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IONIZING RADIATION RISKS TO SATELLITE POWER SYSTEMS (SPS) WORKERS

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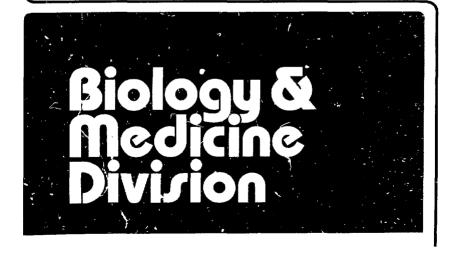
Lawrence Berkeley Laboratory

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IONIZING RADIATION RISKS TO SATELLITE POWER SYSTEMS (SPS) WORKERS

November 1980





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Committee for Assessing the Ionizing Radiation Risks to Satellite Power Systems (SPS) Workers

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SUMMARY

SCOPE

A reference Satellite Power System (SPS) has been designed by NASA and its contractors for the purposes of evaluating the concept and carrying out assessments of the various consequences of development, including those on the health of the space workers. The Department of Energy has responsibility for directing various assessments. Present planning calls for the SPS workers to move from Earth to a low earth orbit (LEO) at an altitude of 500 kilometers; to travel by a transfer ellipse (TE) trajectory to a geosynchronous orbit (GEO) at an altitude of 36,000 kilometers; and to remain in GEO orbit for about 90 percent of the total time aloft.

This report deals with the radiation risks to the health of workers who will construct and maintain solar power satellites in the space environment. The charge to the committee was:

- To evaluate the radiation environment estimated for the Reference System which could represent a hazard;
- b. To assess the possible somatic and genetic radiation hazards:
- c. To estimate the risks to the health of SPS workers due to space radiation exposure, and to make recommendations based on these conclusions.

ENVIRONMENT

GEO has the highest ionizing radiation intensity. The Reference System proposes that most workers will spend most of their tour in space in this environment. The radiation environment in GEO is known only with sufficient accuracy to predict the radiation dose in free space to within a factor of two, and then only for objects in orbit for relatively extended periods for which average values of the radiation environment may be used. The short-term enhancement of the radiation environment yeaused by geomagnetic substorms may result in significant deviations from the average doses and could increase the risk to health of exposed individuals caught outside shielded or protected areas. Solar particle events, which are rare events occurring most frequently in about eleven—year cycles, could increase the dose by a considerable factor.

Shielding is the major factor which influences the dose and thus the potential health risk to the SPS worker. With the exception of the storm cellar, the shielding of the Reference System used in these estimates of dose is the minimum amount of material needed for the

structural integrity of the system. The galactic cosmic ray dose is not readily attenuated by shielding material. This may be a factor for setting a lower limit on the radiation dose if the bremsstrahlung dose is adequately attenuated.

RADIATION DOSE

Based on the Reference System model and subject to its limitations and assumptions, it is estimated that an SPS worker in GEO may be exposed to a dose of as much as 40 rem during any one 90-day mission. This estimate is based on the following assumptions:

- a. That between 87 and 98 percent of the worker-years occur in GEO. For risk estimation, we therefore have assumed that the total mission will occur in GFO.
- b. That effective snielding of the worker will be equivalent to 8 grams per cm² aluminum (which includes 5 g per cm² of body self-shielding).
- c. That the exposure is represented by the conditions in a parking orbit of 160° west longitude. This is considered to be the worst case.
- d. That no contribution to the total dose is made by solar particle events.
- e. That the bulk of the dose is produced by trapped electron bremsstrahlung, and that the small contribution to the total dose made by galactic cosmic rays can be neglected.
- f. That the time spent in GEO is 90 days per space tour.

The uncertainty in these assumptions is of an order that the dose estimate could vary by a factor of two.

HEALTH RISK ASSESSMENTS

Using U.S. life tables of age and sex, the estimated lifetime risk for cancer is 0.8 to 5.0 excess deaths per 10,000 workers per rad of exposure. Thus, for example, in 10,000 workers who completed ten missions with an exposure of 40 rem per mission, 320 to 2,000 additional deaths, in excess of the 1640 deaths from normally occurring cancer, would be expected. These estimates would indicate a 20 to 120 percent increase in cancer incidence in the worker-population. The wide range in these estimates stems from the choice of the risk-projection model and the dose-response relationship. The choice between a linear and a linear-quadratic dose-response model may alter

the risk estimate for some tumors by a factor of at least two. The method of analysis (e.g., relative vs. absolute risk model) can alter the risk estimate by an additional factor of three. Choosing different age and sex distributions can further change the estimate by another factor of up to three. When decisions have been made about the selection of SPS workers, the precise influence of age and sex distribution can be included in risk estimates. However, the choice of dose-respose relationship and projection models at present is a matter of opinion and will not be resolved scientifically for quite some time.

CONCLUSTONS

The conclusions of the committee based on the findings above are:

- a. The risk of excess cancer deaths is assumed to be closer to the lower limit estimated above because the major exposure will be from the low dose rate, lcw-LET irradiation. This being the case, we consider a reasonable estimate to be one excess death per 10,000 workers per rem of exposure. If this level of risk is applied to the worst case reference system exposure level, namely 40 rem, there would be 400 excess cancer deaths in a work force of 10,000 completing ten missions (accumulative dose equivalent of 400 rem).
- b. The potential genetic consequences could be of significance, but at the present time, sufficient information on the age and sex distribution of the worker population is lacking for are: ise estimation of risk.
- c. The potential teratogenic consequences resulting from radiation are considered significant. Radiation exposure of a pregnant worker could result in developmental abnormalities.
- d. Based on the Reference System, dose to space workers from low-LET bremstrahlung approaches the cateratogenic level for man. The appropriate quality factor for the HZE particle portion of the dose is unknown at this time. If its Q is greater than 20, the cateract hazard may be significant. More information is needed regarding this nazard.
- e. In the absence of a radiation accident or some other unexpected situations (e.g., nuclear detonation) and with adequate snielding to protect against the increased radiation levels during solar particle events, there will be no early or acute radiation health effects occurring in the SPS worker population.
- f. No other radiation health effects are considered to be of sufficient consequence to be important for risk estimation.

RECOMMENDATIONS

This Committee strongly emphasizes the need to reduce the uncertainties in the evaluation of radiation health risks to SPS workers in the space environment. It recommends that the following be carried out to achieve these goals:

- a. The short-term variations of the radiation dose rate in space must be better understood so that the range of doses and dose rates to be expected can be established accurately. A radiation environment model should be developed that is appropriate to this SPS mission for study and simulation. The model should include short-term and solar-cycle variations.
- D. An instrumented research satellite should be placed in GEO to measure absorbed dose rate and particle spectra at depth in phantoms and to measure the temporal variations of the radiation field.
- c. The differences in the results of dose and dose-rate estimation obtained from the snielding transport codes must be evaluated. The use of different calculational methods with the same set of assumotions should yield the same results.
- d. When institutional decisions have been made to develop appropriate exposure strategies, engineering decisions for dose control (e.g., improved shielding) should then be made.

INTRODUCTION

BACKGROUND

The need to develop long-term, baseload, electrical energy sources has initiated search for economically competitive and environmentally acceptable alternatives to our limited supply of fossil fuels and other nonrenewable energy sources. Such a search must consider the widest range of available and potential technologies. Satellite-based solar power generation has emerged as one possible source of electrical energy obtained from inexhaustible and renewable sources.

The U.S. Department of Energy (DDE) and the National Aeronautics and Space Administration (NASA) are examining the feasibility of generating baseload electrical power with Satellite Power Systems (SPS) located in geostationary orbit. These systems would collect radiant energy from the Sun, convert it to electrical energy, and then beam it to Earth as microwaves. Ground receiving stations would convert the microwaves to electrical energy to be supplied to power grids.

DOE and NASA are studying the large number of factors involved in bringing such a complex undertaking into being. At the present time, initial system definition studies have been completed. A Reference System (DOE/NASA, 1978) has been designed as a basis for further development and evaluation of important areas of uncertainty and for preliminary assessment of environmental impacts and potential health risks to SPS workers. This report directs attention to the radiation risks to the health of workers who will construct and maintain the SPS satellites in the space environment.

THE SPACE ENVIRONMENT

Man faces substantial risks to health in space. Prolonged periods in space can result in workers being exposed to radiation that can cause adverse health effects such as cancer, genetically related ill health, cataracts, and—with very large doses—even death. Exposure to ionizing radiations is a major factor in the evaluation of potential health risks to workers in space in the SPS program. Assessment of these health risks involves the determination of existing radiation dose levels. Methods must be developed to reduce radiation levels for persons in space to the lowest practical level. Determination must be made of the acceptable radiation levels which may not be absolutely safe but, rather, may be appropriately safe for the special circumstances of the space environment.

Because SPS travel and work will take place in a number of different space environments—primarily low earth orbit (LEO), the transfer ellipse (TE), and geosynchronous earth orbit (GEO)—the worker in space will be exposed to radiation environments with different radiation qualities, varying intensities, and differing difficulty in predicting exposure levels. Assessment of potential health risks in each of the three environments is influenced by many factors, including: (1) the location in space; (2) the type of shielding used in work stations, living quarters, transport systems, and space suits; (3) the types of duties performed; (4) the length of each mission; (5) the age and sex of space workers; and (6) the total number of missions per worker—career. This assessment of the radiation health hazards in space is limited by the lack of adequate information concerning these factors and further complicated by inadequate data concerning the health effects of various types and dose levels of radiation on the human body.

