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# X-ray Bi-prism Interferometry – A Design Study of Proposed Novel Hardware

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Running Title: X-ray Bi-prism Interferometry

11 **Abstract**

12 **Purpose:** Advances in X-ray phase-contrast imaging can obtain excellent soft-tissue contrast of phase-shift,  
13 attenuation, and small angle scatter. Here we present fringe patterns for different design parameters of X-  
14 ray bi-prism interferometry imaging systems. Our aim is to develop bi-prism interferometry imaging  
15 systems with excellent polychromatic performance that produce high-contrast fringes with spatially  
16 incoherent X-ray illumination. We also introduce a novel X-ray tube design that uses temporal multiplexing  
17 to provide simultaneous virtual “electronic phase stepping” that replace “mechanical phase stepping”  
18 popular with grating-based interferometry setups.

19 **Methods:** In our investigation we develop expressions for the irradiance distribution pattern of a bi-prism  
20 interferometer comprised of multiple point sources and multiple bi-prisms. These expressions are used to  
21 plot fringe patterns for X-ray design parameters including size of point source, number of point sources,  
22 and point source separation; and bi-prism design parameters including material, angle, number of bi-  
23 prisms, period, and bi-prism array to X-ray source and to detector distances.

24 **Results:** Results show that the fringe patterns for a bi-prism interferometry system are not longitudinally  
25 periodic as with grating interferometers that produce a Talbot-Lau carpet. It is also shown that at 59 keV X-  
26 rays the bi-prism material should be something comparable to nickel to obtain reasonable fringe visibility.

27 **Conclusion:** The irradiance distribution pattern demonstrates that bi-prism interferometry may provide  
28 comparable or improved fringe visibility to that of gratings. Special care is given to present our findings  
29 within the context of previous advancements. A single-shot image acquisition approach using a temporal  
30 multiplexed, high-power X-ray source provides virtual electronic phase stepping without focal spot  
31 sweeping. This provides multiple images, each at the same exposure and the same projection view, from  
32 different fringe locations that allow one to derive the attenuation, phase, and dark-field images of the  
33 sample without mechanical phase stepping of a grating.

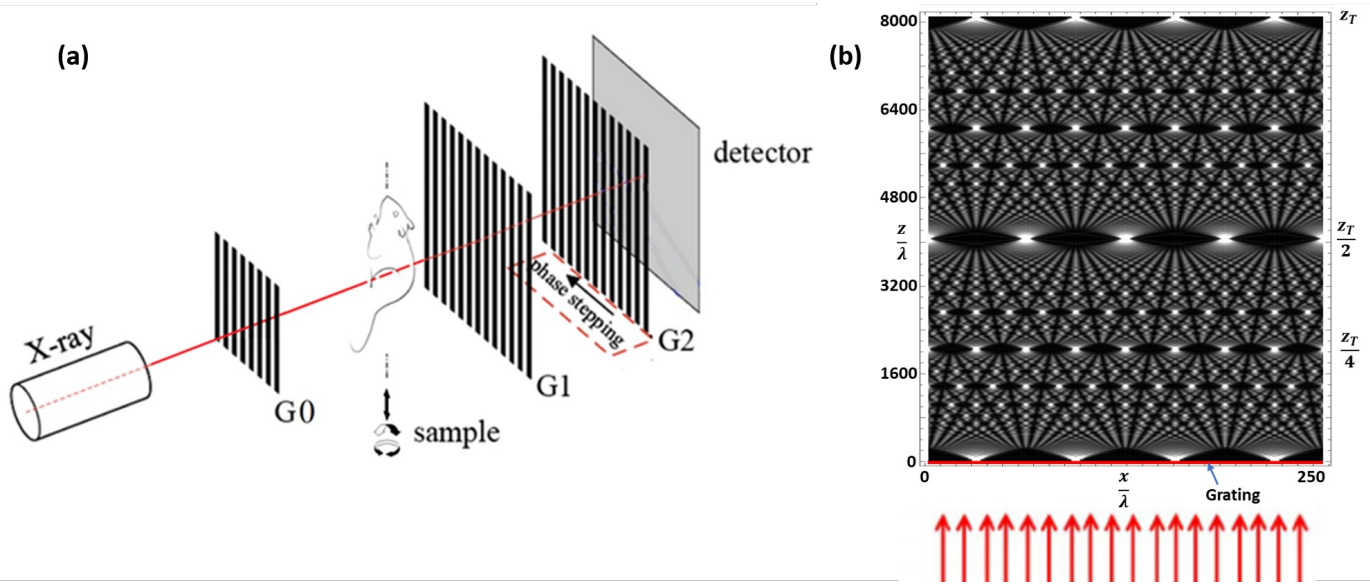
## 34 **1. INTRODUCTION**

35 To take advantage of the full potential of X-ray CT, there is a need for imaging systems that measure the  
36 full extent of the physics (attenuation, differential phase-shift, small angle scatter) involved in the  
37 interaction of X-rays with soft tissues. In a recent review it was demonstrated that interferometry-based X-  
38 ray imaging can provide excellent soft-tissue contrast of phase and directional small angle scatter  
39 information in addition, attenuation properties of tissues common in present day X-ray CT.<sup>1</sup> We propose to  
40 design such imaging systems using refractive bi-prisms<sup>2,3</sup> with optimum material and geometric parameters  
41 to provide fine spatial fringe modulation with high intensity, and to develop novel X-ray tubes that  
42 eliminate the need for mechanical phase stepping.

43

### 44 **1.A. Difference between grating and bi-prism interferometry**

45 There are a variety of X-ray phase contrast techniques all of which are based on the observation of  
46 interference patterns between deviated and un-deviated rays. Some approaches generate a Talbot-Lau  
47 interference pattern shown in Fig. 1 using a grating. This produces a replication of the grating via  
48 overlapping interference fringe patterns at regular distances away from the source. By placing the detector  
49 at the particular distance where the signal of the intensity is amplified and using methods to measure phase  
50 shifts in the interference fringes one is able to separate X-ray attenuation, phase, and scatter of the sample  
51 being imaged by comparing the observed pattern with and without the sample. In Fig. 1 for a plane wave  
52 passing through a grating, we see different fringe patterns mimicking the source images at regular distances  
53 from the source. We will show for the bi-prism interferometer that the amplified fringe patterns are not  
54 found at periodic longitudinal distances. Moreover, we will show that the visibility of many, but not all, of  
55 the fringe patterns is diminished by the increased numbers of bi-prisms.

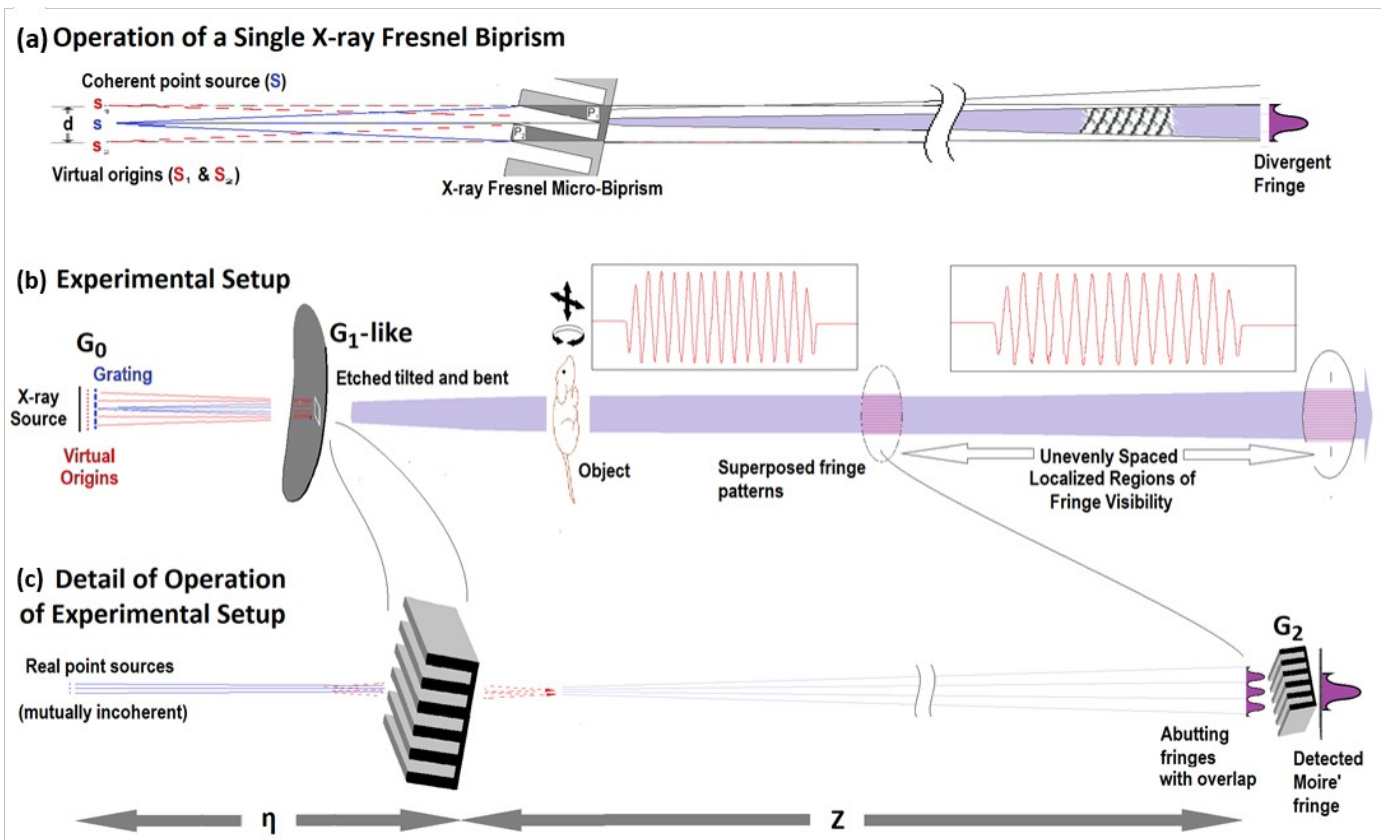


56  
 57 Fig. 1. (a) Schematic diagram of an X-ray grating interferometry system. (b) Talbot-Lau carpet.  
 58 Illuminating plane wave passes through a grating producing a fringe pattern with replicating amplified  
 59 fringe patterns at regular distances from the sources produced by the grating. At  $Z_T/2$  there is a secondary  
 60 Talbot image and at  $Z_T$  a replication of the original Talbot image that emerged from the grating. At  $Z_T/4$   
 61 there is a double frequency fractional image and increased frequency of images at less fractional distances.  
 62 (Modified from Wikipedia: [https://en.wikipedia.org/wiki/Talbot\\_effect](https://en.wikipedia.org/wiki/Talbot_effect).)

63  
 64 In our lab, we are investigating a novel X-ray interferometer (Fig. 2) wherein an array of Fresnel bi-  
 65 prisms produces interference fringes with X-rays from a source grating ( $G_0$ ).<sup>4</sup> The system does not have a  
 66 phase stepping grating ( $G_2$ ) common in most X-ray phase contrast imaging systems, but our intent is to use  
 67 hardware, such as a detector/scintillator with small hexagonal elements,<sup>5,6</sup> which acts as a  $G_2$  analyser  
 68 grating producing a Moiré pattern<sup>7</sup> on the detector. The source grating  $G_0$  forms multiple mutually-  
 69 incoherent sources of X-ray illumination from a single X-ray source. The multiple mutually-incoherent X-  
 70 rays sources refract through each element of a bi-prism, thus overlapping as if proceeding from two slightly  
 71 separated virtual sources of coherent rays (Fig. 3). By selecting bi-prism design parameters of material,  
 72 angle, number of bi-prisms, period, and bi-prism to X-ray source and to detector distances; the bi-prism  
 73 array interferometer provides spatially modulated intensity across a wide field as each center fringe thus  
 74 produced falls in a resonant position at the detector. An X-ray bi-prism material has widely varied  
 75 refracting power relative to wavelength and thus the separation of the virtual sources also varies with X-ray

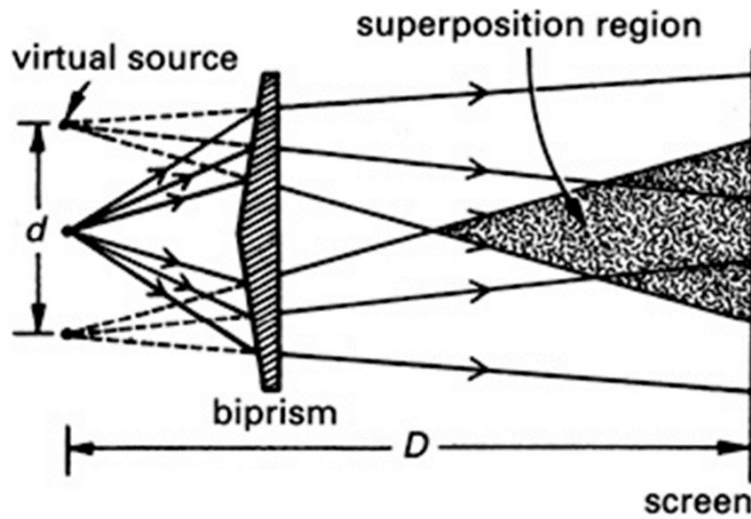
76 wavelength. The setup allows for rotating the sample itself to capture different angular views needed for  
 77 analysing the X-ray scattering resulting from the orientation of a sample's internal microstructure.

