\circ\text{--}29^\circ\text{S}$ at 95% confidence)
136 while younger ca. 458 Ma volcanics give paleolatitudes of $\sim 14^\circ\text{S}$ ($8^\circ\text{--}23^\circ\text{S}$ at 95%
137 confidence) and $\sim 11^\circ\text{S}$ ($6^\circ\text{--}16^\circ\text{S}$ at 95% confidence).

138 Although the Notre Dame arc was at a low latitude by ca. 475 Ma (Johnson et al.,
139 1991) when it collided with hyperextended fragments of the Laurentian margin
140 (Macdonald et al., 2014; van Staal and Barr, 2012), the Taconic seaway separated these
141 terranes from the Laurentian autochthon until they were exhumed ca. 465 Ma. While the
142 width of the Taconic seaway is unconstrained, the hyperextended margin of northeast
143 Australia, which extends >500 km from the craton, may be a modern analog. This ~500-
144 km-wide seaway closed between 475 and 465 Ma.

145 Paleolatitude constraints from ca. 470–465 Ma volcanics of the Taconic arc
146 terranes span ~20° of latitude, suggesting a distended arc system comparable to the
147 modern southwest Pacific arc system (Fig. 1; Mac Niocaill et al., 1997). Although the
148 precise latitudinal spread is difficult to resolve given uncertainty associated with
149 paleolatitude estimates, we interpret the spread of these latitudes to represent the leading
150 and trailing edges of the arc system (Fig. 1). This approach is a simplification; analogous
151 to the modern southwest Pacific, there were probably other active subduction zones.
152 Shortening during the Taconic and subsequent orogenies would have translated these
153 terranes inward toward the craton, further contributing to the interpretation that Laurentia
154 was equatorward of their paleolatitudes. Overall, the paleomagnetic database strongly
155 supports a revised reconstruction wherein the Appalachian Laurentian margin was
156 equatorward of 10°S at 465 Ma (Fig. 1).

157 **WEATHERING PROXY DATA**

158 Strontium and neodymium isotope data were compiled and recalculated (see the
159 Data Repository) using *The Geological Time Scale 2012* (see Cooper and Sadler, 2012).

160

161 $^{87}\text{Sr}/^{86}\text{Sr}$ data developed from the conodont apatite record a broad
162 decline from 0.7090 to 0.7088 between 480 and 465 Ma. This gradual decline is followed
163 by a sharp deflection at 465 Ma toward more juvenile $^{87}\text{Sr}/^{86}\text{Sr}$ values, reaching 0.7079
164 by 450 Ma (Saltzman et al., 2014; Fig. 2). Neodymium isotope (ϵ_{Nd}) data from fine-
165 grained siliciclastic rocks deposited on the distal margin of Laurentia, on the Taconic
166 allochthon, and Sevier basin (Gleason et al., 2002; Macdonald et al., 2017) display an
167 inflection to more positive values at 465 Ma consistent with a substantial increase of
168 sediment being weathered from juvenile lithologies (Fig. 2). This inflection in ϵ_{Nd} values
169 occurs later in more interior basins (Fig. 2) that did not receive arc-derived sediment until
170 subsequent accretionary events thrust arc rocks onto Laurentia between ca. 455 and 450
171 Ma (Macdonald et al., 2014, 2017).

172 **DISCUSSION**

173 The paleogeographic reconstruction presented here suggests that the Appalachian
174 margin was at a significantly lower latitude than is typically depicted, equatorward of
175 10°S by 465 Ma (Fig. 1). Our reconstruction is compatible with paleomagnetic data from
176 the Taconic arc system and is not in conflict with robust paleomagnetic poles from
177 Laurentia.