The Reference System proposal is based on the following assumptions:

- 1. The SPS construction would occur mainly in GEO with all material transported from Earth.
- 2. There would be a two-phase transportation of materials and personnel: movement to LEO, followed by subsequent delivery to GEO.
- 3. After construction of the LEO base, cargo transportation vehicles for the voyage to GEO would be assembled in LEO.
- 4. The construction and operation of a fleet of 60 SPS facilities, assuming a 30-year lifetime for each SPS, would require between 22,000 and 57,000 worker-years in space. The range of worker-years required takes into consideration the number of maintenance workers needed per SPS (from 4 to 20). The number of SPS workers required to achieve these goals would range between 10,000 and 20,000, with ten 90-day missions per worker.
- 5. The time spent in GEO would be between 87 and 98 percent of the total worker-years.

2. REFERENCE SYSTEM EVALUATION

In determining potential radiation risks to the health of SPS workers in the space environment, an estimate must first se made of the radiation environment outside the space vehicle. This environment consists of trapped electrons and protons, galactic cosmic rays (GCR), and sporadic solar particle event (SPE). The average values of the flux densities of the primary radiation are presently known only to within a factor of two. The spacecraft and the space workers comprise a complex distribution of shielding materials which attenuates the primary radiation and which is also a source of secondary radiations. The transport of the radiation through the shield is fairly well understood, but there is uncertainty in the calculations due to assumptions of shield geometry and composition.

Dose estimates for previous space missions have been based on three-dimensional distribution of the shielding materials in the spacecraft and in the astronauts. Other calculations have been based upon solid-angle sectoring of the available shielding. All calculations for an SPS mission thus far have been made assuming simple geometries. Based on a recent analysis (Selzer, 1979), the absorbed dose calculated inside a spherical shell at a given rad us is three to four times that at the same depth within a semi-infinite slab with isotropic radiation impinging on it from one hemisphere. For the purposes of this report, spherical shield geometry is therefore used to provide a worst case estimate of dose. However, this SPS Committee recognizes that until calculations for SPS radiation environments are based on more realistic shielding configurations, these calculations remain inadequate for detailed, accurate estimation of dose equivalent.

In this report, the shielding assumed is $3~g/cm^2$ of aluminum for the habitat and work stations and 20–30 g/cm^2 for the storm cellar to be used during solar particle events (SPE). The assumptions are based on the Boeing contribution to the Reference System (DOE/NASA, 1978). In addition, an average of $5~g/cm^2$ of aluminum equivalent shielding is assumed for the body self-shielding in the present analysis. Thus, the estimates made below can be considered representative average doses at the center of a sphere of radius $3~g/cm^2$ of aluminum equivalent material.

Because the relative biological effectiveness of the different radiations is variable, a quality factor (Q) is assigned to each radiation, permitting calculations of a total dose equivalent (rem). The use of Q is reasonable for biological endpoints of principal concern. However, there is the possibility that other health effects are caused only by high energy heavy ions (HZE particles). Therefore, at the present time, the use of a Q for the HZE particles does not provide a complete assessment of the risk to health of SPS workers

ESTIMATED ABSORBED COSES AND DOSE EQUIVALENTS IN SPS WORKERS

Low Earth Orbit

Dose estimates are most accurate for the LEO phase of the mission. Protons in the South Atlantic Anomaly pose the only major source of external radiation. The dose equivalent in LEO will vary by about a factor of two between solar minimum and solar maximum (Stassinopoulos, 1979). The dose rate estimates at solar minimum, when the doses are higher, range from 0.15 rad per day (Hardy, 1979) to 0.3 rad per day, the latter value corrected from a semi-infinite slab calculation (Stassinopoulos, 1979) to spherical shielding (Seltzer, 1979). Since Q for this radiation is close to unity, these values are good estimates of the dose equivalent rate in rem per day. There will be a negligible contribution to the dose from galactic cosmic rays and solar particle events, due to the large amount of geomagnetic shielding available in the LEO trajectory. The total dose equivalents for a 90-day mission in LEO are therefore estimated to be between 14 and 28 rem at solar minimum and between 7 and 14 rem at solar maximum.

Transfer Ellipse

Dose calculations nave been made for the 5.25 hour transfer ellipse from LEO to GEO. These results vary between 1.0 rad (Hardy, 1979) due primarily to protons, and 0.018 rad (Stassinopoulos, 1979) due primarily to bremsstrahlung. As is the case for LEO, these are the estimates for the dose equivalents in rem as well. The large difference is the result of different assumptions in the trajectory made for the calculations.

Geosynchronous Orbit

In GEO, a majority of the absorbed dose is due to bremsstrahlung produced by the trapped electrons. An estimate of the dose equivalent for a worst case exposure from predictable radiation is 0.43 rem per day inside an aluminum sphere with radius 8 g/cm^2 (Seltzer. 1980).

The contribution to the dose equivalent by the galactic cosmic rays (GCR), particularly from the heavy charged particle component, may be important. The fragmentation characteristics of the heavy particle component must be considered for an accurate estimation of the absorbed dose from GCR (Wilkinson and Curtis, 1972). The quality factors (Q) necessary to convert absorbed dose to dose-equivalent are generally unknown for these radiations. Using a Q of 3 for the GCR, independent of depth, yields a very rough estimation of the dose equivalent as a function of depth. Use of an average value for Q of 3 for the galactic cosmic ray contribution is consistent with current recommendations (ICRP 26, 1977). Q values are under continual reassessment, and it is

possible the Q value could be increased for the HZE particle contributions. This Committee believes that an average Q of 3 may not be conservative for carcinogenic or mutagenic effects produced at low doses and low dose rates over long periods. This is a simplified model and more must be learned about the variation of Q with dose and depth for the GCR before a more accurate value for the dose equivalent can be obtained.

The shielding of manned space vehicles against the proton and helium ion component of the GCR has been examined (Santoro et al., 1973). The absorbed dose and dose equivalent due to primary particles and to secondary particles arising from the self-shielding of the tissue sphere are little affected by increasing the sphere radius from 5 to 20 g/cm² of aluminum. Protons and helium ions and their associated secondary radiations contribute about 55 to 65 percent to the total galactic cosmic ray dose and a smaller percent to the total dose equivalent. The dose and dose-equivalent rates for the combined proton and helium ion components average 31 mrad per day and 73 mrem per day, respectively, yielding an average Q of 2.4. This value of Q is a lower limit when the HZE particle contribution is very small. If the Q of 2.4 is the value for the proton and helium ion components, the average for all components may be higher than the Q value of 3 assumed above.

Solar particle events (SPE) will be a major hazard in GEO; special shielding—a storm cellar—is needed for space workers during a solar particle event. The dose received from a given SPE will depend upon the size of the SPE, the length of time of the warning before the particle buildup, the time required for the workers to get inside the storm cellar and the storm cellar shielding thickness. The size and time of occurrence of an SPE are not currently predictable. known that a correlation exists with the sunspot number and that the event frequency has an 11-year cycle. Within this cycle, there is a three- to five-year period that is almost event free. During the remaining six to eight years of the period, there is about a 40 percent probability of a large SPE each year. The total dose from the solar particles within the 11-year cycle is generally dominated by the contribution of the largest event within the cycle. This makes the accuracy of the prediction of size of an event that is about to occur, or is just starting, very important because of the special precautions which must be taken.

For a 30 g/cm² tissue sphere, the Wilson and Denn (1976) calculations provide a dose equivalent of an additional 2.5 rem for an SPE with the size and energy spectrum of the August 1972 event and 25 rem for the February 1956 event (Fig. 1). Webber (1963) made a similar estimate for the 1956 event. Rossi and Stauber (1977) estimated the uose equivalent behind 40 g/cm² of aluminum (equivalent to about 30 g/cm² of tissue) to be 25 rem for the August 1972 event, a factor

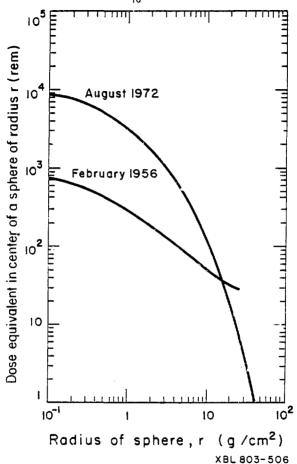


Figure 1. Dose equivalents from two major solar particle events (SPE) plotted as a function of the radius of tissue equivalent sphere. Adapted from Wilson, 1979.

of ten greater than the Wilson and Denn (1976) calculations. These differences in dose calculations for the 1972 event are still to be resolved.

Total 90-Day Dose Equivalent

Daily dose equivalents in the three phases of the SPS mission are summarized in Table 1. From the table, the best estimate of dose equivalent from predictable radiation sources at the worst parking orbit is approximately 40 rem for 90 days in GEO, assuming a Q value of unity. This dose equivalent value is derived from the calculations of Seltzer (1980) for a geostationary orbit with an altitude of $35,790~{\rm km}$, an inclination of 0° , and a parking longitude of $160^{\circ}{\rm W}$, the worst condition for radiation exposure to trapped electrons (Stassinopoulos, 1980). The incident electron spectrum was integrated for the epoch 1979.0 using the AEI7-HI environmental model. This model is based on a recent compilation of trapped electron data yielding conservatively high average values for the flux densities. A total spherical shield of 8 g/cm 2 aluminum equivalent material (3 g/cm 2 of spherical spacecraft shielding plus 5 q/cm² effective body selfshielding) is assumed to obtain the above value of the absorbed dose in water. The shape of the dose vs depth curve is such that a 50 percent change in the total shield assumed will affect the volume of the dose by a factor or two.