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Fig. 2. Schematic diagram of an X-ray bi-prism interferometry system. (a) Operation of a Fresnel X-ray bi-prism, (b) experimental setup with multiple X-ray origins each producing fringes in resonant position, and (c) detail of production of Moiré fringes via a tilted grating  $G_2$  at the detector. Our intention is to eliminate the  $G_2$  grating and instead use only a detector/scintillator with small hexagonal elements to produce the Moiré fringe pattern. (Modified from our earlier publication.<sup>4</sup>)



88

89 Fig. 3. X-ray beam deflection by a bi-prism illustrating the distance  $d$  between the two virtual coherent  
 90 point sources produced by a bi-prism. (Copied from The Fresnel Biprism, David-Alexander Robinson and  
 91 Jack Denning; Daniel Tanner 08332461; 10th December 2009.)

92 **1.B. Mathematical manipulation of a single exposure of an X-ray phase contrast interferometer**

93 To reduce the acquisition time and dose, mathematical data processing methods have been developed to  
 94 extract the data from a single exposure. One method<sup>8</sup> “reverse projection” processing is based on  
 95 manipulation of the phase relation between different X-ray projections and uses tomography to reconstruct  
 96 the absorption coefficient and refractive index by filtered backprojection of sampled projections. Another  
 97 method using a single lead grid,<sup>9,10</sup> relies on Fourier analysis of spatial harmonics to obtain absorption,  
 98 differential phase contrast, and diffraction images, all with a single exposure. The method assumes that  
 99 refraction and diffraction in the imaged object is manifested as position shifts and amplitude attenuation of  
 100 the fringes, respectively. In an additional method,<sup>11</sup> the scan of the phase-stepping grating is combined with  
 101 the rotation of the sample so that a single detector frame per projection angle is obtained, yielding images  
 102 corresponding to all three contrast modalities. Most single shot methods produce Moiré fringe patterns and  
 103 depends on Fourier analysis to extract the attenuation, phase, and small angle scatter.<sup>12-17</sup> In particular one  
 104 paper<sup>13</sup> used continuous wavelet transforms to extract the phase information from Moiré interferograms. In  
 105 another paper,<sup>16</sup> Moiré fringe patterns are produced by positioning a two-dimensional checkerboard grating  
 106 at the first Talbot position beyond the object being imaged. Differential phase-contrast and absorption

107 images are obtained by Fourier analysis of the Moiré fringe patterns. In two other papers,<sup>15,17</sup> in addition to  
108 the source grating ( $G_0$ ), the phase grating ( $G_1$ ) and the analyzer grating ( $G_2$ ) are designed and manipulated  
109 to produce a Moiré pattern. In the future, our goal is to develop an imaging system that produces a Moiré  
110 pattern without a  $G_2$  grating.

111

#### 112 **1. D. Applications of bi-prism interferometry**

113 Our bi-prism setup is the first to explore X-ray phase contrast bi-prism interferometry in a laboratory  
114 setup. The bi-prism is a different approach from gratings to produce refracted waves for extracting the  
115 properties of X-rays. Its fringe pattern is divergent in most setups. Bi-prisms have had many applications in  
116 the optical regime.<sup>2</sup> Most of the investigations in X-ray bi-prism applications has been using hard X-rays  
117 generated in synchrotrons.<sup>3,18-21</sup> However, it has also found application to illustrate the wave-particle  
118 behaviour in the single-photon regime.<sup>22,23</sup> Bi-prisms have also been used in interference electron  
119 microscopy.<sup>24</sup>

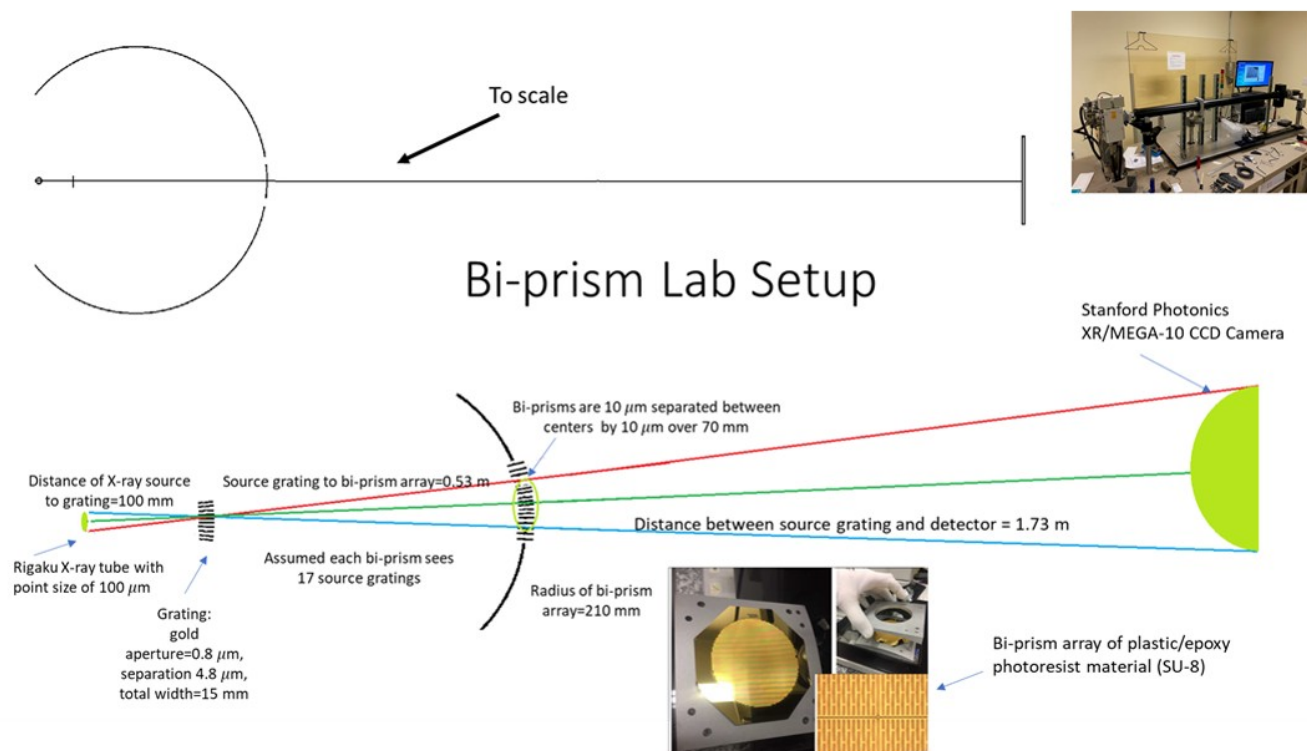
120 **Outline of paper.** In the following we develop analytical expressions for the irradiance distribution of an  
121 array of bi-prisms. Using the expressions, we demonstrate the non-periodic pattern characteristic of a bi-  
122 prism interferometry system. We also evaluate results of various design parameters for a bi-prisms  
123 interferometry system. We provide results for 17.5 keV X-rays with silicon and a plastic/epoxy photoresist  
124 material (SU-8) demonstrating irradiance distributions for different number of point sources, point source  
125 separations, and bi-prism angles. We then investigate the possibility of a nickel bi-prism array for 59 keV  
126 that one might consider in the design of a high energy X-ray CT scanner. The results are followed by a  
127 discussion of the potential merits of bi-prism interferometry.



## 128 2. MATERIALS AND METHODS

### 129 2.A. X-ray bi-prism interferometry system

130 In this paper, we propose a modified X-ray interferometer (Fig. 2) wherein the phase grating ( $G_1$ ) is  
131 replaced with an array of Fresnel micro-bi-prisms. A novel design for the source grating ( $G_0$ ) is also  
132 provided and the analyzer grating ( $G_2$ ) might be eliminated, though in our present analysis we use a rotated  
133 grating ( $G_2$ ) to produce a Moiré pattern. The Fresnel bi-prism produces high-contrast fringes with spatially  
134 incoherent X-ray illumination. The bi-prism array is composed of 7,000 sets of two counter-positioned  
135 refractive prisms illuminated with multiple sources. X-rays refracting through each prism element overlap  
136 and form divergent (magnified) interference fringes on a distant detector. A further advantage when using  
137 polychromatic radiation is that the central fringe produced by a bi-prism is always a “white light” fringe.<sup>25</sup>



138

139 Fig. 4. Lab setup of the bi-prism interferometry system at UCSF. The bi-prism interferometry system  
140 includes a molybdenum-target X-ray tube operating at 35 kVp with predominate energy of the 17.5 keV  
141 characteristic line (designed energy/wavelength). Shown also is a gold source grating with 0.8 μm apertures  
142 separated by 4.8 μm, a bi-prism array made of plastic/epoxy photoresist material (SU-8), and a Stanford  
143 Photonics iCCD camera.

144

145 Our laboratory bi-prism interferometry system (Fig.4) includes a Rigaku X-ray tube (Rigaku Corp.,  
 146 Akishima-shi, Tokyo, Japan) operating at 17.5 keV with a point-size of  $100 \mu\text{m}$  and a 15 mm wide gold  
 147 source grating with  $0.8 \mu\text{m}$  apertures separated by  $4.8 \mu\text{m}$ . The bi-prism array consisting of multiple  $10 \mu\text{m}$   
 148 bi-prisms with centers separated by  $10 \mu\text{m}$  is curved with radius of 210 mm (We assumed flat in our  
 149 numerical calculations.), is located 0.53 m from the source grating, and is a plastic/epoxy photoresist  
 150 material (SU-8) with refraction index decrement of  $\delta=8.7 \times 10^{-7}$ . A gold source grating with  $0.8 \mu\text{m}$   
 151 apertures separated by  $4.8 \mu\text{m}$  provides multiple point sources to the bi-prisms of coherent sources of X-  
 152 rays, which are each mutually incoherent. The X-ray tube source size determines the number of point  
 153 source apertures that illuminate a single bi-prism. We assume that each point source aperture will radiate 17  
 154 separate bi-prisms. The detector is a Stanford Photonics CCD camera and is located 1.73 m from the source  
 155 grating.

156 Our work is aimed at imaging small objects at energies of 17.5 keV; however, our X-ray source to detector  
 157 distance is appropriate for a human size X-ray CT scanner. The equations developed in the following section  
 158 provide parameters that one can change to design a bi-prism interferometry system at higher X-ray energies.

159

## 160 **2.B. Analytical expression for the irradiance distribution pattern**

161 In the following section, we develop analytical equations for the irradiance distribution pattern for our bi-  
 162 prism interferometry system. We use these equations to plot the distribution of fringes for different  
 163 parameters of a bi-prism interferometry system.