178 We propose that the broad rise in oxygen isotope values and decline in strontium
179 isotope values between 490 and 465 Ma (Fig. 2) are related to the movement of the
180 Taconic arc system into the tropics and collision of the leading edge with distended
181 fragments and promontories of the Laurentian margin (Taconic orogenic phase 2 of van
182 Staal and Barr, 2012). A concomitant increase in global weatherability would have
183 caused cooling through CO_2 drawdown, moderated by the silicate weathering feedback.

184 In addition, we argue that the sharp drop in $^{87}\text{Sr}/^{86}\text{Sr}$ values, the shift toward more
185 juvenile ϵ_{Nd} values in shale from the distal margin of Laurentia, and the additional
186 increase in oxygen isotope values between 465 and 455 Ma (Fig. 2) are due to the uplift
187 and exhumation of the Taconic arc system in the tropics (peak of Taconic 2) followed by
188 continued Late Ordovician arc accretion (Taconic 3). This exhumation led to uplift and
189 erosion of island arc volcanics and suprasubduction ophiolites, as evidenced by the
190 presence of detrital chromite in Middle to Late Ordovician foreland basins (e.g., Hiscott,
191 1978).

192 Increased weathering of volcanic arcs associated with the Taconic orogeny was
193 previously invoked to explain the Ordovician drop in $^{87}\text{Sr}/^{86}\text{Sr}$ values (Young et al.,
194 2009). The feasibility of this scenario was supported with a model in which global
195 weatherability was increased by 25% and a new flux of riverine $^{87}\text{Sr}/^{86}\text{Sr}$ was introduced
196 from weathering basalt with a composition of 0.7043 (Young et al., 2009). The ϵ_{Nd}
197 compilation from the Appalachian margin of Laurentia, which records local provenance,
198 is consistent with the hypothesis that the Taconic orogeny played a significant role in the
199 inferred increase in global weatherability and riverine $^{87}\text{Sr}/^{86}\text{Sr}$ input to the ocean. The
200 inflection in ϵ_{Nd} data from distal margin basins occurs a few million years prior to the
201 inflection in the global $^{87}\text{Sr}/^{86}\text{Sr}$ curve (Fig. 2). This lead time is predicted if the
202 weathering of Taconic terranes is a significant driver of the global strontium signal.
203 Juvenile ϵ_{Nd} values should be imparted in siliciclastic rocks over the time scale that
204 sediment transits from source to sink (thousand year time scales), whereas strontium has
205 a multimillion year residence time in the ocean such that a prolonged interval of arc
206 weathering would be necessary to significantly change seawater $^{87}\text{Sr}/^{86}\text{Sr}$. A complication

207 in this interpretation is that the age model for the ϵ_{Nd} data is anchored by U-Pb zircon
208 ages from ashes within the same stratigraphic sections (Macdonald et al., 2017), whereas
209 the $^{87}\text{Sr}/^{86}\text{Sr}$ age model is based on Cooper and Sadler (2012; Saltzman et al., 2014), so
210 the estimated temporal offset is as accurate as the calibration of the geological time scale.

211 Although other arc systems likely enhanced global weatherability in the
212 Ordovician, such as those in the paleo-Asian Ocean and the Fammetanian
213 arc of present-day Argentina, the Taconic arcs likely played an
214 outsized role as they were exhumed along an east-west belt in the tropics during the
215 closure of the Iapetus Ocean (Fig. 1). Exhumation would have created significant
216 topography composed of mafic and ultramafic lithologies through a wide swath across
217 the tropics. This scenario has similarities to the low-latitude closure of the Neo-Tethys
218 Ocean, and two-phase collision of the trans-Tethyan subduction system, which coincided
219 with the two-pronged cooling trend from the Cretaceous to Oligocene (Jagoutz et al.,
220 2016). The closure of major oceanic basins along east-west belts in the tropics may have
221 been a significant driver of long-term cooling trends throughout Earth history. Following
222 the Taconic orogeny, the Appalachian margin moved away from the tropics, so that
223 collisions associated with the Salinic orogeny in the Silurian would have occurred at
224 $\sim 20^\circ\text{S}$, where there would have been a lesser effect on global weatherability (Fig. 2).