The largest contribution to the dose equivalent is the bremsstrahlung. In addition, there is about a 10 percent probability of an additional 2.5 rem from solar particle events, and a smaller probability that the SPE dose equivalent might be as high as 25 rem. A smaller contribution will be made by GCR. Therefore, the worst case 90-day mission dose equivalents will most likely be within the range of 40 rem (no SPE) to 65 rem (with large SPE). The Committee emphasizes that these are preliminary dose-equivalent estimates with large uncertainties.

ASSESSMENT OF RISK FOR CANCER INDUCTION

This SPS committee considered radiation-induced cancer the major health risk associated with exposure to ionizing radiations at dose levels most likely to be encountered in the SPS space environment. The risks can be calculated based on the whole-body exposure and a linear-quadratic dose-response model (NAS-BEIR, 1980), and average career dose-equivalent values may be used. As an illustrative example,

fable 1. Estimate to se equivalents at the center of an aluminum sphere of 8 g/cm 2 dius for the three phases of SPS*.

| Mission phase | Dose equivalent (rem) | Reference |
|--|-----------------------|---|
| LEO: | | |
| Average daily dose equivalent at solar minimum | 0.15 0.30 | A. Hardy, 1979 Stassinopoulos, 1979 modified to spherical shields (Seltzer, 1979) |
| TE: | | |
| Average one way trip from LEO to GEO | -1 -0.2 | Hardy, 1979 Stassinopoulos, 1972 modified to spherical shields (Selczer, 1979) |
| GEO: | | |
| Average daily dose equivalent at solar minimum (excluding solar particle events) assuming a "worst" case longitude of 160 W. | 0.43 | Seltzer, 1979 |

 $^{^{\}star}$ A Q of unity is assumed for the conversion from the absorbed dose to dose equivalent.

among 10,000 SPS workers from the general population of all age groups, about 1,640 persons would be expected to die of cancer, in the absence of any additional radiation exposure. Based on the dose-equivalent estimate in GEO for the Reference System design (40 rem per 90-day mission) and ten missions per career (accumulated dose equivalent of 400 rem), the linear-quadratic dose response model predicts increases of between 20 and 60 percent depending on the projection model (about 160 to 1,000 excess cancer deaths). The linear dose-response model predicts values about two times larger.

Such cancer risk predictions are subject to a large number of uncertainties, as outlined in this report. In spite of these uncertainties, this SPS Committee concludes that the increased potential in cancer-induction risk due to radiation exposure in the SPS environment, as presently envisaged with the present Reference System design, can be substantially reduced.

There is also the possibility that this radiation might increase genetically-related ill health, developmental abnormalities in the newoorn, lens cataracts, and temporarily decrease fertility. If the radiation dose is substantially reduced, as suggested above for the purpose of reducing cancer incidence, the probability of these other health risks will also be reduced.

Additional details regarding the space radiation environment and radiation health effects are given in Appendices A and B.

3. CONCLUSIONS AND RECOMMENDATIONS

HEALTH RISK ASSESSMENTS

for ionizing radiation, the major concern will be late or delayed health effects, particularly the increased risk of radiation-induced cancer. The estimated lifetime risk for cancer is 0.8 to 5.0 excess deaths per 10,000 workers per rad of exposure. Thus, for example, in 10.000 workers who completed ten missions with an exposure of 40 rem per mission, 320 to 2,000 additional deaths in excess of the 1640 deaths from normally occurring cancer, would be expected. These estimates would indicate a 20 to 120 percent increase in cancer deaths in the worker-population. The wide range in these estimates stems from the choice of the risk-projection model and the dose-resonnse relationship. The choice between a linear and a linear-quadratic dose-response model may alter the risk estimate by a factor of about two. The method of analysis (e.g., relative vs absolute risk model) can alter the risk estimate by an additional factor of three. Choosing different age and sex distributions can further change the estimate by another factor of up to three. When decisions have been made about the selection of SPS workers, the precise influence of age and sex distribution can be included in later risk estimates. However, the choice of dose-response relationship and projection models is a matter of opinion and will not be resolved scientifically for quite some time.

CONCLUSIONS

The conclusions of the committee based on the findings above are:

- The risk of excess cancer deaths will be closer to the lower limit estimated above because the major exposure will be from low dose rate, low-LET irradiation. This being the case, we consider a reasonable estimate to be one excess death per 10,000 workers per rem of exposure. If this level of risk is applied to the worst case reference system exposure level, namely 40 rem/mission, there would be 400 excess cancer deaths in a work force of 10,000 completing ten missions (accumulated dose equivalent of 400 rem).
- The potential genetic consequences, which vary with the population age distribution, could be of significance. At the present time, sufficient information on the age and sex distribution of the worker population is lacking for precise estimation of impact.
- Similarly, the radiation exposure of a pregnant worker could lead to developmental abnormalities in the empryo.

- 4. Based on the Reference System (estimated 400 rem from ten missions), dose to space workers from low-LET bremstrahlung approaches the cataractogenic level for man. The appropriate quality factor of the HZE particle portion of the dose is unknown at this time. If its Q is greater than 20, the cataract hazard may be significant. More information is needed about the cataractogenic risk of exposure to high-LET radiations.
- 5. With adequate information on the radiation environment, with well-designed areas protected from the increased levels during solar particle events, and in the absence of a radiation accident or some other unexpected situations (e.g., nuclear detonation), there will be no early or acute radiation health effects occurring in the SPS worker population.
- No other radiation health effects are considered to be of sufficient consequence to be important for risk estimation.

RECOMMENDATIONS

The SPS Committee strongly emphasizes the need to reduce the uncertainties in the evaluation of radiation health risks to SPS workers in the space environment. It recommends that the following be carried out to achieve this goal:

- A radiation environment model, appropriate to this SPS mission, should be developed for study and simulation. The model should include short-term and solar-cycle variations. The short-term variations of the radiation dose rate in space must be better understood so that the range of doses and dose rates to be expected can be established accurately.
- An instrumented research satellite should be placed in GEO to measure absorbed dose rate and particle spectra at depth in phantoms and to measure the temporal variations of the radiation field.
- 3. The differences in the results of dose and dose-rate estimation obtained from the shielding transport codes must be evaluated. The use of different calculational methods with the same set of assumptions should yield the same results.
- When institutional decisions have been made to develop appropriate exposure strategies, engineering decisions for dose control should then be made.
- Radiation shielding of transport vehicles, work stations, habitats, and space suits should be designed to achieve minimal radiation exposure levels. The use of laminar layering, where

possible, to reduce the dose and dose rate from bremsstrahlung to the lowest practical values provides a practical example.

- 6. Studies should be initiated promptly to define the space worker population profiles, particularly the age and sex distributions. Age-specific deaths rates from all causes, in general, and cancers, in particular, should be obtained for a worker population composition similar to that which will work in space. These studies on the SPS worker populations will then provide the data base for calculations of potential health risks of ionizing radiations in the space environment.
- 7. The biological effects of the HZE particle radiations must be investigated in detail in order to determine appropriate quality factors and risks from lesions unique to HZE. Interactions between high- and low-LET radiation that could occur in the space radiation environment should be considered, particularly the possibility of synergistic effects on carcinogenesis or mutagenesis.
- 8. The health effects of exposure to low-level ionizing radiations must be considered in the context of the potential health effects of other physical and chemical agents in the space environment. Such competing effects may interact with other host or constitutional factors to mask, enhance, or diminish any radiation health effects, such as cancer. Environmental factors to be considered include cabin atmosphere, temperature and pressure, nutrition, non-ionizing radiation and weightlessness.

This SPS Committee concludes that the radiation environment estimated for the Reference System represents a health risk to SPS workers. However, the Committee emphasizes that none of the problems identified are considered sufficiently intractable to preclude achieving a minimal risk to SPS workers.

A number of areas have been considered which we believe impact on the potential health effects on workers in space. These include, for example: the biological effects of HZE particles; the effects of environmental agents in space, other than ionizing radiation, which may affect the radiation health effects; and the RBE/LET relationships. However, much more needs to be known about these factors before they may be used to improve the accuracy of quantitative estimation of health risks in space. The Committee has chosen not to include these uncertainties in its estimations at the present time but urges increased studies in these areas to provide greater precision for future SPS health assessments.

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APPENDIX A

SPACE RADIATION ENVIRONMENT

The preliminary SPS Reference System (DOE/NASA, 1978) calls for workers to move from Earth to low Earth orbit (LEO) (altitude 500 km) for stays of varying lengths of time. Workers then travel to geosynchronous Earth orbit (GEO) (altitude about 36,000 km), following an elliptical trajectory (transfer ellipse, TE). These three SPS Environments—LEO, TE, and GEO—have ionizing radiations of different quality, time dependence, and predictability of dose levels. The various components of the radiation environment are described in this section, and those important to each SPS stage are identified.

TRAPPED ELECTRONS

Large flux densities of electrons, trapped in the earth's magnetic field, are contained in an inner and outer zone separated roughly by the magnetic shell parameter,* L = 2.8 earth radii. The low energies of electrons in the inner zone are important only if persons are protected by very thin shielding ($<0.5~g/cm^2$ aluminum) (Stassinopoulos, 1979). The outer zone contains flux densities of electrons which are greater in magnitude than those of the inner zone and have a larger fraction of high-energy particles. Maximum flux densities occur in the region of about 3.5 earth radii (approximately 21,300 km); the trapping region extends out to about 12 earth radii (76,540 km). Space venicles transporting workers and materials from LEO must travel through the heart of the outer electron zone to reach GEO which, at 6.6 earth radii, is well within the outer zone of trapped electrons. Figure A-1 illustrates the spatial variation of trapped electrons as a function of altitude and geographic latitude, the positions of LEO and GEO, and a representative pass for the transfer trajectory (TE).