164 Assuming an X-ray beam proceeding from a spatially incoherent planar source with wavelength  $\lambda$   
 165 illuminates a thin bi-prism (Fig. 3), we write the irradiance distribution  $I$  of X-ray sources on an arbitrary  
 166 plane placed at a distance  $z$  beyond the bi-prism as<sup>2</sup>

$$167 I(\vec{x}, z; \eta) = \frac{\eta^2}{z^2} I_s\left(\frac{-\eta \vec{x}}{z}\right) \otimes_2 I_0(\vec{x}, z; \eta), \quad (1)$$

168 where  $\eta$  is the distance between the source and the bi-prism,  $z$  is the distance between the bi-prism and the  
 169 imaging plane,  $\otimes_2$  is the 2D convolution performed over the transverse coordinates  $\vec{x}=(x, y)$ ,  $I_S$  is the  
 170 irradiance distribution of the planar source,  $I_0$  is the irradiance distribution of the Fresnel bi-prism, and  $I$  is  
 171 the irradiance distribution at the observation plane.

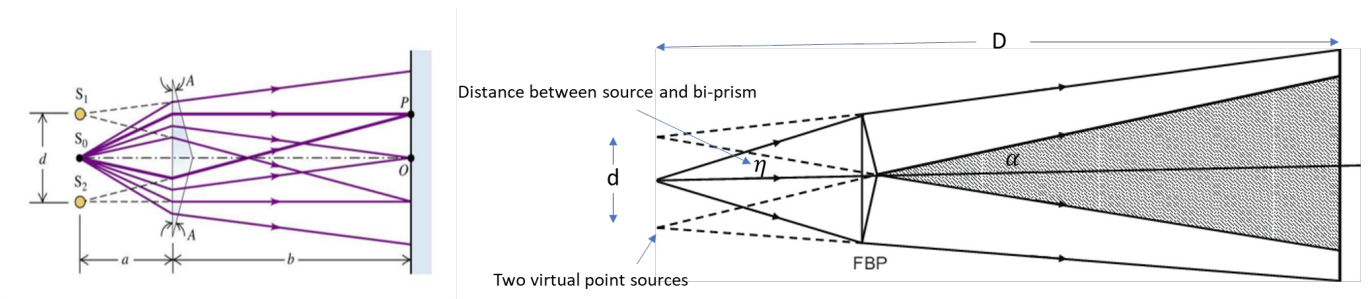
172 For a spherical wavefront proceeding from a point source illuminating a Fresnel bi-prism, the exiting  
 173 wavefront produces an interference pattern  $I_0$  beyond the Fresnel bi-prism whose irradiance distribution is

174 
$$I_0(\vec{x}, z; \eta) = 1 + \cos\left(\frac{2\pi x}{p}\right),$$

175 where  $p = \lambda(\eta + z) / (2\eta \tan \alpha)$  is the period of the interference pattern,  $\alpha = \delta \tan \chi$  is the angle of the beam  
 176 deflection,<sup>3</sup>  $\delta$  is the refraction index decrement, and  $\chi$  is the angle of the bi-prism shown in Fig. 5. Here the  
 177 expression for the period of the interference pattern for X-rays differ from the expression for light:<sup>2</sup>

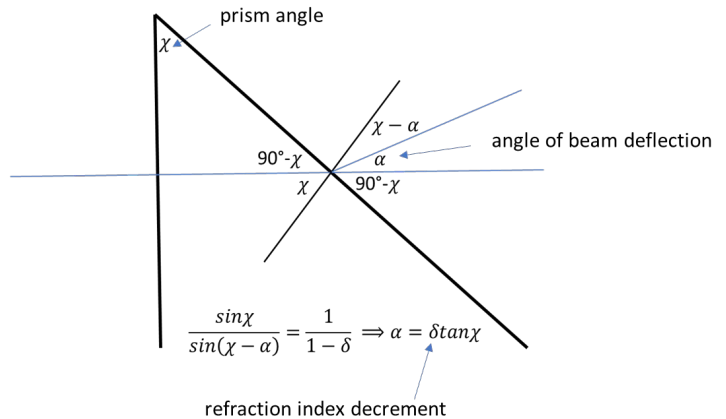
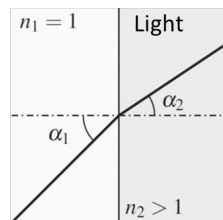
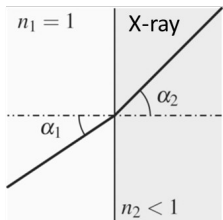
178  $p = \lambda(\eta + z) / (2(n - 1)\eta\theta)$ , where  $n$  is the index of refraction and  $\theta$  is the refringence angle of the bi-prism.

179 Note that light rays refract toward the normal to the material interface and X-rays refract away from the  
 180 normal (Fig. 5).



distance between two virtual point sources:  $d = 2\eta \tan \alpha$

angle of beam deflection:  $\alpha = \delta \tan \chi$



181

182 Fig. 5. X-ray beam deflection by bi-prism. Here is given the expression for the angle  $\alpha$  of the beam  
 183 deflection and the distance  $d$  between the two virtual coherent point sources produced by a bi-prism.  $\delta$  is  
 184 the refraction index decrement. Notice the difference in refraction between X-rays and light. (Modified from  
 185 Chegg Study, www.x-ray-optics.de, and researchgate.net)

186 Of special interest is a source composed by an array of  $N_s$  mutually incoherent point sources with the  
 187 same irradiance,  $I_p$ , and arranged equidistant perpendicular to the bi-prism edge and distributed  
 188 symmetrically to the optical axis. In a realistic experimental situation, the width  $\Delta$  of the sources along the  
 189  $x$ - and  $y$ -direction would not be infinitesimal. In that case, the irradiance distribution of each source can be  
 190 written as the convolution between a delta function and a rectangular function of width  $\Delta$ :

$$191 \quad h(x, y) = I_p \delta(x, y) * \text{rect} \left( \frac{x}{\Delta}, \frac{y}{\Delta} \right).$$

192 (Note: Here  $\text{rect}$  refers to the 2D convolution.) If the point sources are located with  $x$ -coordinates

$$193 \quad x_i = \left( \frac{N_s + 1}{2} - i \right) x_0, \quad i = 1, \dots, N_s,$$

194 where  $x_0$  is the separation between neighbor point sources, then from Eq. (1) the irradiance distributions of  
 195 each point source is

$$196 \quad g(x, y, z; \eta, x_i) = h \left( \frac{-\eta x}{z} - x_i, \frac{-\eta y}{z} \right) * \text{rect} \left( \frac{-\eta x}{z} - x_i, \frac{-\eta y}{z} \right) * \left[ 1 + \cos \left( \frac{2\pi x}{p} \right) \right],$$

197 where  $-z/\eta$  is the magnification factor between the source and observation plane. The irradiance  
 198 distribution of all the sources is therefore the summation of the distribution of source irradiances

$$199 \quad I(\vec{x}, z; \eta) = \frac{\eta^2}{z^2} \sum_{i=1}^{N_s} g(x, y, z; \eta, x_i).$$

200 Using Mathematica (Wolfram Research, Champaign, Illinois), we come up with the following analytical  
 201 expression for the irradiance distribution pattern for  $N_b = 1$  (one bi-prism):<sup>4</sup>

$$202 \quad I_{N_b=1}(x, z; \eta) = \begin{cases} I_p \left( \Delta^2 + \frac{p \eta \Delta}{\pi z} \cos \left[ \frac{2\pi x}{p} \right] \sin \left[ \frac{\pi z \Delta}{p \eta} \right] \right) & \text{if } N_s = 1 \\ \text{rect} \left( \frac{x}{\Delta}, \frac{y}{\Delta} \right) * I_p \left( N_s \Delta^2 + \frac{p \eta \Delta}{\pi z} \cos \left[ \frac{2\pi x}{p} \right] \sin \left[ \frac{\pi z \Delta}{p \eta} \right] \csc \left[ \frac{\pi x_0 z}{p \eta} \right] \sin \left[ \frac{\pi N_s x_0 z}{p \eta} \right] \right) & \text{if } N_s > 1 \end{cases}, \quad (2)$$

203 where

204

$$p = \lambda(\eta + z) / (2\eta \tan \alpha) = \lambda(\eta + z) / [2\eta \tan(\delta \tan \chi)]$$

205

is the period of the interference pattern. Notice that the expression is given only as a function of the  $x$ -

206

coordinate, assuming the irradiation distribution in  $y$  is minimal.

207

Equation (2) gives an expression for one bi-prism with one and multiple point sources. To obtain an

208

expression for multiple bi-prisms, we used Mathematica to sum the irradiance distribution of  $N_b$  shifted bi-

209

prisms separated by  $x_b$  using the expression

210

$$I(x, z; \eta) = \sum_{j=1}^{N_b} I_{N_b=1} \left[ \left( (N_b + 1) / 2 - j \right) x_b \right] ,$$

211

to obtain the following expression for multiple bi-prisms and multiple sources:

212

$$I(x, z; \eta) = I_p \left[ N_b N_s \Delta^2 + \frac{p \eta \Delta}{\pi z} \cos \left( \frac{2 \pi x}{p} \right) \csc \left( \frac{\pi x_0 z}{p \eta} \right) \csc \left( \frac{\pi x_b}{p} \right) \sin \left( \frac{N_s \pi x_0 z}{p \eta} \right) \sin \left( \frac{N_b \pi x_b}{p} \right) \sin \left( \frac{\pi z \Delta}{p \eta} \right) \right].$$

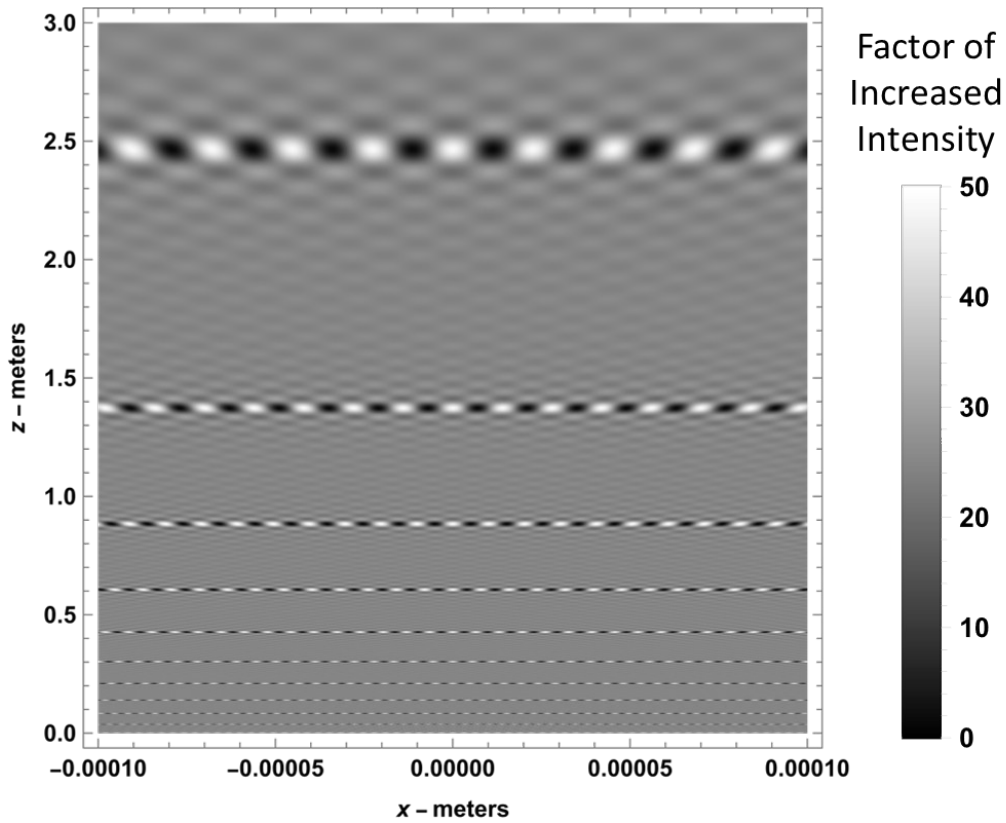
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(3)