225 Lower $p\text{CO}_2$ resulting from elevated global weatherability could have set the
226 stage for the growth of ice sheets during the Hirnantian. However, these tectonic
227 boundary conditions may not be the sole driver for the Hirnantian ice advance, and other
228 factors such as orbital forcing, changing ocean circulation, organic carbon burial, or rapid

229 changes in albedo may have caused the shorter term cooling associated with the
230 Hirnantian glacial maximum.

231 **CONCLUSIONS**

232 Our paleogeographic reconstruction constrained by the paleolatitude of
233 allochthonous volcanic rocks demonstrates that Laurentia moved toward the equator
234 during the Ordovician such that the Appalachian margin was equatorward of 10°S at 465
235 Ma. This movement into the tropics coincided with (1) collision and exhumation of the
236 Taconic arc system marked by the appearance of detrital chromite in foreland basins; (2)
237 a shift in ϵ_{Nd} data from fine-grained siliciclastic rocks on the Laurentian margin to more
238 juvenile values; (3) a drop in seawater $^{87}\text{Sr}/^{86}\text{Sr}$ values to more juvenile values; and (4) a
239 continued trend to higher values in the oxygen isotopic composition of both brachiopod
240 carbonate and conodont phosphate. These data are consistent with tropical weathering of
241 the Taconic arc-continent collision as a driver of Ordovician cooling.

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345 **FIGURE CAPTIONS**

346 Figure 1. Paleogeographic reconstruction ca. 465 Ma, after the arrival of the leading edge
347 of the Taconic arc system in the tropics along with the paleolatitude from allochthonous
348 volcanic rocks shown with 95% uncertainty. The reconstructed positions of these
349 paleomagnetic localities are shown on the classic position of Laurentia (as in Torsvik and
350 Cocks, 2017) and the new position proposed herein. While Laurentia most have been
351 north of these volcanics, in the classic reconstruction their positions
352 are south of the paleolatitudinal constraints
353 rather than equatorward, as in the revised position. The positions of other continental
354 blocks are as in Torsvik and Cocks (2017), other than Carolina, which is modified to be
355 traveling in unison with Ganderia.

356

357 Figure 2. Paleomagnetic and geochemical data from 500 to 400 Ma. A: Paleolatitude
358 constraints for Laurentia, Taconic arc terranes (Popelogan-Victoria, Bronson Hill,
359 Annieopsquotch, and Notre Dame), and the peri-Gondwana Ganderia and Avalonia
360 terranes. Laurentia paleolatitudes are calculated for two localities on the margin from
361 paleomagnetic poles with the implied position of New York (NY) shown for the classic

362 and new models. B: Strontium isotope data from conodont apatite and brachiopod calcite
363 with a locally weighted scatterplot smoothing (LOWESS) regression curve to the data of
364 Saltzman et al. (2014). C: Neodymium isotope data from fine-grained siliciclastic rocks
365 on the Appalachian margin of Laurentia with a LOWESS curve for distal margin data. D:
366 Oxygen isotope data from conodont apatite and brachiopod calcite with a LOWESS
367 curve for the brachiopod data. VPDB—Vienna Peedee belemnite; VSMOW—Vienna
368 standard mean ocean water. E: Orogenic phases wherein Taconic 2 spans the collision of
369 the leading edge of the arc system with promontories of the Laurentian margin. The peak
370 of Taconic 2 coincides with arc exhumation in the tropics and weathering of ophiolite and
371 arc detritus into Laurentian foreland basins. Late Ordovician arc accretion composes
372 Taconic 3. Data sources are provided in the Data Repository (see footnote 1).

373

374 ¹GSA Data Repository item 2017238, details of the paleomagnetic and
375 chemostratigraphic data compilations, is available online at
376 <http://www.geosociety.org/datarepository/2017/> or on request from
377 editing@geosociety.org.