Two large temporal variations of the outer zone electron flux densities at the position of a GEO satellite have been identified (Stassinopoulos, 1980):

1. Diurnal variations—At GEO, electron flux densities vary over factors between 6 and 16 (depending on L value) between day and night. The maxima occur at about 1000 to 1100 hours, and the minima at about 2200 to 2300 hours, local time. The extent and times of these extremes also depend on electron energy and, to a small degree, on position in the eleven-year solar cycle.

^{*}The magnet shell parameter, L, denotes roughly a geomagnetic field line. The value gives the approximate geocentric distance, in earth radii, of the intersection of the field line with the geomagnetic equatorial plane.

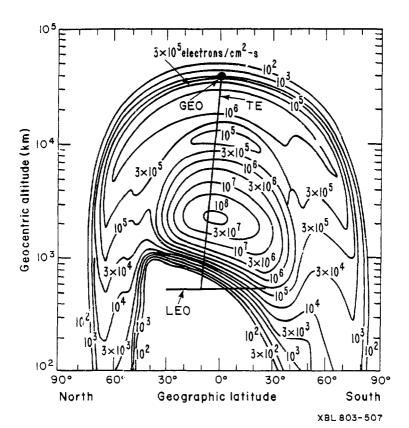


Figure A-1. The spatial variation of trapped electrons, plotted as a function of altitude and geographic latitude. The positions of LEO, GEO, and a representative pass for the transfer trajectory are illustrated. Adapted from Wilson, 1979.

 Short-term ennancements—Electron flux densities can increase markedly due to intermittent changes in the solar plasma associated with small substorms. Flux densities may rise two to three orders of magnitude in several hours followed by a decay lasting several days (Stassinopoulos, 1980).

BREMSSTRAHLUNG

In general, the absorped dose from bremsstrahlung is not dominant behind thin shielding (e.g., <3 g/cm2 aluminum) because of the overriding preponderance of primary electrons. However, at greater depths, the dose from electrons drops sharply, dependent largely on the shape of the incident electron energy spectrum. This causes the bremsstranlung to dominate the absorbed dose at large shielding thicknesses. (e.g., >3 g/cm² aluminum). For inner zone electrons at LEO, the premsstranlung dose is completely dominated at all thicknesses by the dose from trapped protons (see below). Here, the bremsstrahlung dose is sufficiently small to be negligible. For outer zone electrons at GEO, on the other hand, the premsstrahlung dose dominates benind shielding of 3 q/cm² aluminum or greater thickness (Stassinopoulos. 1980). Thus, the premsstrahlung dose is an important component of the radiation environment in GEO and in the transfer ellipse between LEO and GEO, taking into account certain assumptions of trajectory and vehicle speed.

TRAPPED PROTONS

Protons are also trapped in large numbers by the geomagnetic field. A region of the trapped proton zone dips close to the earth in the southern Atlantic Ocean, southeast of the Brazilian coast. This region, called the South Atlantic Anomaly, is the most important contributor to the radiation environment for space workers in LEO. A spacecraft in LEO, nowever, will pass through this region on only 60 percent of its revolutions. Stassinopoulos (1979) has estimated that 36 percent of the time in LEO is flux-free, i.e., in a radiation environment of less than one particle/cm²-sec of electrons with energies greater than 0.5 MeV and protons with energies greater than 5 MeV. Each day, for a maximum duration of ten hours, there are about six consecutive flux-free revolutions without danger of increased radiation exposure.

Trapped protons may be important in the TE between LEO and GEO, depending on the trajectory and venicle speed selected. Two different calculations of the proton environment encountered by a transfer vehicle have been made, assuming proton environments varying by several orders of magnitude (see "Transfer Ellipse," p. 3). In GEO, the trapped protons are of such low energy that they will not present a health risk to workers at nominal snielding thicknesses.

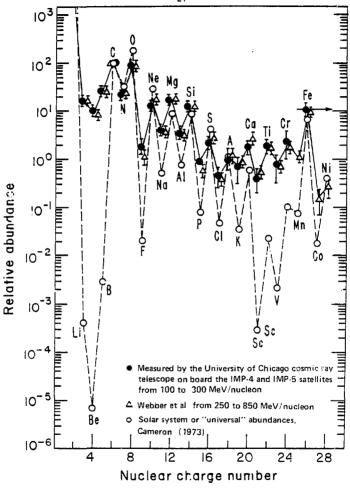
GALACTIC COSMIC RAYS

Galactic cosmic radiation consists primarily of hign-energy nuclei with origins outside our solar system; these particles appear to pervade regions outside our magnetosphere isotropically. Approximately 38 percent are protons, 10 percent are helium nuclei, 1 percent are electrons, and 1 percent are heavy nuclei (Z > 2). Ratios of abundances relative to carbon (Z = 6) are shown for the heavy nuclei as a function of nuclear charge number or Z in Fig. A-2. If each nuclear species has a similarly shaped energy spectrum, the relative absorbed dose in free space in a small sphere of tissue from each species would be the GCR abundance ratios scaled by Z^2 . Due to differences in the spectral shapes at low energies, there are variations noted from the simple Z^2 scaling. However, the difference at low energies is important only behind thin shielding.

Table A-l presents three calculations of the estimated absorbed dose rate in a small tissue sphere in free space and the relative composition from particles grouped in broad ranges of Z (Curtis, 1974). The first two calculations (Curtis and Wilkinson, 1968; Schaefer, 1968) were obtained by integration of the appropriate energy spectra. The third calculation was based on experimental data from a balloon at high altitude (Anderson, 1968). The daily dose rate of between 30 and 40 mrad/day is the maximum expected, since the spectra used were appropriate for solar minimum when the galactic flux densities are known to be the highest. The daily absorbed dose within the body will be less than this value because, although the dose from the orotons will not decrease with depth, there will be a decrease in the contribution from the heavier components (see Section 2).

Curtis (1973) calculated that the flux densities of HZE particles with LET > $100~\text{keV/}_\text{um}$ decrease from 8 to 3 cm² hr as the shielding is increased from 1 to 10 g/cm² of aluminum. Approximately 5 percent of the particles with charge greater than two have charges in the range of 20–26, and this percentage does not vary significantly with thickness of shielding.

Measurements made on Soviet satellites Prognoz 1 and Prognoz 2 over an eight month period in 1972, during quiet radiation conditions outside the magnetosphere, yield an estimated dose-rate of about 24 mrad/day (Logacnev et al., 1974). The estimated dose rate is about 10 mrad/day at solar maximum. Data from U.S. satellites Mariner II and Mariner IV from late 1962 to 1964 provide estimated dose-rate values of 30 to 45 mrad/day at solar minimum. Pioneer IV measurements correspond to dose rates of about 9.6 and 24 mrad/day, the latter at solar minimum (Janni and Holly, 1969). Pioneer V data yield an estimated dose rate value of about 14.4 mrad/day at solar maximum. These estimates are summarized in Table A-2.



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Figure A-2. Comparison of the abundances of the elements in galactic cosmic rays with the solar system abundances. Adapted from Wefel, 1979.

Table A-1. Primary galactic cosmic-ray dose rates at solar minimum*

| Reference | Dose-rate (mrad/day) | Charge (Z) | Relative Composition (% of total dose) |
|----------------------------------|-------------------------|---------------------------------|---|
| Curtis and Wilkinson, 1968 | 34.5 | 1 2 6-9 10-14 26-28 | 37 28 15 10 |
| Schaefer, 1968 | 30.6 | 1 2 6 10 20 | 40 24 15 9 12 |
| Anderson, 1968 | ~42.5 ± 4.3 | 1 2 > 2 Electrons | 32 2. 38 5 |

^{*} Curtis, 1974.

Table A-2 Comparison of measured radiation dose-rates in space satellites.

| Satellite | Dose Rate mrad/day | Comment | | | |
|-------------------------------------|-----------------------|---|--|--|--|
| Prognoz 1 ^a Prognoz 2 | 24 | Measured at quiet radiation conditions outside magneto-sphere | | | |
| | 10 | Estimated at solar maximum | | | |
| Mariner II Mariner IV | 30–45 | Measured at solar minimum | | | |
| Pioneer IV ^b | 9.6 24 | Measured at solar maximum Measured at solar minimum | | | |
| Pioneer V | 14.4 | Measured at solar maximum | | | |
| a Logacher et al., 1974 | | | | | |
| D Janni and Holly, 1969 | | | | | |

SOLAR PARTICLE EVENTS

Giant solar particle events (SPE) are caused by large upheavals on the solar surface which accelerate protons and, to a lesser extent, heavier nuclei to high energies. These particles are then transported through the solar magnetic field and can increase radiation levels to high values for several hours or days in the vicinity of the earth outside our magnetosphere. GEO, at 6.6 earth radii, is in such minimal geomagnetic shielding that particles down to very low energy arriving in the vicinity of the earth will be able to reach it.

Large events occur with highest probability during the rising or falling portions of the eleven-year solar activity cycle. A very small number of events have dominated the total fluence of particles arriving in an eleven-year period. Figure A-3 shows that the events occurring in August 1972 dominated the total fluence arriving in cycle 20 and compares these events with others which occurred in 1967, 1969, and 1971. In Figure A-4, two different estimates of the dose equivalent at the center of a sphere of radius r from the August 1972 event are plotted as a function of the radius; a discrepancy exists between the two calculations, and this is still to be resolved. The figure shows, nowever, that the sphere radius must be at least 10 q/cm² of tissue before the dose equivalent drops below 100 rem. The occurrence of such large events, therefore, results in high dose-rate exposures which will be a considerable radiation hazard to SPS workers in GEO. For workers in LEO, on the other hand, enough shielding is provided by the geomagnetic field to make dose equivalents from solar events negligible.