214

215 **3. RESULTS**

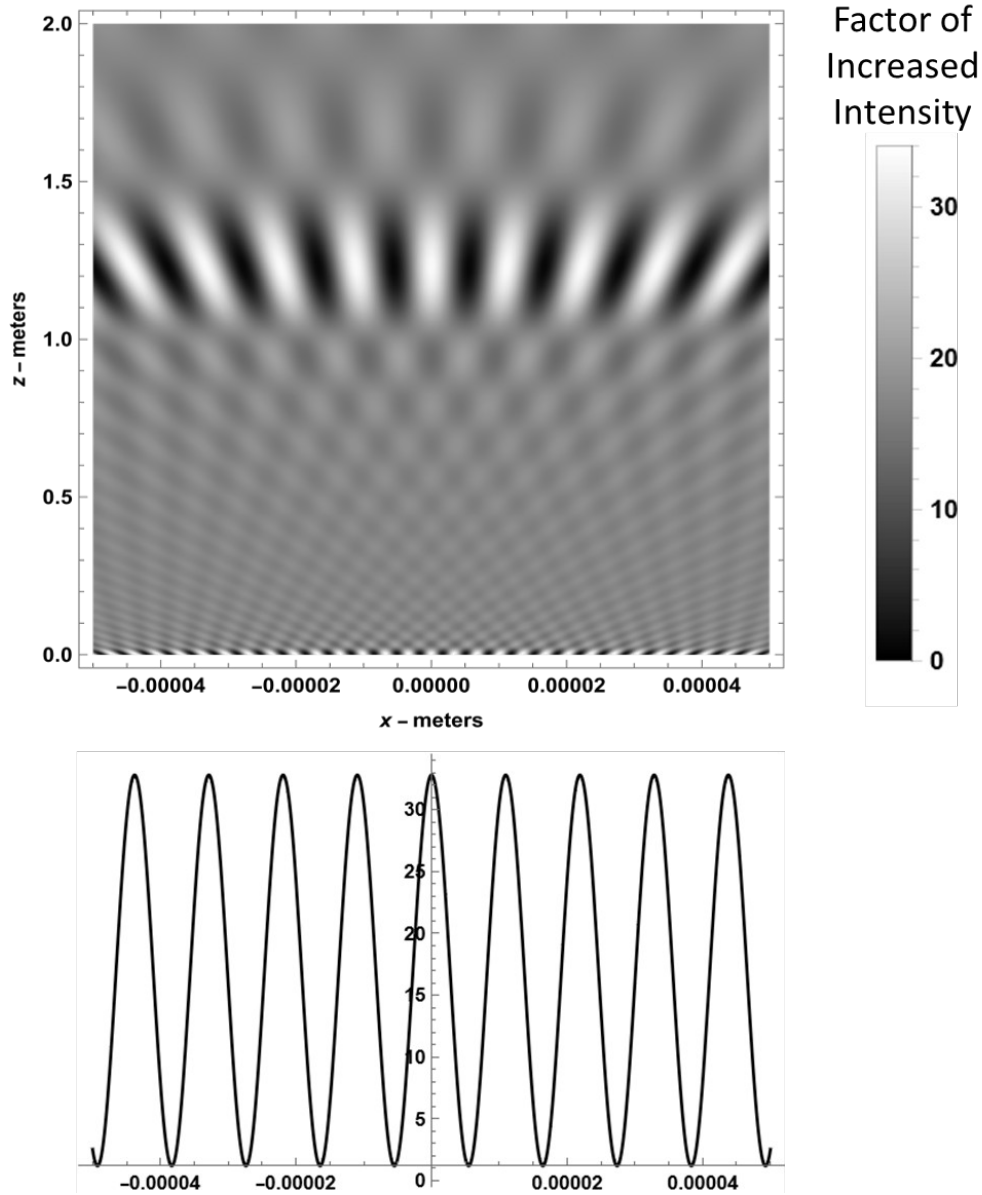
216 Mathematica was used to numerically evaluate Eq. (2) to obtain fringe patterns behind a bi-prism for  
 217 incoherent X-ray point sources located at  $\eta=0.4$  m in front of the bi-prism (see Fig. 6). For the calculation  
 218 we set  $\lambda=7.1\times 10^{-11}$  m,  $I_p=1/\Delta^2$ ,  $\Delta=7.00\times 10^{-7}$ , and  $\alpha=\delta \tan(\chi)$ , where  $\delta=1.57\times 10^{-6}$ . Using these  
 219 parameters, calculations were made for 25-point sources with a bi-prism angle of  $\chi=82^\circ$ . The scale  
 220 increases in proportion to the number of point sources.<sup>4</sup> Non-periodic visibility is seen in Fig. 6. This is  
 221 characteristic of a fringe pattern for a bi-prism and differs from the periodic pattern of the well-known  
 222 Talbot–Lau interferometer with gratings in Fig. 1.



223 Fig. 6. Density plot of a fringe pattern for 25-point sources with 1 bi-prism. The amplification of the  
 224 interference pattern is repeated at non-periodic distances away from the plane of the bi-prism. For the  
 225 calculation we set  $\lambda=7.1\times 10^{-11}$  m (17.5 keV),  $I_p=1/\Delta^2$ ,  $\Delta=7.00\times 10^{-7}$  m,  $\alpha=\delta \tan(\chi)$ ,  $\delta=1.58\times 10^{-6}$   
 226 (silicon),  $\chi=82^\circ$ ,  $\eta=0.4$  m, and  $x_0=36.7\ \mu\text{m}$ . This figure is modified with new parameters from that of our  
 227 previous publication.<sup>4</sup>  
 228  
 229

230 Notice that in our calculations using Eq. (2) we set the irradiance  $I_p$  of each point source to  $1/\Delta^2$ . One  
231 might want to consider this to be a flux of 1 photon per  $\Delta^2$ . Then the resultant intensity in Eq. 2 is a total  
232 number of photons. However, since we do not know the photon flux of the source, we interpret the result of  
233 Eq. (2) as a relative measure or factor of increase in intensity. Keep in mind that our equations represent  
234 refraction and do not include attenuation.

235 Mathematica was also used to numerically evaluate Eq. (3) to obtain a density plot in Fig. 7 of a fringe  
236 pattern behind one bi-prism with 17-point sources for set of parameters used in our laboratory setup that  
237 differed from those in Fig. 6.



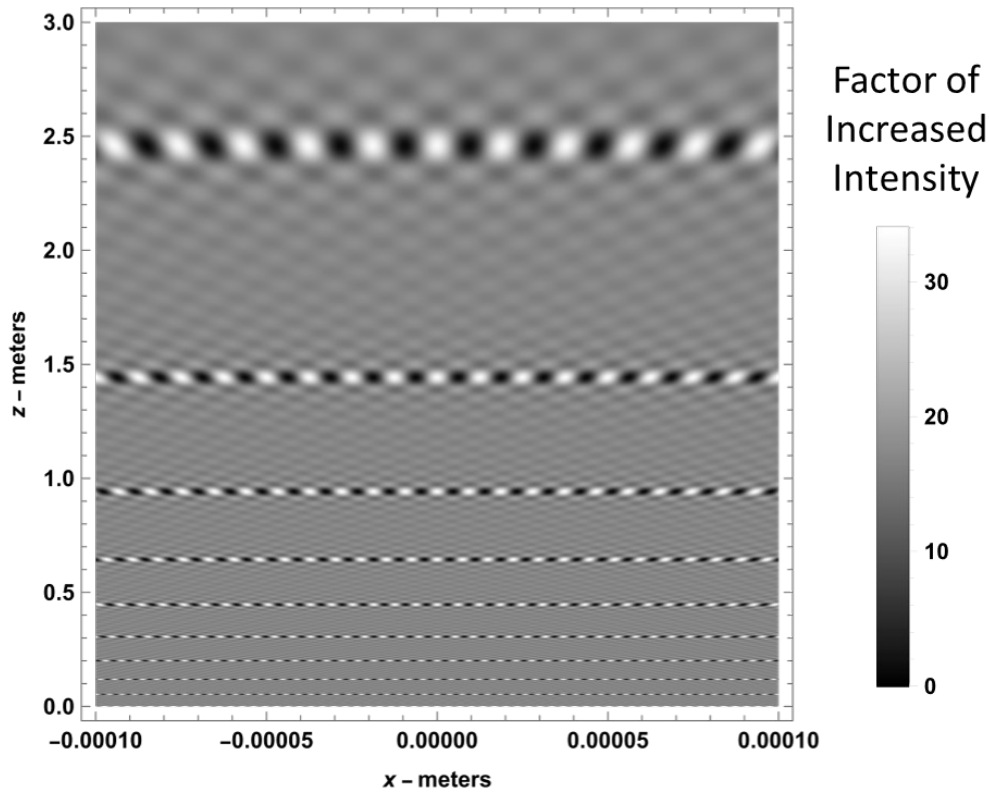
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240 Fig. 7. Density plot of a fringe pattern for 17-point sources with 1 bi-prism for a different set of  
 241 parameters than those in Fig. 6. For the calculation we set  $\lambda=7.1 \times 10^{-11}$  m (17.5 keV),  $I_p=1/\Delta^2$ ,  
 242  $\Delta=8.00 \times 10^{-7}$  m,  $\alpha=\delta \tan(\chi)$ ,  $\delta=8.7 \times 10^{-7}$  [plastic/epoxy photoresist material (SU-8)],  $\chi=85.3^\circ$ ,  
 243  $\eta=0.53$  m, and  $x_0=4.8 \mu\text{m}$ ,  $x_b=10 \mu\text{m}$ . There is a fringe separation of  $1.09 \times 10^{-5}$  m at  $z=1.2$  m.

244 If we change the source separation in Fig. 7 from  $x_0=4.8 \mu\text{m}$  to  $x_0=36.7 \mu\text{m}$ , we see a very different  
 245 fringe pattern in Fig. 8. This demonstrates that the fringe pattern is sensitive to the distance  $x_0$  between the  
 246 point sources.

247





248

249 Fig. 8. Density plot of a fringe pattern for 17-point sources with 1 bi-prism for the same set of parameters  
 250 as those in Fig. 7, except  $x_0$  was changed from  $x_0=4.8 \mu\text{m}$  to  $x_0=36.7 \mu\text{m}$ . For the calculation:  
 251  $\lambda=7.1 \times 10^{-11}\text{m}$  (17.5 keV),  $I_p=1/\Delta^2$ ,  $\Delta=8.00 \times 10^{-7}$  m,  $a=\delta \tan(\chi)$ ,  $\delta=8.7 \times 10^{-7}$  [plastic/epoxy  
 252 photoresist material (SU-8)],  $\chi=85.3^\circ$ ,  $\eta=0.53$  m, and  $x_0=36.7 \mu\text{m}$ ,  $x_b=10 \mu\text{m}$ .  
 253