DIFFERENT EFFECTS OF HIGH- AND LOW-LET RADIATIONS

In the evaluation of the potential health effects caused by space radiations, the radiations may be divided into general lategories based on linear energy transfer (LET) or collision stopping power. Low-LET radiations, such as nigh energy electrons and protons, have ionizations which are relatively far apart, with only a small probability of interaction between the ionization products created by a single ionizing particle. Those interactions which do occur are primarily the result of multiple particle tracks. High-LET radiations, such as heavy charged particles, are characterized by ionizations which are normally more closely spaced, and there is therefore a correspondingly greater probability of interactions between the ionization products created by the passage of a single particle. This difference in the microscopic dose distribution generally causes the higher LET radiations to have a greater biological effectiveness.

QUALITY FACTOR

A quality factor, Q, has been defined to account for the varying degrees of potential adverse health effects in man caused by different

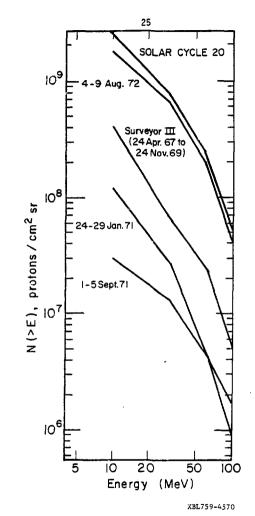
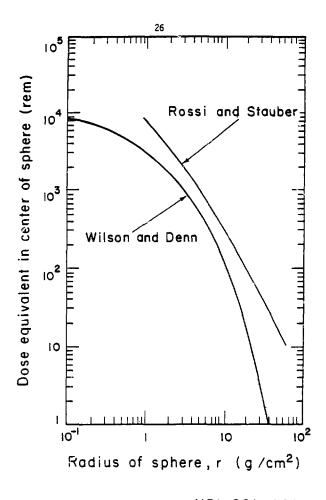


Figure A-3. Degree of domination of total fluence arriving in cycle 20 by solar events which occurred on 4-9 August 1972. Adapted from Crawford et al., 1975.



XBL 801-161

Figure A-4. Dose equivalents from the August 1972 solar event at the center of a sphere as calculated by Rossi and Stauber (1977) and Wilson and Denn (1976).

ionizing radiations of different LET (ICRP 26, 1977). Currently recommended quality factors very from 1 (for radiations with LET equal to or less than 3.5 keV/ μ m) to 20 (for radiations with LET equal to or greater than 175 keV/ μ m). These values have been selected on the basis of relevant low dose and/or low dose—rate RBE values. They have guided our estimation of a quality factor for the galactic cosmic radiation. Q values are used for assessing radiation health effects for radiation protection purposes, and not to assess the potential health effects of severe accidental high-dose exposures.

There may be some health effects caused only by HZE particles and not by other radiations, and here the conventional use of a quality factor does not provide an assessment of potential health risks.

APPENDIX B

RADIATION HEALTH EFFECTS IN THE SPACE ENVIRONMENT

Radiation doses and dose rates which may be encountered in space present potential health risks to space workers. This section discusses the biological and health effects of the different radiations encountered in the space environment and is based upon data currently available from human epidemiological surveys, laboratory animal studies, and space experimentation. The effects on health are considered in two general categories: early or acute effects, and late or delayed effects. Early effects are those occurring within hours, days, or a few weeks following high-dose, whole-body exposure. Late or delayed effects usually occur months to years following exposure and include cancer-induction, developmental abnormalities in the newborn, genetically related ill-health, lens cataracts, shortened lifespan, and impairment of fertility. The special problems of HZE-particle induced health effects are also discussed.

EARLY EFFECTS

Early radiation health effects assume clinical significance only with whole-body dose equivalent greater than 150 rem received in relatively short time periods (minutes to hours). Such exposure levels are likely to be encountered only during major solar particle events or during nuclear detonations in space.

Radiation is similar to other potentially hazardous physical or chemical agents in that high doses can produce tissue and organ injury, illness, and possibly death. The principle site of biological action of ionizing radiation is the proliferating cells of the renewal system of the organism, such as the bone marrow and intestinal epithelium and spermatogonia. Within these cell renewal populations, the most sensitive cell is the progenitor or stem cell. When the supply of functional cells is temporarily disrupted, the result is impaired function of that tissue or organ and potentially serious injury to the individual. If the therapeutic measures are inadequate or regeneration of the depleted cell population does not occur soon enough, the individual may die. Dawage to bone marrow and the intestine may cause death within 7 to 60 days after acute exposure to whole-body dose of radiation greater than a few hundred rad of low-LET radiation.

Cellular depletion in tissues and organs that contain large numbers of dividing cells, e.g., in the bone marrow, lymphopoietic tissues, and testis, can be detected at doses as low as 50 rad delivered in short time periods, and are readily evident at doses of 100 to 150 rad. Death does not occur at these dose levels, elthough severe cellular depletion can be observed in some non-vital renewing organs, such as the testis. Whole-body radiation doses in the 200- to 400-rad

range causes severe bone marrow depletion leading to symptoms related primarily to a decrease of circulating neutrophils and platelets in the blood. This results in the bone-marrow syndrome. The signs and symptoms are infection in a variety of body tissues due to neutropenia, severe bleeding in tissues and organs due to thrombocytopenia, and possibly death within 20 to 40 days after exposure. Effective treatment consists of neutrophil transfusions, protection of the individual against infection, and fresh platelet transfusions to control bleeding. Replacement of hematopoeitic stem cells by bone marrow transfusions from compatible donors may also provide effective therapy after exposure to doses where bone-marrow damage is a lethal threat, (Bond et. al., 1965; Mathe and Schwarzenberg, 1979).

LATE EFFECTS

Consideration here is given primarily to those delayed or late health effects in human beings following exposure to low-LET radiation. x-rays, and damma rays from radioactive sources since these are the ionizing radiations most often encountered on Earth (e.g., in the nuclear industry and in medicine) and about which we have sufficient human data. Little consideration is given to high-LET neutrons and alpha particles. Briefly, low-level radiation can affect the cells and tissues of the body in three important ways. First, if the lesion occurs in one or a few cells, such as those of the hematopoietic tissues, the irradiated cell can occasionally transform into a cancer cell, and after a period of time, there is an increased risk of cancer developing in the exposed individual (carcinogenesis). Second, if the embryo or fetus is exposed during gestation, injury can occur to the proliferating and differentiating cells and tissues, leading to abnormal growth and development (teratogenesis). Third, if the lesion occurs in the reproductive cells of the testis or the ovary, the genome of the germ cell can be altered, and the injury can be expressed in the descendants of the exposed individual as genetically-related ill nealth.

There are a number of other biological effects of ionizing radiation, such as cataracts of the lens of the eye, or impairment of fertility, but the three important late effects—carcinogenesis, mutagenesis and teratogenesis—stand out as those of greatest concern. A considerable amount of information is available from epidemiological studies of exposed human populations and from laboratory animal experiments. Furtnermore, we believe that exposure to ionizing radiations, even at very low levels of dose, carries some risk of such deleterious effects. As the dose of radiation increases above very low levels, the risk of these deleterious health effects increases in exposed human populations. Because some risk is present at any level of exposure the exposure must be kept to a minimum commensurate with accompliring the mission.

A number of observations on the late health effects of low-level radiation have been made about which there is reasonably good general agreement. These observations are based primarily on careful evaluation of epidemiological surveys of exposed human populations, on extensive research in laboratory animals, on analysis of dose-response relationships of carcinogenic, teratogenic and genetic effects, and on known mechanisms of cell and tissue injury in vivo and in vitro (NAS-BEIR, 1972, 1977, 1980; UNSCEAR, 1977; NCRP, 1980).

- 1. Cancer is the most important late somatic effect of low-dose ionizing radiation. Different organs and tissues appear to vary greatly in their relative susceptibility to cancer-induction by radiation. The most frequently occurring radiation-induced cancers in man include, in decreasing order of susceptibility: the female breast; the thyroid gland; the hematopoietic tissues; the lung; certain organs of the gastrointestinal tract; and bone. There are influences, however, of sex, age at the time of irradiation, and age at the time of expression of the disease.
- 2. The effects of growth and development in the irradiated embryo and fetus are related to the gestational stage at which exposure occurs. It appears that a threshold level of radiation dose and dose rate may exist below which gross teratogenic effects will not be coserved. However, these dose levels would vary greatly depending on the particular developmental abnormality and on the radiation types and qualities.
- 3. Estimations of the radiation risks of genetically-related ill-nealth are based essentially entirely on laboratory animal observations, primarily from aboratory mouse experiments, because of the lack of data on exposed human populations. Scientific knowledge of fundamental mechanisms of radiation injury at the genetic level is far more complete than, for example, of mechanisms of radiation carcinogenesis, thereby permitting greater assurance in extrapolating information on genetic mutagenesis from laboratory animal experiments to man. Mutagenic effects are related linearly to radiation dose even at very low levels of exposure at low dose rates.

In spite of a thorough understanding of these late health effects in exposed human populations, there is still considerable uncertainty about the potential delayed health effects of low-level radiation. These uncertainties are discussed in the sections that follow.

CANCER

Cancer is the second leading cause of death in the general population in the United States. It is estimated that about 25 percent of the U.S. population will eventually develop cancer (not including non-melonoma skin cancer) and that about 16.5 percent of the population will eventually die of cancer.