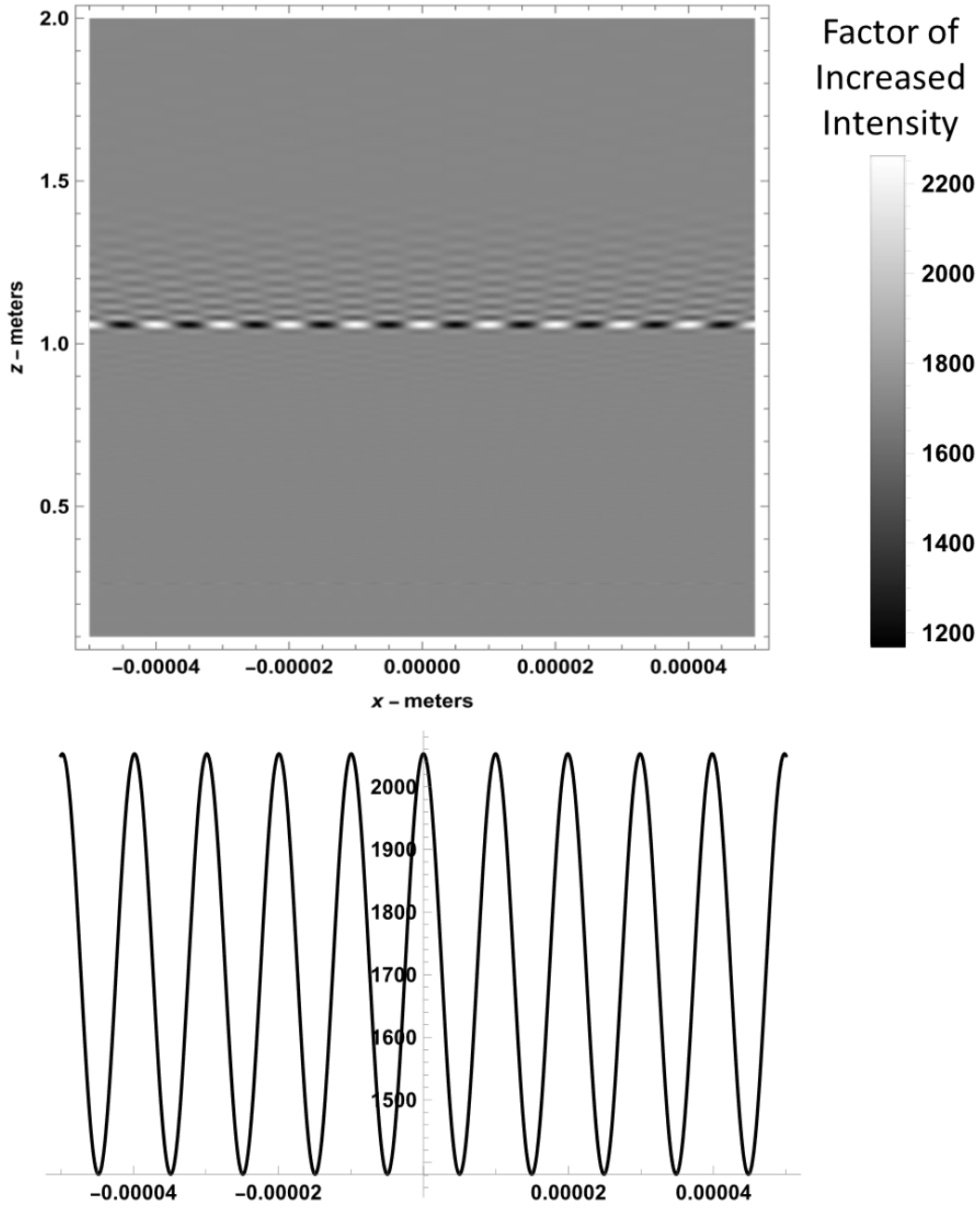
254 The fringe pattern is sensitive to the bi-prism material, the bi-prism angle  $\chi$ , and the distance  $x_0$  between  
 255 the point sources. In Fig. 6 the fringe pattern corresponds to silicon ( $\delta=1.58 \times 10^{-6}$ ) with parameters  
 256  $\chi=82^\circ$ ,  $\eta=0.4$  m,  $x_0=36.7 \mu\text{m}$ , and  $\Delta=7.00 \times 10^{-7}$  m; whereas, in Fig. 7 the fringe pattern corresponds to  
 257 SU-8 material ( $\delta=8.7 \times 10^{-7}$ ) used in our laboratory setup with parameters  $\chi=85.3^\circ$ ,  $\eta=0.53$  m,  $x_0=4.8 \mu$   
 258 m, and  $\Delta=8.00 \times 10^{-7}$ . Since the refractive index decrement of SU-8 is less than silicon, thus decreasing  
 259 the angle of deflection ( $a=\delta \tan \chi$ ); in our design we increased the angle of each bi-prism to  $\chi=85.3^\circ$ .  
 260 Probably the significant increase in the number of non-periodic fringes in Fig. 8 is due to changing the  
 261 source separation from  $x_0=4.8 \mu\text{m}$  to  $x_0=36.7 \mu\text{m}$ .

262 The visibility,  $(I_{max}-I_{min})/(I_{max}+I_{min})$ , was calculated for a few examples by finding the minimum and  
 263 maximum intensity for  $\pm 5 \mu\text{m}$  in  $x$  about the maximum central fringe. If we use 17-point sources with one

264 bi-prism, as in Fig. 8, we obtain a visibility of 0.93. If the point source separation is change from 36.7  $\mu\text{m}$   
265 to 4.8  $\mu\text{m}$  the visibility is improved to 0.95. If for the same point source separation of 36.7  $\mu\text{m}$ , the point  
266 source size  $\Delta$  is changed from  $8 \times 10^{-7}$  to  $4 \times 10^{-7}$ , the visibility is improved even more to 0.98. These trends  
267 in the visibility calculations are as one would expect; however, are higher than one would expect in a  
268 laboratory X-ray phase contrast interferometer. The work of Isakovic et al.<sup>19</sup> provided bi-prism  
269 measurements performed at the Brookhaven National Laboratory synchrotron where a single bi-prism was  
270 evaluated in an X-ray beam of 9.5 keV. Their data showed a difference in fringe visibility in the vertical  
271 and the horizontal axis, with a maximum fringe visibility in the vertical above 0.9 and approximately 0.6 in  
272 the horizontal axis. From this, they calculated an effective asymmetric source of their beamline as a vertical  
273 source of  $10 \pm 5 \mu\text{m}$  and a horizontal source of  $130 \pm 20 \mu\text{m}$ . From generally known parameters, the  
274 transverse coherence length of that beamline would be  $\sim 160 \mu\text{m}$  in the vertical and  $\sim 12.3 \mu\text{m}$  in the  
275 horizontal. In the simulation of our laboratory setup, we use an array of 0.8 -  $\mu\text{m}$  sources at 0.53-meters in  
276 front of an array of bi-prisms. At 17.5 KeV, the transverse coherence length of the operable axis is 11.8 -  
277  $\mu\text{m}$  and one would expect to produce one central fringe per each source/bi-prism pair. Of course, our  
278 proposed setup is an array of bi-prisms with a source grating as an array of small sources; thus, the detector  
279 records many such fringes each falling in resonant positions.

280 Again, using Mathematica to numerically evaluate Eq. (3), we obtained a fringe pattern behind 101 bi-  
281 prisms for 17-point sources. Notice in Fig. 7 that for a single bi-prism the maximum intensity is at  
282 approximately  $z = 1.2 \text{ m}$ , whereas, in Fig. 9 for 101 bi-prisms, the maximum intensity appears to be at  $z = 1.05 \text{ m}$ .  
283 Increasing the number of bi-prisms increases the fringe intensity and shifts the maximum intensity  
284 toward the sources while bringing the maximum fringe intensity more in focus and greatly reducing fringe  
285 intensity in other regions.

286



287

288 Fig. 9. Density plot of a fringe pattern for 17-point sources with 101 bi-prism for the same set of  
 289 parameters as in Fig. 7. For the calculation we set  $\lambda=7.1 \times 10^{-11}$  m (17.5 keV),  $I_p=1/\Delta^2$ ,  $\Delta=8.00 \times 10^{-7}$  m,  
 290  $\alpha=\delta \tan(\chi)$ ,  $\delta=8.7 \times 10^{-7}$  [plastic/epoxy photoresist material (SU-8)],  $\chi=85.3^\circ$ ,  $\eta=0.53$  m, and  $x_0=4.8 \mu$   
 291 m,  $x_b=10 \mu$  m. There is a fringe separation of  $1.00 \times 10^{-5}$  m at  $z=1.05$  m.  
 292  
 293

294 The equations developed in this section provide plots of the non-periodic fringe pattern, the intensity of  
 295 the maximum fringe intensity, and the fringe separation at the maximum fringe intensity for our laboratory  
 296 bi-prism design. Equation (3) was used to compare different possible designs of a bi-prism interferometry

297 system for parameters given in Table 1. Designs were compared between a bi-prism material of our  
298 present lab setup of plastic/epoxy photoresist material (SU-8) and silicon at 17.5 keV. The refractive index  
299 decrement for SU-8 was not listed in the CXRO online calculator,<sup>27</sup> but is a similar plastic to PMMA. The  
300 paper of Mamyrbayev et al.<sup>26</sup> states on page 7 that the refractive index of SU-8 is  $9.23 \times 10^{-7}$  at 17 keV,  
301 which is the value CXRO calculates for PMMA at 17 keV. The evaluations in our paper were performed  
302 for 17.5 keV, thus we used the value of  $\delta = 8.7 \times 10^{-7}$  for SU-8 at 17.5 keV which is the value for PMMA  
303 at 17.5 keV in the CXRO online calculator.<sup>27</sup>

304 By selecting different bi-prism materials, one can design a system that would be appropriate for an X-ray  
305 CT system with higher X-ray source energy than the 17.5 keV used here in our validations. In our  
306 demonstrations here, we have attempted to keep the distance between the X-ray source and detector with  
307 reasonable real world X-ray imaging applications. By modifying the bi-prism material and angle of the bi-  
308 prism one maintains reasonable X-ray source to detector distances. Silicon and our laboratory SU-8  
309 material are not ideal for a higher energy of 59 keV. A better material would be nickel. A value of  
310  $\delta = 5.06 \times 10^{-7}$  was obtained for nickel at 59 keV from the X-Ray Optics Calculator.<sup>28</sup> As one sees from  
311 Table 1, the fringe separation decreased by a factor of approximately 3. The angle of the bi-prism is also a  
312 function of the material with silicon requiring a smaller angle to bring the detector distance close to that of  
313 our SU-8 material, whereas in Table 1 we see that an increased angle of  $87^\circ$  is required for nickel.

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320 Table 1. Equation (3) was used to calculate the fringe separation and the fringe intensity for our laboratory  
321 bi-prism SU-8 material and compared to silicon at 17.5 keV. The fringe separation was nearly  $10 \mu m$  for  
322 our laboratory setup and would be the same for the designated parameters for a bi-prism material of silicon.

323 The same equation was used to design a bi-prism nickel interferometry system at 59 keV. The distances  
 324 between the bi-prisms and the selected detector positions were determined by observing the density plots  
 325 like the one in Fig. 9.

326  
 327

	Design 1 (Fig. 6)	Design 2	Design 3	Design 4 (Fig. 9)	Design 5
Bi-prism material	Silicon	plastic/epoxy photoresist (SU-8)	Silicon	plastic/epoxy photoresist (SU-8)	Nickel
Energy	17.5 keV	17.5 keV	17.5 keV	17.5 keV	59 keV
Wave length $\lambda$	$7.1 \times 10^{-11}$ m	$7.1 \times 10^{-11}$ m	$7.1 \times 10^{-11}$ m	$7.1 \times 10^{-11}$ m	$2.14 \times 10^{-11}$ m
Refractive index decrement $\delta$	$1.58 \times 10^{-6}$ *	$8.7 \times 10^{-7}$ **	$1.58 \times 10^{-6}$ *	$8.7 \times 10^{-7}$ *	$5.06 \times 10^{-7}$ +
Bi-prism angle $\chi$	$82^\circ$	$85^\circ$	$82^\circ$	$85.3^\circ$	$87^\circ$
Angle of deflection $\alpha$	$11.2 \times 10^{-6}$	$9.9 \times 10^{-6}$	$11.2 \times 10^{-6}$	$10.63 \times 10^{-6}$	$9.66 \times 10^{-6}$
Bi-prism separation $x_b$	---	---	$10 \mu\text{m}$	$10 \mu\text{m}$	$3.65 \mu\text{m}$
Number of bi-prisms $N_b$	1	1	101	101	101
Number of X-ray sources $N_s$	25	25	17	17	17
Source size $\Delta$	$7.00 \times 10^{-7}$ m	$7.00 \times 10^{-7}$ m	$8.00 \times 10^{-7}$ m	$8.00 \times 10^{-7}$ m	$8.00 \times 10^{-7}$ m
Pt. source separation $x_0$	$36.7 \mu\text{m}$	$36.7 \mu\text{m}$	$4.8 \mu\text{m}$	$4.8 \mu\text{m}$	$6 \mu\text{m}$
Distance between source and bi-prism $\eta$	0.4 m	0.4 m	0.53 m	0.53 m	0.51 m
Distance between bi-prism and detector $z$	1.38 m	1.4 m	1.14 m	1.05 m	1.19 m
Fringe separation	$14.1 \times 10^{-6}$ m	$16.1 \times 10^{-6}$ m	$10.0 \times 10^{-6}$ m	$10.0 \times 10^{-6}$ m	$3.69 \times 10^{-6}$ m
Factor of increase in fringe intensity	46.2	48.6	2176	2052	1733

328

\* CXRO online calculator<sup>26</sup>

329

\*\*SU-8 is similar to PMMA in CXRO online calculator<sup>26,27</sup>

330

+ X-Ray Optics Calculator ([http://purple.ipm.ru/xcalc/xcalc\\_mysql/ref\\_index.php](http://purple.ipm.ru/xcalc/xcalc_mysql/ref_index.php))<sup>28</sup>

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In a previous publication<sup>4</sup> we showed that the distance to the maximum fringe pattern is sensitive to

336 small changes in the angle of the bi-prism. To study this, one can look at a simple case of one point source

337 and one bi-prism. We first write the intensity in Eq. (1) as a function of the angle of the bi-prism  $\chi$ :

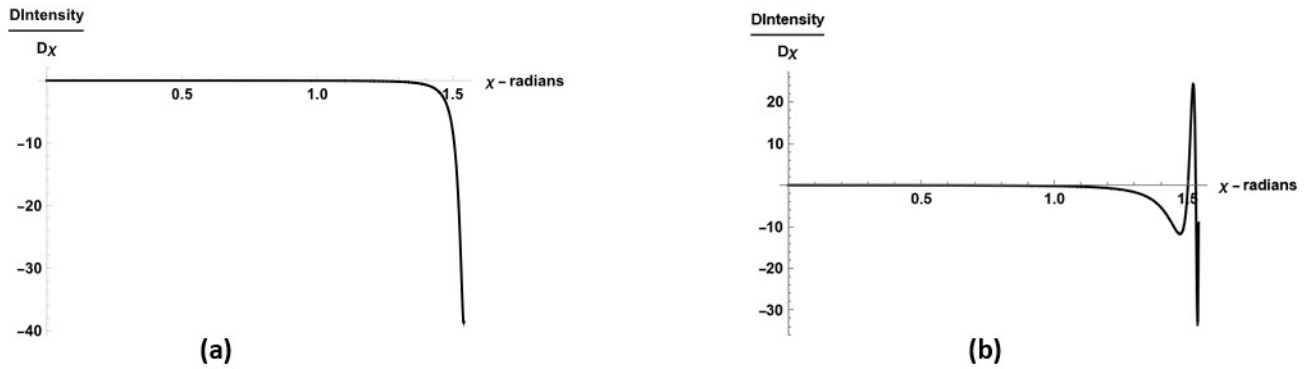
338  $Intensity[\chi, x, z]=\dot{I}$

$$339 \quad I_p \left( \Delta^2 + \frac{\Delta}{\pi z} ((\eta+z)*\lambda) / (2 \tan[\delta \tan(\chi)]) \cos \left\{ \frac{4 \pi x \eta \tan[\delta \tan(\chi)]}{(\eta+z)*\lambda} \right\} \sin \left\{ \frac{2 \pi z \Delta \eta \tan[\delta \tan(\chi)]}{(\eta+z)*\lambda \eta} \right\} \right), \quad (4)$$

340 where  $z$  is the distance between the bi-prism and the fringe pattern and  $x$  is the lateral distance on the  
 341 detector. One can then take the partial derivative of the intensity with respect to the angle to come up with  
 342 the expression:

$$343 \quad \frac{\partial Intensity(\chi, x, z)}{\partial \chi} = \frac{-1}{2 \pi z} I_p \delta \Delta \csc[\delta \tan(\chi)] \sec(\chi)^2 \dot{I}$$

344 Using this expression, we obtain plots of the sensitivity of the Intensity as a function of the angle of the  
 345 bi-prism (Fig. 10). One sees that the derivative of the Intensity is fairly flat near the central fringe  $x=0$  for  
 346  $z=1$  m until the angle nears  $\pi/2=1.5708$ ; then it changes rapidly for an energy of 17.5 keV [Fig.10(a)]. For  
 347 59 keV, the change is more dramatic near  $\pi/2$  [Fig.10(b)].



348

349 Fig. 10. Sensitivity as a function of bi-prism angle  $\chi$  for X-ray source energies of (a)  $\lambda=7.1 \times 10^{-11}$  m  
 350 (17.5 keV) and (b)  $\lambda=2.14 \times 10^{-11}$  m (59 keV). For the calculation we set  $N_s=1$ ,  $N_b=1$ ,  $I_p=1/\Delta^2$ ,  
 351  $\Delta=8.00 \times 10^{-7}$  m,  $\delta=8.7 \times 10^{-7}$  (SU-8),  $\eta=0.53$  m,  $x=0.000001$  m, and  $z=1$  m.

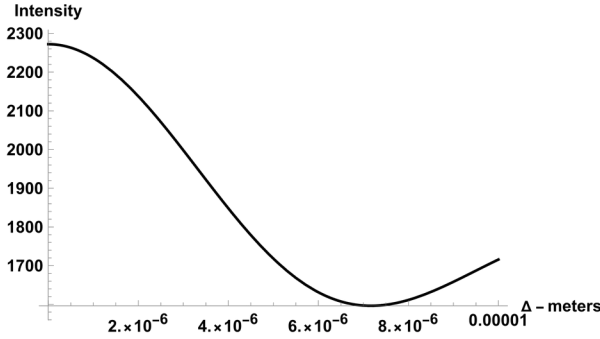
352

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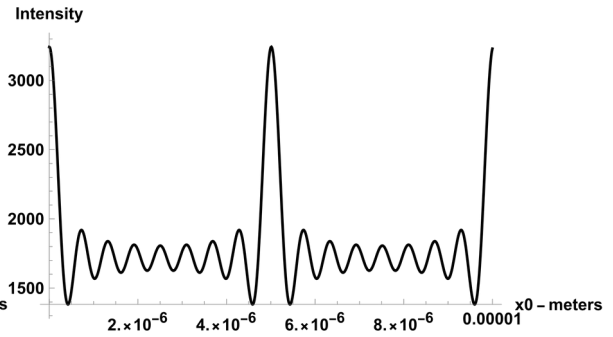
354 Using Eq. (4) one can study the Intensity variation due to the size of the point source aperture  $\Delta$  and the  
 355 point source separation  $x_0$ . For the parameters in Fig. 9, increasing the point source aperture  $\Delta$  will as  
 356 shown in Fig. 11 decrease the maximum fringe intensity until the aperture  $\Delta$  approaches that of a point  
 357 source aperture separation of  $4.8 \mu\text{m}$  with a minimum around  $7 \mu\text{m}$ . The Intensity then increases as the

358 apertures continue to overlap the separation between point sources. In Fig. 12 we see that the Intensity will  
 359 be maximum for point source separations  $x_0$  at factors of  $x_b/2=5\mu\text{m}$ . This is helpful because we can see  
 360 that if we change the value of  $x_b$  to 5 in Fig. 9, we would have come up with a better factor of increase in  
 361 intensity of 3238 with a fringe separation of  $10.0 \times 10^{-6}$  m at  $z=1.06$  m.

362



363



364

365

Fig. 11. Intensity as a function of point source size  $\Delta$  :

366

$Intensity = \frac{13736 \Delta + \sin(626988 \Delta)}{8 \Delta}$ . For the  
 calculation we set  $N_s=17$ ,  $N_b=101$ ,  $I_p=1/\Delta^2$ ,  
 367  $\delta=8.7 \times 10^{-7}$  (SU-8),  $\chi=85.32^\circ$ ,  $\eta=0.53$  m,  
 $x_0=4.8 \mu\text{m}$ ,  $x_b=10 \mu\text{m}$ ,  $x=0$  m, and  $z=1.06$  m.

368

Fig. 12. Intensity as a function of separation between point  
 sources  $x_0$ . For the calculation we set  $N_s=17$ ,  $N_b=101$ ,  
 $I_p=1/\Delta^2$ ,  $\Delta=8.00 \times 10^{-7}$  m,  $\delta=8.7 \times 10^{-7}$  (SU-8),  
 $\chi=85.32^\circ$ ,  $\eta=0.53$  m,  $x_b=10 \mu\text{m}$ ,  $x=0$  m, and  
 $z=1.06$  m.

## 369 **4. DISCUSSION**

370 This study provides expressions for fringe patterns of bi-prism interferometry and demonstrates  
371 maximum fringe visibility regions that are not periodic as is common with Talbot–Lau interferometry  
372 systems using diffraction gratings. Using the same analytical expressions, we calculate maximum fringe  
373 contrast and fringe periods for different X-ray source sizes and separations; and bi-prism design parameters  
374 of material, angle, size, number, period, and distances between the source and detector. The expressions  
375 are used to design and calibrate our laboratory system so that experiments can be performed that would be  
376 appropriate in the design of X-ray CT systems with higher energies and with high-power X-tubes capable  
377 of providing phase images almost instantaneously.

378

### 379 **4.A. Advantages of a single exposure bi-prism interferometry system**

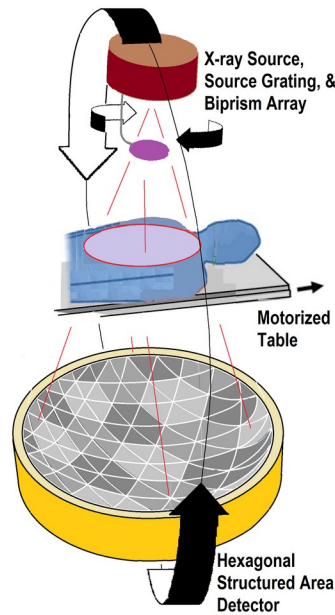
380 The well-established grating-based method of phase contrast X-ray imaging may yet prove capable of  
381 clinical applications. Still, alternative approaches such as the bi-prism method herein discussed holds  
382 several possible advantages. For good flux throughput, we have constructed a bi-prism array from the  
383 polymer SU8, a popular photoresist useful with LIGA processes. SU8 is a photoactive compound with some  
384 high Z components in low percentages. For 17 keV, the data yields 1/e absorption lengths for the polymer of  
385 1825  $\mu\text{m}$ , or a transparency of 97% for our optic design.

386 A bi-prism works well with polychromatic X-rays, as the fringe period is, in part, determined by the  
387 separation of the two virtual sources. This spacing is not fixed, as with a grating or Young's double slit, but  
388 rather varies dependent on the refractive index of the prism material relative to X-ray wavelength. Whereas,  
389 under polychromatic illumination the design of a grating system should be optimized for a particular  
390 frequency of X-rays.<sup>29</sup>



391 Moreover, the fringe period generated from a laboratory bi-prism setup has a relatively large  
392 magnification factor of approximately 2.5X. Some publications<sup>30,31</sup> have shown that modifying the standard  
393 Talbot-Lau interferometer where by the  $G_1$  grating is placed close to the source can also increase  
394 magnification factors similar to those in our bi-prism system. A larger fringe period at the detector could  
395 allow using a structured scintillator to record Moiré patterns. A hexagonal structured scintillator (Fig. 13)  
396 might aid the fast acquisition of phase contrast data of X-ray attenuation, phase, and small angle scatter.  
397 However, there are some challenges with using bi-prisms as well, mostly related to poorer spatial resolution.  
398 Even though, bi-prism-based interferometry, or other refractive-optical approaches, hold the potential to  
399 revolutionize phase-contrast and dark-field X-ray CT.

400



401

402

403 Fig. 13. Proposed cone-beam tensor tomography system. The X-ray source and optics rotate around an  
404 axis aligned to the center of an area detector with hexagonally segmented crosstalk-free scintillation  
405 elements. The components are mounted on a revolving gantry (for clarity, not shown here). X-ray scatter  
406 data is captured for three specific orientations 120° apart for every angular view.

407

408

409 In our previous paper<sup>32</sup> we performed simulations using wave optics to evaluate X-ray phase contrast  
410 imaging with grating interferometry using phase stepping and with bi-prism interferometry using single

411 exposure Moiré fringe analysis. Results of reconstructed vector coefficients of small angle scatter showed  
412 that bi-prism interferometry with single-exposure Moiré fringe analysis has an increase in noise compared  
413 to grating interferometry using phase stepping because of 1/8th the number of projection samples but  
414 provides better CNR due to better photon flux with less dose and faster acquisition times.