Cancer is expected to be the most important delayed somatic effect in workers exposed to radiation in the space environment. Radiation causes an increase in the cancer risk after whole-body exposure to low LET radiation at dose equivalents greater than 50 to 100 rem. At lower levels, it is difficult or impossible to demonstrate an increased risk even from epidemiological surveys of large populations exposed to low-level radiation. Risk estimates of possible cancer-induction at low-dose levels must necessarily depend more on the dose-response models and on the projection models used in the risk coefficient calculations than on the epidemiological data themselves. The risk of cancer from low-dose, low-LET whole body radiation is a problem of detection of statistically increased incidence of the disease over the normal expectation in the population in the absence of any additional radiation exposure.

Radiation-induced cancer may, as at least an initial step, involve random changes in DNA. Hence, late effects occur only on a chance basis in the relatively rare individual (i.e., almost all survivors of high-dose exposure will live out their normal life span). For example, in the 1974 ABCC study in Hiroshima and Nagasaki of the 70,000 deaths from all causes among 285,000 Japanese atomic bomb survivors exposed at all dose levels, it is estimated that approximately 400 to 500 persons, or 0.7 percent, nad died from radiation-induced cancer (Finch, 1979).

Cancers induced by radiation are indistinguishable from their normally-occurring counterparts; cause and effect cannot be established with certainty on an individual basis. This, together with the low overall induced cancer incidence rate even at high doses, make it difficult or impossible to detect directly or define the carcingenic effects of low dose equivalents of radiation (in the range of a few rem or less).

Although the cancer risk may be increased at the time of irradiation, a cancer induced does not begin to appear until after a minimal latent period has passed. The duration (or plateau) of the apparent radiation—induced increase in incidence of solid cancers (i.e., cancers other than leukemia), which tend to have long latent periods, are not

yet completely known. On the other hand, the duration of apparent radiation-induced increase in incidence of leukemia, which has a relatively short modal latency, is better known.

Uncertainties in Dose-Response Relationships for Radiation-Induced Cancer

A general hypothesis for estimation of excess cancer risk in irradiated human populations, based on theoretical considerations, on extensive laboratory animal studies, and on limited epidemiological surveys of exposed human populations, suggests various and complex doseresponse relationships between radiation dose and observed cancer incidence (NAS-BEIR, 1980). Models, with increasing complexity, include the linear, the pure quadratic, the quadratic (with a linear term), and finally, the multicomponent quadratic form with a linear term and with an exponential modifier (Figure 8-1). One of the most widely considered models for cancer-induction by radiation, pased on the available information and consistent with both knowledge and theory, takes the complex quadratic form: $I(D) = (a_0 + a_1D + a_2D^2) \exp(-a_1D - a_2D^2)$, where I is the cancer incidence in the irradiated population at radiation dose D in rad, and an, al, al, al, and all are coefficients (Figure 8-1). This multicomponent dose-response curve contains (1) initial upward-curving linear and quadratic functions of dose, which represent the process of cancer-induction by radiation; and (2) a modifying exponential function of dose, which represents the competing effect of cell-killing at high doses. on is the ordinate intercept at zero dose, and defines the natural incidence of cancer in the population. at is the initial slope of the curve at zero dose. and defines the linear component in the low-dose range. as determines the upward-curving quadratic function of dose. 81 and 82 determine the clope and curvature of the downward-curving function in the high-dose range, and define the cell-killing function. In the case of epidemiological surveys, this complex general dose-response form cannot be universally applied. Therefore, the model is simplified by eliminating the parameters which have the least effect on the form of the goseresponse relationship in the dose range of low-level radiation.

The present SPS Committee believes:

1. Some experimental and human data, as well as theoretical considerations, suggest that for exposure to low-LET radiation, such as x-rays and gamma rays, at low doses, the linear model probably leads to parestimates of the risk of most radiation-induced cancers in man. Generally, data from high dose exposures in man and animals are used for estimating risk coefficients for the various models. To the extent that a quadratic term plays a role in actual response, nigh dose data extrapolated linearly to zero will tend to overestimate a, the linear term. The use of data in this fashion can certainly be used to define the upper limits of risk.

SHAPES OF DOSE RESPONSE CURVES

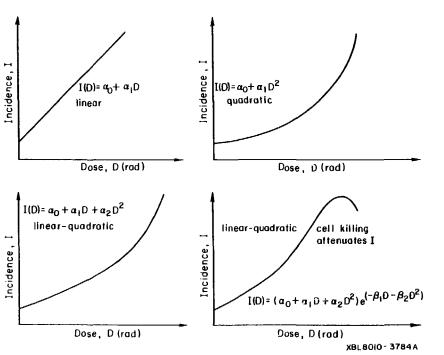


Figure B-1. Shapes of dose response curves (NAS-BEIR, 1980).

- 2. The pure quadratic model may be used to define the lower limits of risk from low-dose low-LET radiation.
- 3. For exposure to high-LET radiation, such as neutrons and neavy charged particles, the linear model is considered to be the most appropriate (NAS-BEIR, 1980).
- 4. For low dose and/or dose-rate exposure where reliable human data do not exist, the evidence from laboratory animal data indicates that a more accurate excess cancer risk may be smaller by a factor of perhaps two or more than that estimated using the linear model (NRC, 1975; UNSCEAR, 1972 and 1977; NCRP, 1975 and 1980). This reduction, however, would probably apply only to the low-LET fraction of the total dose.

Estimation of Space Radiation Induced Carcinogenic Risk in Man

Additional poorly understood factors affect the quantitative estimation of carcinogenesis at low-dose radiation. These include: the length of the latent period; the RBE for high-LET radiations, particularly for fast neutron and heavy charged-particle radiation relative to gamma and x-radiation: the special health effects of HZE particles: the period during which the radiation risk is expressed; the model used in projecting risk beyond the period of observation; the effect of dose rate or dose fractionation; and the influence of differences in the natural incidence of specific types of cancer. In addition, uncertainties are introduced by the biological risk characteristics of numans. for example, the effect of age at irradiation, the influence of any disease for which radiation was given therapeutically, and the influence of length of observation or follow-up of the study populations. The collective influence of these uncertainties limits the precision with which estimates of human cancer risk can be hade for radiation exposure in the space environment.

The &EIR-III Report (NAS-BEIR, 1980) is used here as the source for cancer risk estimation. The chief sources of epidemiological data used in that report are the Japanese populations exposed to whole-body irradiation in Hirosnima and Nagasaki, the patients with ankylosing spondylitis and other patients wno were exposed to partial body irradiation therapeutically or to diagnostic medical x-rays and the various occupationally-exposed populations, such as uranium miners and radium dial painters. Most epidemiological data do not systematically cover the range of low to moderate radiation doses for which the Japanese atomic bomb survivor data appear to be fairly reliable (NAS-REF, 1977; UNSCEAR, 1977). Analysis in terms of dose-response, therefore, necessarily rely greatly on the Japanese data. The substantial neutron component of dose in this shima makes this data of great value in estimating cancer risk for high-LET radiation. The Nagasaki data for which the neutron component of dose is small, are not reliable for

doses below 100 rad, but can be useful for estimating cancer risk from higher doses of low-LET radiation. The recent calculations of the BEIR III Report (NAS-BEIR, 1980) provides the most current basis for estimating the carcinogenic risk from radiation in SPS workers in the space environment. This report (NAS-BEIR, 1980) chose three exposure situations for illustrative computations of lifetime cancer risk of low-dose, low-LET, whole-body radiation:

- A single exposure of a representative (life-table) population to 10 rad;
- 2. A continuous, lifetime exposure of a representative (lifetable) population to 1 rad per year:
- 3. An exposure to 1 rad per year over several age intervals exemplifying conditions of occupational exposure.

These three exposure situations reflect circumstances that might normally occur—i.e., the general and worker populations and single and continuous exposure.

The selected level of chronic exposure of one rad per year, although only one-fifth the maximal permissible dose for conventional occupational exposure is nevertheless consistent with the occupational exposure experience in the nuclear industry. The United Status 1969–1971 life-table was used as the basis for the calculations (NAS-BEIR, 1980). The expression time was taken as 25 years for leukemia and the remaining years of life for other cancers. Two projection models were used, the absolute risk and the relative risk models. The <u>absolute</u> risk is the expression of excess cancer risk due to radiation exposure as the arithmetic <u>difference</u> between the risk among those in the exposed population and that obtaining in the population in the absence of radiation exposure. The <u>relative</u> risk is the expression of excess cancer risk due to radiation exposure as the arithmetic <u>ratio</u> of the risk among those in the exposed population to that obtaining in the population in the absence of radiation exposure.

In the absence of any increased radiation exposure, among one million persons of life-table age and sex composition in the United States, about 164,000 persons would normally be expected to die from cancer, according to present cancer mortality rates. For a situation in which these one million persons are exposed to a single dose increment of 10 rad of low-LET radiation, the linear-quadratic diserresponse model (LQ-L)* predicts increases of about 0.5 percent to 1.4 percent (about 750-2,300 cancer cases) over the normal expectation of cancer

^{*}Linear-Quadratic (LQ) for low-LET radiation, linear (L) for nign-LET radiation.

mortality, according to the projection model used. The linear model (L-L)** predicts increases of about 1 percent to 3 percent (about 1.700-5,000 cancer cases) over the normal expectation, depending on the projection model. The pure quadratic model (Q-L) increases of about 0.06 to 0.2 percent (about 100-300 cancer cases) over the normal expectation of cancer mortality, according to the projection model used. The calculations are summarized in Table B-1 taken from the BEIR III Report (NAS-BEIR, 1980). The upper and lower limits of these cancer mortality-risk estimates suggest a wide range of values which may differ by as much as an order of magnitude. The uncertainty derives mainly from the dose-response models used, from the alternative absolute and relative projection models, and from the sampling variation in the source data. The lowest risk estimates-the lower bound of the range-are obtained from the pure quadratic model: the highest—the upper bound of the range—from the linear model; and the linear-quadratic model provides estimates intermediate between these two extremes.