415

#### 416 ***4.A.1. Design of structured scintillators***

417 The implementation of a structured scintillator<sup>5,6</sup> fits well into our design of an X-ray bi-prism  
418 interferometry system with single-exposure Moiré fringe analysis because it reduces crosstalk between  
419 pixels of our high-resolution large fringe separation. Though the authors of this referenced work aimed  
420 their development for grating systems, a structured scintillator is ideal for our bi-prism interferometry  
421 system. The structure scintillator consists of filling an array of silicon pores with a thallium-doped cesium  
422 iodide [CsI(Tl)] scintillator. The pores act as a mold which are filled with melted CsI(Tl) powder. Before  
423 filling, the walls are coated thermally with an oxide layer to provide a wave guide to direct the light and  
424 reduce noise between pixels. The design of a structured scintillator prevents crosstalk and allows  
425 opportunity for Moiré fringe development when illuminated with spatially modulated X-ray illumination.  
426 Moiré patterns occur because of the interaction between the pixelated image sensor and the periodic pore  
427 structure of the scintillator. The magnitude of this occurrence depends on the ratio of the respective pitches,  
428 the geometries involved and the amount of angular misalignment. The Moiré pattern is ideal for the  
429 extraction of dark field scatter by taking the first harmonic of the Fourier transform. The work of  
430 Rutishauser et al.<sup>6</sup> showed an increase in fringe visibility at X-ray energies around 60 keV (greater than our  
431 simulations of 17.5 keV), leading to an improvement in the quality of the differential phase and dark field  
432 data compared to a conventional grating interferometry system.

433

#### 434 ***4.A.2. Parameter range for a bi-prism interferometry system***

435 In our laboratory setup, we have been using a Rigaku X-ray tube with molybdenum anode operating at  
436 35kVp, where 17.5 keV is the predominate energy with a point-size of 100  $\mu\text{m}$  and a 15 mm wide gold  
437 source grating with 0.8  $\mu\text{m}$  apertures separated by 4.8  $\mu\text{m}$ . We assume the multiple point source apertures  
438 are the “open windows” of the source grating. These are much smaller than the “closed bars” of a typical  
439 source grating and provide to the bi-prism array multiple coherent sources of X-rays, which are each  
440 mutually incoherent. There is some reduction in image resolution by using laboratory X-ray tube sources.  
441 The actual laboratory X-ray source size is that of a typical X-ray tube and determines the number of “open  
442 window” point sources that illuminate a single bi-prism. We assume that each point source aperture will  
443 radiate 17 separate bi-prisms. The intensity of the fringe visibility increases with the number of sources and  
444 the number of sources seen by each bi-prism. The intensity is maximum if the source separation is a factor  
445 of one-half the width of the bi-prism. The intensity also increases with the increase in the angle of the bi-  
446 prism. The rate of change of this increase is fairly flat until the angle of the bi-prism nears  $90^\circ$ .

447 Our bi-prism system design, like that proposed by Yaroshenko *et al.*<sup>33</sup> for a grating interferometry  
448 system, offers the potential for future X-ray CT application by providing high fringe visibility at shorter  
449 propagation distances. They demonstrated that using phase gratings with triangular-shaped structures in an  
450 X-ray interferometer can yield high visibilities for significantly shorter propagation distances than  
451 conventional gratings with binary structures. This is especially important in the design of human systems  
452 where bore sizes are approximately 75-85 cm.

453 Although similar in construction as Yaroshenko *et al.*<sup>33</sup> our interferometry methods follow Doblas *et al.*<sup>2</sup>-  
454 albeit with X-rays rather than visible light. Where Doblas *et al.* use multiple coherent sources and a single  
455 Fresnel bi-prism, we use an array of Fresnel bi-prisms. The full setup yields multiple fringe patterns in  
456 resonant superposition with a large field of illumination. Fundamentally, this is a different way to make  
457 interference fringes. A single coherent X-ray source can illuminate a single Fresnel bi-prism and produces  
458 interference fringes downfield. The fringe patterns produced by our array of bi-prisms are only indirectly

459 related to the array period, but are the direct result of the periodicity of the virtual sources created by the  
460 individual bi-prisms in the array. Interference fringes produced by a grating are directly related to the  
461 period of the grating, and this includes a triangular-profile grating such as used by Yaroshenko *et al.*<sup>33</sup>  
462 Moreover, their work highlights the benefit of closer spacing of 10 cm between the phase grating and  
463 analyzer grating. This is the opposite effect we are pursuing in our approach with a bi-prism array to  
464 analyzer/structured detector of 1.06 m.

465 Periodic replication of a grating, which is the Talbot effect, results from the overlapping zero-order, 1st  
466 order and -1st order diffracted beams, when those beams encounter a screen (detector). Yaroshenko *et al.*<sup>33</sup>  
467 rightly calls for future work on sinusoidal amplitude gratings as those gratings produce only those  
468 diffraction orders. Their triangular-shaped grating system clearly demonstrates periodic longitudinal  
469 spacing of fringe visibility regions.

470 Our goal has been to keep the dimensions of our laboratory system within something that would fit a  
471 present-day large bore X-ray CT scanner. Though our initial work is aimed at imaging small objects at  
472 energies of 17.5 keV, we have performed calculations using Eq. (3) to see how the parameters of the bi-  
473 prism system would change at higher X-ray energies. For example, if we were to image at 59 keV, because  
474 of the change in refraction index decrement at higher energies, we would have to modify the material of the  
475 bi-prism to something like nickel in Table 1 to obtain reasonable fringe visibility at 1.19 m from the bi-  
476 prism array with a fringe separation of  $3.69 \mu m$ .

477 Recent advances in grating fabrication techniques<sup>34,35</sup> alleviate some concerns regarding high-energy X-  
478 ray phase contrast imaging. The pattern height of our laboratory  $G_0$  grating is 24 microns, however 200  
479 microns are now available, which could be available for higher energies. Much of the X-ray flux is  
480 absorbed by  $G_0$ . Thus, with higher energies the thicker  $G_0$  grating will have its effect on flux. However,  
481 current CT tube technology, can operate at >100 kW for the scan duration necessary to acquire clinical  
482 tensor tomographic data including attenuation, phase, and small angle scatter information.

483

484 **4.A.3. Advantages of single shot approaches**

485 There are single shot approaches,<sup>7-11</sup> which render phase, scattering, and attenuation data from single  
486 projection measurements without mechanical phase stepping measurements at each projections angle. The  
487 challenges of mechanical phase stepping are the need to obtain precise grating posing at multiple  
488 measurements per projection angle, whereas single shot approaches of data acquisition have shorter  
489 acquisition times and have better mechanical stability. However, the paper of von Teuffenbach *et al.*<sup>36</sup> does  
490 point out some negative aspects of previous single shot methods. For example, the reverse projection  
491 method<sup>8</sup> combines two interferograms taken exactly 180 degrees apart, which requires stability of the  
492 grating alignment for each view and gives only access to attenuation and differential phase-contrast  
493 projection images. The implementation of helical and cone beam tomography would be difficult. The Moiré  
494 analysis method<sup>7</sup> using a slight detuning of the interferometer suffers from loss in resolution in one  
495 dimension of the detector. The sliding window method<sup>11</sup> requires only a single interferogram per projection  
496 angle, but changes the grating position at each angular position. Missing interferograms are simply  
497 interpolated from adjacent projection angles. This works well for scans with high angular sampling.

498 The paper of von Teuffenbach *et al.*<sup>36</sup> compares three single shot approaches and demonstrate that it is  
499 possible to successfully reconstruct the linear attenuation coefficient, the refractive index and the linear  
500 diffusion coefficient (measure of small angle scatter attenuation) using a single measurement per projection  
501 angle without grating movements. In their paper using the interferogram-based statistical iterative  
502 reconstruction (IBSIR) method of data reconstruction, they compare: 1. an equidistant phase stepping (PS)  
503 acquisition where the analyzing grating was shifted by 1/3rd the grating period to obtain 3 measurements at  
504 101 equal angular positions over 360° (identified as PS-IBSIR); 2. a sliding window phase-stepping (SW)  
505 acquisition that records only one measurement at each of 303 equal angular positions over 360° with the  
506 analyzing grating shifted by 1/3rd its period before the next acquisition (SW-IBSIR); and 3. a single shot

507 electromagnetic phase-stepping acquisition (SSEPS) recording a single image at each of 303 equal angular  
508 positions over 360° with the position of the X-ray focal spot shifted from the previous acquisition before  
509 each measurement (SSEPS-IBSIR). These three methods provided the same number of measurements. The  
510 PS-IBSIR showed artifacts at sharp boundaries, probably due to the finer angular sampling even though the  
511 same number of measurements. The SW-IBSIR and SSEPS-IBSIR methods showed similar X-ray  
512 attenuation reconstruction with the SW-IBSIR method showing a slightly better contrast but showing more  
513 artifacts in the phase reconstruction and somewhat more noise in the reconstruction of the scatter diffusion  
514 coefficient. As can be seen from the author's very nice display of the Radon space sampling, even though  
515 both methods provide the same number of angular sampling, the regular sampling of SW-IBSIR method  
516 probably has more aliasing than the irregular sampling of the SSEP-SIBSIR method with the motion of the  
517 X-ray focus spot. This Radon space sampling proposes interesting future investigations similar to those that  
518 have been pursued using data compression with irregular Fourier space sampling in MRI.<sup>37,38</sup> The proposed  
519 electronic phase stepping acquisition protocol incorporated into an interferometry system would be  
520 perfectly suited for application with a clinical CT system because the required focal spot sweeping is  
521 already available in most modern clinical CT systems and the technique is well suited for a continuously  
522 rotating gantry; however, it supposes one could design an X-tube with electronic phase stepping without  
523 focal spot sweeping.

#### 524

#### 525 **4.B. X-ray tube design for virtual electronic phase stepping**

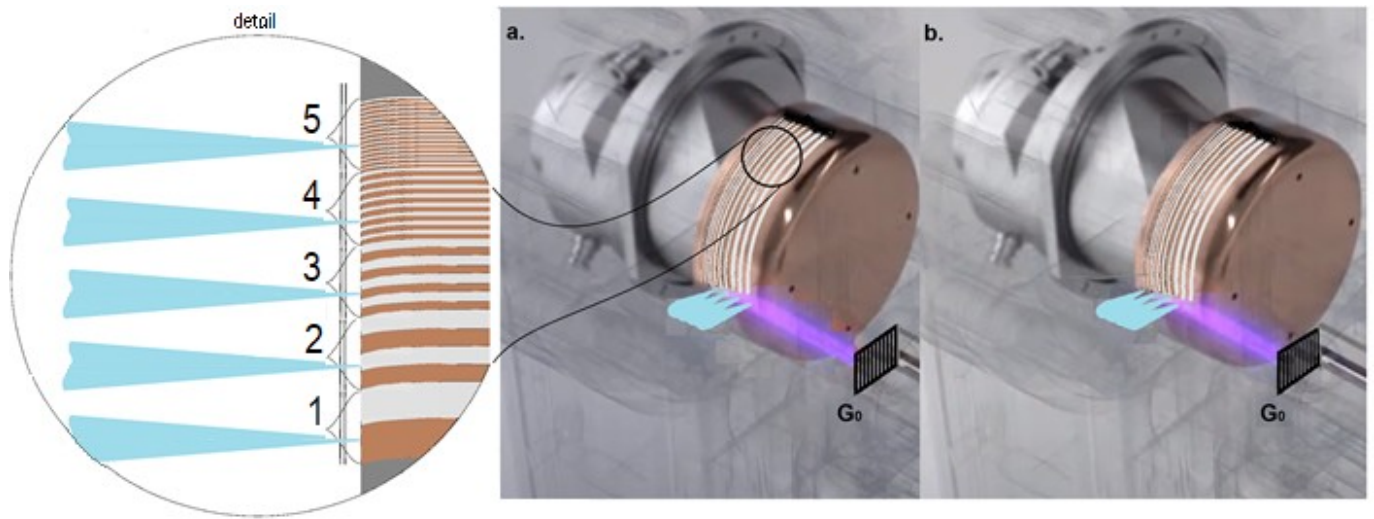
526 Electromagnetic phase-stepping (EPS)<sup>39</sup> moves an electrical beam to move projections of the object  
527 across the detector whereby the shifted images are combined to form pseudo phase-stepping images. Here  
528 we propose a single-shot image acquisition approach using a novel X-ray source that provides for “virtual”  
529 electronic phase stepping without focal spot sweeping. This approach involves using an X-ray  
530 interferometer imaging system where the X-ray tube has multiple X-ray origins/spots. The tube is operated

531 in such a way that each spot is temporally modulated at differing frequencies, and spatially modulated to  
532 form illumination beams with differing fringe locations as observed in the interferometer. Composite  
533 images of an object acquired with the interferometer are subsequently separated according to the flashing  
534 source that provided the illumination. This is accomplished via de-multiplexing transform processes<sup>40</sup> based  
535 on the known temporal multiplexing transform of the source. After an alignment step, the isolated images  
536 from a single composite image/exposure have the same view of the object but with different fringe  
537 locations relative to a feature in the object. An isolated image group of each single view are then analyzed  
538 to yield images representing the differential phase shift, attenuation, and X-ray scattering features of the  
539 object. All of the X-ray optical components in the interferometer, either a Talbot-Lau or a biprism setup,  
540 remain fixed and motionless.