For continuous lifetime exposure to one rad per year, the increase in cancer mortality, according to the linear-quadratic dose-response model (LQ-L), ranges from about three percent to eight percent (about 5,000-13,000 cancer cases) over the normal expectation, depending on the projection model. The linear model (L-L) predicts increases of about 7 percent to 18 percent (about 11,000 to 30,000 cancer cases) over the normal expectation. The quadratic dose-response model (Q-L) cannot be applied to this calculation. The calculations are summarized in the BEIR III Report (NAS-BEIR, 1980).

The calculations for continuous exposure to one rad per year from the BEIR III Report (NAS-BEIR, 1980) segments the population into several age intervals, viz. 20-65 years, 35-65 years, and 50-65 years, exemplifying conditions of occupational exposure over a working career. For the increasing age groups, much of the variation in cancer risk estimates is due to the total dose received for the different periods of continuous exposure. The excess cancer mortality decreases with increasing age interval due almost completely to the total dose received rather than on the projection model used. For exposure at 35-65 years and 50-65 years, the two projection models give nearly identical results.

^{**}Linear (L) for low-LET radiation, linear (L) for high-LET
radiation.

^{***}Quadratic (Q) for low-LET radiation, linear (L) for high-LET radiation.

Table 8-1 Estimated excess mortality per million persons from all forms of cancer, single exposure to 10 rads of low-LET radiation, y dose-response model.*

| Dose-response model** | | | Absolute-risk projection model | Relative-risk projection model |
|-----------------------|-----------------|--|--------------------------------|-----------------------------------|
| Leukemia and bone | Other cancer | | | |
| | | Normal expectation of cancer death | 163,800 | 163,800 |
| LQL | LQ-L | <pre>£xcess deaths: number % of normal</pre> | 766 0.47 | 2,255 1.4 |
| LL | L-L | Excess deaths: number % of normal | 1,671 1.0 | 5,014 3.1 |
| Q-L | Q-L | Excess deaths: number % of normal | 95 0. 058 | 276 0.17 |

^{*} NAS-BEIR (1980)

^{**}See text.

GENETICALLY RELATED ILL-HEALTH

At low-level radiation exposure the effects of radiation include genetic disorders which depend on changes in individual genes (gene mutations) and in chromosomes (either in total number or in gene arrangement, or chromosomal aberrations). Low-level ionizing radiation causes genetic disorders (genetic mutations) in laboratory animals, and it is assumed that the same applies in man. These effects may be manifested in man as genetically related ill-health (e.g., mental retardation in descendants of irradiated parents). Gene mutations are expected to have greater health consequences than changes in chromosomes. The genetic effects arising from either are proportional to dose. In estimating genetic disorders caused by gene mutations, the BEIR III Report (NAS-BEIR, 1980) considered two exposure situations: continuous exposure over the lifetime of an individual and single exposure of the parents only.

It is estimated that since genetic effects produced are proportional to radiation dose, only about one percent to six percent of all spontaneous mutations in humans are due to natural background radiation. The incidence of numan genetic disorders in the absence of any additional radiation exposure is about 107,000 cases per million (about 10.7 percent) liveporn offspring (NAS-BEIR, 1980). A small increase in radiation exposure above background will lead only to a correspondingly small relative increase in the mutation rate. The doubling dose equivalent, i.e., the radiation dose equivalent required to produce as many more mutations as are already occurring spontaneously, is estimated to be 50 to 250 rem. The lower the doubling dose equivalent, the greater is the genetic risk from a given exposure. It is estimated that at equilibrium the increase in human genetic disorders in the population could be about 60 tp 1.100 per million liveborn offspring per rem of parental exposure received each generation before conception.

An average parental exposure of one rem before conception is expected to produce 5 to 65 additional genetic disorders per million live-oorn offspring in the first generation. These estimates are taken from the BEIR III Report (NAS-BEIR, 1980). Since the risk is conservatively taken to be linear at the dose levels of concern, the risk can be scaled readily for any dose equivalent.

DEVELOPMENTAL ABNORMALITIES IN THE NEWBORN

Among the somatic effects of low-level ionizing radiation other than cancer-induction, developmental abnormalities in the newborn are of greatest concern (NAS-BEIR, 1980). Exposure to high doses (hundreds of rads) of the embryo or fetus during gestation causes deatn, developmental abnormalities, retardation of growth, and functional impairment. There appears to be a threshold level for the induction of gross malformations in the newborn depending on the developmental stage

during gestation at which the exposure occurred and for the particular abnormality considered. Threshold doses for some abnormalities have been demonstrated, but such thresholds vary for different abnormalities. Recent data from the Japanese atomic bomb studies demonstrate measurable teratogenic damage (e.g., small head size associated with mental retardation) at doses of 10-19 rad (kerma) (NAS-BEIR, 1980). Decreasing the dose rate generally leads to a decrease in the developmental effects.

EFFECTS ON THE EYE

Cataracts. The most important response of the human eye to chronic radiation exposures is the production of cataracts resulting in impairment or loss of vision. The committee believes that cataract induction in space workers may be a problem, depending in part on the fluence and the RBE of the HZE particles. Impairment of vision in man can occur after single and fractionated exposure to low- and high-LET radiations (Britten et al., 1966). The response, including the length of the latent period and the severity of the effect, is dose-dependent above a threshold level. The dose-response curve for low-LET radiation-induced cataracts is highly sigmoid (Merriam and Focht, 1962). Above the threshold, the incidence of lens changes increases non-linearly with dose, the time to appearance is shortened. and the number of opacities that progress to cataracts with visual impairment increases. Opacities occur with doses from 200 to 500 rad. but do not progress to cataracts. Opacities may occur with 400 rac when the dose is fractionated over a period of three weeks and with 600 rad if the fractionation extends beyond three months. For low-LET radiations, protraction of the exposure is thought to reduce the effect on the lens less than for some other tissues.

Workers exposed to fast neutrons from cyclotrons have developed cataracts (Abelson and Kruger, 1949); however, precise lose levels are not known, and a reliable quality factor cannot be derived. In radio—therapy patients exposed to 14-MeV cyclotron neutrons, no change in the lens was seen following 80 rad in 12 fractions. However, slight permanent vision loss was found after 220 rad (Roth et al., 1976). Despite a knowledge of the neutron dose in the Japanese atomic bomb survivors, the neutron dose effect on the induction of cataracts is not known. Therefore, the RBE for cataract induction in man remains unknown, although the RBE will be dependent on the dose.

Although rodents are considered much more susceptible than humans to radiation-induced cataracts, rodent studies provide data on LET-RBE relationships that may obtain for human risk estimation. The mouse lens is sensitive to high-LET neutrons; the RBE at doses less than 10 rad could approach 100 or greater when "fleck" opacities are scored. The human lens would also be expected to be sensitive to densely ionizing radiations, such as neutrons or HZE particles (Merriam et al.,

1965; Batemann and Snead, 1969). Current research with rabbits exposed to single doses of various heavy charged particles or photons should provide additional information for risk estimation (Lett et al., 1980).

FERTILITY

Female.

Sensitivity to radiation-induced sterility varies with age (NAS-BEIR, 1980); women under 40 years require larger doses to induce menopause than women over 40 years. Doses below 100 rad are likely to have no effect on fertility or may produce transient sterility for a few months. Doses on the order of 170 rad can result in temporary sterility for from one to several years. However, a small percentage of women may be permanently sterilized by doses as low as 125 rad. Doses of 200 to 650 rad are required to sterilize five percent of women for more than five years. Doses of 625 to 2000 rad or more are required to sterilize 50 percent of women (Lushbaugh and Casarett, 1976). Protraction or fractionation of dose reduces the injury to the ovary (UNSCEAR, 1962). Fractionated doses greater than 2000 rad does not always produce sterility (Baker, 1971).

Male.

The seminiferous epithelium is among the most radiosensitive tissues in the adult. An acute dose of 15 rad will cause a significant decrease of the sperm count in about 40 percent of normal men within approximately two months (Langham, 967; Paulson, 1973). A dose of 30 to 50 rad results in aspermia and temporary sterility or infertility from a low sperm count. At this dose level, all sperm counts will return to normal after 9 to 19 months (Paulson, 1973). Doses up to 400 rad cause temporary sterility and/or infertility lasting up to 30 months (Paulson, 1973; Rowley et al., 1974). With testis exposures up to 500 rad, sterility followed by infertility may last for a period of five years or more (Rowley et al., 1974); but recovery may occur without serious physiological alterations.

Human data suggest that long periods of exposure to low dose rates can cause infertility. Men receiving radiotherapy for Hodgkin's disease and unavoidably receiving a daily dose of 10 to 15 rad to the testes (total of 140 to 300 rad) became sterile with no evidence of recovery for up to 40 months (Speiser et al., 1973). In occupationally exposed Roumanian workers, there was about 35 to 75 percent incidence of infertility (Popescu and Lancranjan, 1975) for periods of occupational exposure ranging from 2 to 22 years at 0.5 to 9.3 rad pc year $(0.01\ to\ 0.2\ rad\ per\ week)$. Thus, infertility could result with 90-day mission dose equivalents of 40 rem repeated over several years duration.

Laboratory animal experiments (Lushbaugh and Casarett, 1976; Langham, 1967), and limited human data on effects of protracted or fractionated low-dose exposures on spermatogenesis, indicate that fractionation of dose at low levels may be more effective on spermatogenesis than acute exposure because of the cyclic process of spermatogenesis (Lushbaugh and Casarett, 1976; Kramer et al., 1974).

OTHER HEALTH EFFECTS OF RADIATION

As stated above, carcinogenesis, mutagenesis, and cataractogenesis are the health effects of principal concern. At low doses of ridiation, cancer induction or promotion probably accounts for all excess mortality in experimental animals and man (NAS-BEIR, 1980). However, low doses of radiation may accelerate the appearance time of some non-lethal conditions (e.g., vascular or renal diseases) and thereby contribute to late life morbidity. The impact of such morbidity from non-neoplastic diseases is thought to be small, in comparison with neoplasia, based on life-span studies on mice (Storer et al., 1979; Sacher, 1976).

Vascular disease occurs as a consequence of aging in man and experimental animals. The vasculature of the mouse heart appears quite sensitive to high-LET neutron radiation, even at low doses (20 rad), and the expression of vascular disease occurs earlier than in aged controls (Yang et al., 1978; Stearner et al., 1979). Also, the degree of vascular injury is increased, in comparison with the same single dose, by fractionation of fission neutron doses over several months. Vascular sensitivity of man to HZE particles is unknown, but neither vascular injury nor other adverse health effects are considered to be of significance for space workers.

HZE Particle Radiation Effects

The potential health hazard of HZE particles (stripped atomic nuclei with Z > 2) from galactic space was suggested soon after their discovery. Tobias (1952) predicted that a single HZE particle could cause light sensation in the retina and suggested that a single HZE particle might kill or modify a column of living cells in tissues. Neon and argon particles at energies of about 0.5 GeV/nucleon have been snown by several authors (Worgle, 1980; Malacnowski, 1978; A. Nelson, 1980) to produce lesions in various eye structures of the rodent which appear to be track related phenomena as predicted by Tooias. Some workers have unsuccessfully tried to relate neuronal killing in the CNS to juch track related events.

The significance of this track and track structure related phenomenology to human hazard remains to be evaluated. It appears

possible, at least, that HZE particles can indeed act in ways that are qualitatively different from the more classical interaction models associated with other radiations. For purposes of this report, dose from HZE particles is ignored, considering all the other uncertainties in dose estimation.

APPENDIX C

GLOSSARY

Aspermia: The absence of sperm or a marked decrease in sperm

count.

Bremsstrahlung: Secondary photon radiation produced by deceleration

of charged particles.

Cataractogenesis: The induction of cataracts in the lens of tre eye.

Dose: A general form denoting the quantity of radiation

or energy absorbed. For special purposes, it must be appropriately qualified. If unqualified, it

refers to absorbed dose.

Absorbed dose: The energy imparted to matter by ionizing radiation per unit mass of irradiated material at the place of interest. The unit of absorbed dose is the rad.

Accumulated dose: Total dose resulting from repeated exposure to radiation.

Dose equivalent (DE): Quantity that expresses all radiations on a common scale for calculating the effective absorbed dose. It is defined as the product of the absorbed dose in rad and certain modifying factors (See Quality Factor). The unit of DE is the rem. For example, the radiation risk from one rem of any ionizing radiation is assumed to be equivalent to that from one rad of low-LET

radiation.

Dose rate: Absorbed dose delivered per unit of time.

Electron: A subatomic particle with a negative electrical

charge.

Electron flux

density: Electrons per square centimeter per second.

Electron volt: A unit of energy equivalent to the energy gained by an electron in passing through a potential differ-

ence of one volt. Larger multiple units of the electron volt are frequently used: KeV for thousand or kilo electron volts; MeV for million or mega electron volts (Abbr.: eV, 1 eV = 1.6x10-12erg.).

Exposure:

A measure of the ionization produced in air by X or gamma radiation. It is the sum of the electrical charges on all ions of one sign produced in air when all electrons liberated by photons in a volume element of air are completely stopped in air, divided by the mass of the air in the volume element.

The special unit of exposure is the roentgen.

Fluence

Particles per square centimeter.

Galactic cosmic

ravs (GCR):

High-energy nuclei with origins outside our solar system, probably originating in supernova explosions of stars or in the remnants of these explosions.

Gensynchronous

orbit (GEO):

(or Geostationary orbit). An orbit at an altitude

of approximately 36,000 kilometers.

Homeostasis:

The steady state for equilibrium in the living body with respect to various functions and to the chemical compositions of the fluids and tissues: the processes through which such body equilibrium is maintained.

High-energy, heavy

ions (HZE)

High energy heavy particles having a nuclear charge

(Z) in excess of 2.

Incidence:

The rate of occurrence of a disease within a specified period of time: usually expressed in a number of cases per million (106) per year.

Ion:

Atomic particle, atom, or chemical radical bearing an electrical charge, either negative or positive.

[onization:

The process by which a neutral atom or molecule

acquires a positive or negative charge.

Primary ionization: In collision theory, the ionization produced by the primary particles as contrasted to the "total ionization" which includes the "secondary ionization" produced by delta rays. Secondary ionization: Ionization produced by delta

rays.

Ionization

Number of ion pairs per unit volume. density:

Ionization

path (track): The trail of ion pairs produced by an ionizing

radiation in its passage through matter.

Ionizing

radiation: Any electromagnetic or particulate radiation capable

of producing ions, directly or indirectly, in its

passage through matter.

Secondary radiation: Radition resulting from absorption or other radiation in matter. It may be

either electromagnetic or particulate.

Kerma:

The quotient of dE_{tr} by dm, where dE_{tr} is the sum of the initial kinetic energies of all the charged particles liberated by indirectly ionizing particles in a volume element of the specified material, and dm is the mass of the matter in that volume element.

 $K = \frac{d\vec{E}_{tr}}{dm}$

Kerma has the same dimension as absorbed dose and both quantities have the same special unit, the rad. The unit in the International System of Units (SI) is the gray, symbol Gy; 1 Gy = 1 J κq^{-1} = 100 rad.

Latent period:

The period or state of apparent non-effectiveness between the time of exposure of tissue to an injurious agent and response to that exposure.

Linear energy

transfer (LET): The average amount of energy lost per unit of

particle track length.

Low-LET: Radiation energy transfer characteristic of electrons, x-rays, and Gamma rays.

High-LET: Radiation energy transfer characteristic

of heavy charged particles or fast neutrons.

Linear no thresh-

nold nypothesis: The assumption that a dose-effect curve derived from data in the nigh dose and dose-rate ranges may be

extrapolated linearly through the low dose and low dose range to zero, implying that, theoretically, any amount of radiation may cause some damage.

Low earth orbit

(LEO): Orbit at an altitude of approximately 500 km.

Alterations in the genes that result in a mutation. Mutagenesis:

Neoplasm: Any new abnormal growth. The term "neoplastic

disease" refers to the occurrence of either

malignant or benign tumors.

In the quantum theory of light, light consists of Photon:

tiny bundles of energy. Each bundle is a photon.

Guality

Rad:

factor (0): The linear-energy-transfer-dependent factor by which absorbed doses are multiplied to obtain (for radia-

tion protection purposes) a quantity that expresses-on a common scale for all ionizing

radiations-the effectiveness of the absorped dose.

The standard unit of absorbed dose, equal to energy absorption of 100 ergs per gram or 0.01 Joules per kilogram; supercedes the roentgen as the unit of dosage. One mrad is 0.001 rad. ICRU currently recommends use of the gray (Gy) as the measure of

absorbed dose; I Gy is equal to 100 rad.

Relative susceptability of cells, tissues, organs, Radiosensitivity: organisms of any living substance to the injurious action of radiation. Radiosensitivity and its

antonym radioresistance are currently used in a comparative rather than an apsolute sense.

Relative Biological Effectiveness

(RBE):

The RBE is a factor used to compare the prological effectiveness of absorbed radiation doses (i.e., rads) due to different types of ionizing radiation; more specifically, it is the experimentally determined ratio of an ansorped dose of a radiation

in question to the apsorped dose of a reference radiation required to produce an identical biological effect in a particular experimental organism or tissues. The RBE is the ratio of rem to rad. (If I rad of fast neutrons equalled in

letnality 3.2 rads of KVP x-rays, the RBE of the

fast neutrons would be 3.2).

A special unit of dose equivalent. The dose Rem: equivalent in rem is numerically equal to the

absorbed dose in rad multiplied by the quality factor, Q, and any other necessary modifying factors. The rem represents that quantity of radiation that is equialent—in biological damage of a specified sort—to 1 rad of 250 KVP x-rays.

ICRU currently recommends use of the sievert (Sv) as the measure of dose equivalent; 1 Sv is equal to 100 rem.

Solar cell:

A photovoltaic cell that converts light energy into

electrical energy.

Solar radiation:

The total electromagnetic radiation emitted by the

sun.

Solar particle

event (SPE):

An event caused by large magnetic upneavals on the solar surface which accelerate protons—and to a lesser extent, heavier nuclei—to nigh energies and

projects these particles into space.

South Atlantic Anomaly:

A region of the trapped particle zones which dips close to Earth in the southern Atlantic Ocean southeast of the Brazilian coast. This area is the most important source of radiation for space

workers in LEO.

SPS:

Satellite power systems.

Transfer

ellipse (TE):

The path followed by a space venicle in moving from one point in space to another (for example, from

LEO to GEU).

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