541 A high-power X-ray tube design that does this full work of virtual electronic phase stepping during a  
542 single exposure is illustrated in Fig. 14. Here, a binary multiplexing transform of the X-ray output is  
543 conducted by normal tube operation. Multiple X-ray spots are generated on the side of a rotating and  
544 linearly-translated cylinder. An X-ray window allows that  $G_0$  is illuminated by each spot, but at slightly  
545 different angles. The oscillating linear translation, about 2 Hz, carries the anode past separate, adjacent,  
546 focused, and identical high-current electron guns. As the anode cylinder walls are patterned, either with  
547 metal/insulator or with ridges/grooves, in varying stepped spatial periods, each X-ray spot turns on/off at  
548 different frequencies.

549 Other multiplexing transforms can result from different patterning, which might work better with slower  
550 detectors. Lower power tubes can control the electron guns independently to accomplish a similar  
551 multiplexed X-ray tube output.

552



553

554 Fig. 14. Illustration of multiple-spot cylindrical rotating anode X-ray tube with oscillating translation for  
 555 phase/dark-field single-shot data acquisition. (a.) Anode at minimum extension, (b.) Anode at maximum  
 556 extension, **detail:** showing multiple spot tracks/flash rates 1.) Low frequency spot, 2.) Lower mid-  
 557 frequency spot, 3.) Higher mid-frequency spot, 4.) High-frequency spot, 5.) Very high-frequency spot.  
 558 (Image courtesy Rigaku Corp., Akishima-shi, Tokyo, Japan.)

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561

562



## 563 5. CONCLUSION

564 Phase-contrast and dark-field X-ray CT for clinical human use will require advancements in X-ray optics,  
565 X-ray sources, and X-ray detectors. As an alternative X-ray optical component, a Fresnel bi-prism array  
566 adds to the set of tools available for the development of phase contrast imaging systems. Here, we present  
567 one possible approach that is somewhat like an “inverse geometry Talbot-Lau” setup, but with improved  
568 fringe visibility. Other bi-prism-based interferometry approaches are possible, including combinations with  
569 phase gratings. We also present an advanced X-ray tube design that offers single-shot “virtual electronic  
570 phase stepping.” These two technologies, when combined with previously demonstrated structured-  
571 scintillator detectors, can provide separate phase, dark-field, and attenuation signals without the need for  
572 mechanical phase stepping. A system employing these components could be made to capture X-ray dark  
573 field projections for direct tomographic tensor reconstruction of small angle scatter. The review of the  
574 literature presents evidence of advancement in hardware (without grating phase stepping) and of new  
575 algorithm developments that could provide fast single shot acquisition with mechanical simplicity  
576 delivering less radiation exposure and faster acquisition times.

577 **ACKNOWLEDGEMENTS**

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579 R01 EB026332, and R01 HL135490.

580

581 **CONFLICT OF INTEREST**

582 The authors have no conflict of interest.

583

584 **DATA AVAILABILITY**

585 Data distribution is available through an appropriate data sharing agreement between institutions.

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## FIGURE CAPTIONS

680

681 Fig. 1. (a) Schematic diagram of an X-ray grating interferometry system. (b) Talbot-Lau carpet.  
682 Illuminating plane wave passes through a grating producing a fringe pattern with replicating amplified  
683 fringe patterns at regular distances from the sources produced by the grating. At  $Z_T/2$  there is a secondary  
684 Talbot image and at  $Z_T$  a replication of the original Talbot image that emerged from the grating. At  $Z_T/4$   
685 there is a double frequency fractional image and increased frequency of images at less fractional distances.  
686 (Modified from Wikipedia: [https://en.wikipedia.org/wiki/Talbot\\_effect](https://en.wikipedia.org/wiki/Talbot_effect).)

687

688 Fig. 2. Schematic diagram of an X-ray bi-prism interferometry system. (a) Operation of a Fresnel X-ray  
689 bi-prism, (b) experimental setup with multiple X-ray origins each producing fringes in resonant position,  
690 and (c) detail of production of Moiré fringes via a tilted grating  $G_2$  at the detector. Our intention is to  
691 eliminate the  $G_2$  grating and instead use only a detector/scintillator with small hexagonal elements to  
692 produce the Moiré fringe pattern. (Modified from our earlier publication.<sup>4</sup>)

693

694 Fig. 3. X-ray beam deflection by a bi-prism illustrating the distance  $d$  between the two virtual coherent  
695 point sources produced by a bi-prism. (Copied from The Fresnel Biprism, David-Alexander Robinson and  
696 Jack Denning; Daniel Tanner 08332461; 10th December 2009.)

697

698 Fig. 4. Lab setup of the bi-prism interferometry system at UCSF. The bi-prism interferometry system  
699 includes a molybdenum-target X-ray tube operating at 35 kVp with predominate energy of the 17.5 keV  
700 characteristic line (designed energy/wavelength). Shown also is a gold source grating with  $0.8 \mu m$  apertures  
701 separated by  $4.8 \mu m$ , a bi-prism array made of plastic/epoxy photoresist material (SU-8), and a Stanford  
702 Photonics iCCD camera.

703

704 Fig. 5. X-ray beam deflection by bi-prism. Here is given the expression for the angle  $\alpha$  of the beam  
705 deflection and the distance  $d$  between the two virtual coherent point sources produced by a bi-prism.  $\delta$  is  
706 the refraction index decrement. Notice the difference in refraction between X-rays and light. (Modified from  
707 Chegg Study, www.x-ray-optics.de, and researchgate.net)

708

709 Fig. 6. Density plot of a fringe pattern for 25-point sources with 1 bi-prism. The amplification of the  
710 interference pattern is repeated at non-periodic distances away from the plane of the bi-prism. For the  
711 calculation we set  $\lambda=7.1 \times 10^{-11}$  m (17.5 keV),  $I_p=1/\Delta^2$ ,  $\Delta=7.00 \times 10^{-7}$  m,  $\alpha=\delta \tan(\chi)$ ,  $\delta=1.58 \times 10^{-6}$   
712 (silicon),  $\chi=82^\circ$ ,  $\eta=0.4$  m, and  $x_0=36.7 \mu\text{m}$ . This figure is modified with new parameters from that of our  
713 previous publication.<sup>4</sup>

714

715 Fig. 7. Density plot of a fringe pattern for 17-point sources with 1 bi-prism for a different set of  
716 parameters than those in Fig. 6. For the calculation we set  $\lambda=7.1 \times 10^{-11}$  m (17.5 keV),  $I_p=1/\Delta^2$ ,  
717  $\Delta=8.00 \times 10^{-7}$  m,  $\alpha=\delta \tan(\chi)$ ,  $\delta=8.7 \times 10^{-7}$  [plastic/epoxy photoresist material (SU-8)],  $\chi=85.3^\circ$ ,  
718  $\eta=0.53$  m, and  $x_0=4.8 \mu\text{m}$ ,  $x_b=10 \mu\text{m}$ . There is a fringe separation of  $1.09 \times 10^{-5}$  m at  $z=1.2$  m.

719

720 Fig. 8. Density plot of a fringe pattern for 17-point sources with 1 bi-prism for the same set of parameters  
721 as those in Fig. 7, except  $x_0$  was changed from  $x_0=4.8 \mu\text{m}$  to  $x_0=36.7 \mu\text{m}$  For the calculation:  
722  $\lambda=7.1 \times 10^{-11}$  m (17.5 keV),  $I_p=1/\Delta^2$ ,  $\Delta=8.00 \times 10^{-7}$  m,  $\alpha=\delta \tan(\chi)$ ,  $\delta=8.7 \times 10^{-7}$  [plastic/epoxy  
723 photoresist material (SU-8)],  $\chi=85.3^\circ$ ,  $\eta=0.53$  m, and  $x_0=36.7 \mu\text{m}$ ,  $x_b=10 \mu\text{m}$ .

724

725 Fig. 9. Density plot of a fringe pattern for 17-point sources with 101 bi-prism for the same set of  
726 parameters as in Fig. 7. For the calculation we set  $\lambda=7.1 \times 10^{-11}$  m (17.5 keV),  $I_p=1/\Delta^2$ ,  $\Delta=8.00 \times 10^{-7}$  m,



727  $\alpha = \delta \tan(\chi)$ ,  $\delta = 8.7 \times 10^{-7}$  [plastic/epoxy photoresist material (SU-8)],  $\chi = 85.32^\circ$ ,  $\eta = 0.53$  m, and  
728  $x_0 = 4.8 \mu\text{m}$ ,  $x_b = 10 \mu\text{m}$ . There is a fringe separation of  $1.00 \times 10^{-5}$  m at  $z = 1.05$  m.

729

730 Fig. 10. Sensitivity as a function of bi-prism angle  $\chi$  for X-ray source energies of (a)  $\lambda = 7.1 \times 10^{-11}$  m  
731 (17.5 keV) and (b)  $\lambda = 2.14 \times 10^{-11}$  m (59 keV). For the calculation we set  $N_s = 1$ ,  $N_b = 1$ ,  $I_p = 1/\Delta^2$ ,  
732  $\Delta = 8.00 \times 10^{-7}$  m,  $\delta = 8.7 \times 10^{-7}$  (SU-8),  $\eta = 0.53$  m,  $x = 0.000001$  m, and  $z = 1$  m.

733

734 Fig. 11. Intensity as a function of point source size  $\Delta$ :  $Intensity = \frac{13736 \Delta + \sin(626988 \Delta)}{8 \Delta}$ . For the  
735 calculation we set  $N_s = 17$ ,  $N_b = 101$ ,  $I_p = 1/\Delta^2$ ,  $\delta = 8.7 \times 10^{-7}$  (SU-8),  $\chi = 85.32^\circ$ ,  $\eta = 0.53$  m,  $x_0 = 4.8 \mu\text{m}$ ,  
736  $x_b = 10 \mu\text{m}$ ,  $x = 0$  m, and  $z = 1.06$  m.

737

738 Fig. 12. Intensity as a function of separation between point sources  $x_0$ . For the calculation we set  $N_s = 17$ ,  
739  $N_b = 101$ ,  $I_p = 1/\Delta^2$ ,  $\Delta = 8.00 \times 10^{-7}$  m,  $\delta = 8.7 \times 10^{-7}$  (SU-8),  $\chi = 85.32^\circ$ ,  $\eta = 0.53$  m,  $x_b = 10 \mu\text{m}$ ,  $x = 0$  m,  
740 and  $z = 1.06$  m.

741

742 Fig. 13. Proposed cone-beam tensor tomography system. The X-ray source and optics rotate around an  
743 axis aligned to the center of an area detector with hexagonally segmented crosstalk-free scintillation  
744 elements. The components are mounted on a revolving gantry (for clarity, not shown here). X-ray scatter  
745 data is captured for three specific orientations  $120^\circ$  apart for every angular view.

746

747 Fig. 14. Illustration of multiple-spot cylindrical rotating anode X-ray tube with oscillating translation for  
748 phase/dark-field single-shot data acquisition. (a.) Anode at minimum extension, (b.) Anode at maximum  
749 extension, **detail**: showing multiple spot tracks/flash rates 1.) Low frequency spot, 2.) Lower mid-

750 frequency spot, 3.) Higher mid-frequency spot, 4.) High-frequency spot, 5.) Very high-frequency spot.

751 (Image courtesy of Rigaku Corp., Akishima-shi, Tokyo, Japan.)