

UC San Diego

UC San Diego Previously Published Works

Title

Modelling gravity currents without an energy closure

Permalink

<https://escholarship.org/uc/item/26b8n9hq>

Authors

Konopliv, NA
Smith, Stefan G Llewellyn
McElwaine, JN
et al.

Publication Date

2016-02-25

DOI

10.1017/jfm.2015.755

Peer reviewed

Modeling gravity currents without an energy closure

N.A. Konopliv¹, Stefan G. Llewellyn Smith², J. N. McElwaine³
and E. Meiburg^{1†}

¹Department of Mechanical Engineering, University of California at Santa Barbara, Santa Barbara, CA 93106, USA

²Department of Mechanical and Aerospace Engineering, Jacobs School of Engineering, UCSD, 9500 Gilman Drive, La Jolla, CA 92093-0411, USA

³Department of Earth Sciences, Durham University, Science Labs, Durham, DH1 3LE, UK

(Received ?; revised ?; accepted ?. - To be entered by editorial office)

We extend the vorticity-based modeling approach of Borden & Meiburg (2013) to non-Boussinesq gravity currents and derive an analytical expression for the Froude number without the need for an energy closure or any assumptions about the pressure. The Froude number expression we obtain reduces to the correct form in the Boussinesq limit and agrees closely with simulation data. Via detailed comparisons with simulation results, we furthermore assess the validity of three key assumptions underlying both our as well as earlier models, *viz.* i) steady-state flow in the moving reference frame; ii) inviscid flow; and iii) horizontal flow sufficiently far in front of and behind the current. The current approach does not require an assumption of zero velocity in the current.

1. Introduction

Three quarters of a century ago, von Kármán (1940) introduced the idealized gravity current model shown in figure 1a. He considered the flow in the reference frame moving with the current front, and invoked three main simplifying assumptions: i) the flow is steady in this reference frame; ii) the flow is inviscid; and iii) the fluid inside the current is at rest. By neglecting the flow in the ambient and applying Bernoulli's law along the streamlines C-O and O-A, *i.e.*, by assuming that the mechanical energy is conserved along these streamlines, he obtained for the Froude number

$$F_h = \frac{U}{\sqrt{g'h}} = \sqrt{\frac{2}{\sigma}}. \quad (1.1)$$

Here, U denotes the front velocity of the gravity current, h represents its height, $g' = g(\rho_1 - \rho_2)/\rho_1$ indicates the reduced gravity, and $\sigma = \rho_2/\rho_1$ refers to the density ratio.

Benjamin (1968) objected to von Kármán's analysis on the grounds that Bernoulli's equation should not be assumed to hold along streamline O-A, due to the dissipation that occurs in this interfacial region as a result of the velocity shear between the current and the ambient, which causes the development of Kelvin-Helmholtz billows and turbulence. Benjamin instead considered a corresponding gravity current in a channel of finite depth H , as shown in figure 1b. By applying the same three simplifying assumptions as von Kármán, and also considering the pressure distributions far up- and downstream of the current front to be hydrostatic, Benjamin was able to write the conservation laws for

† Email address for correspondence: meiburg@engineering.ucsb.edu

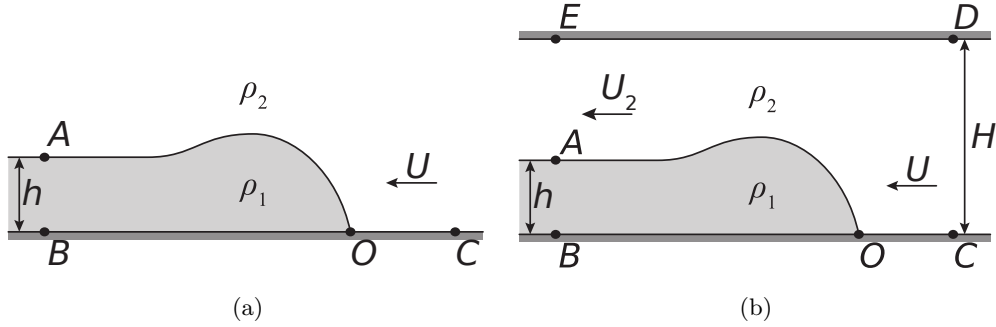


Figure 1: Idealized gravity current in a deep ambient (a) and a channel (b).

mass and horizontal momentum flux as

$$UH = U_2(H - h) \quad (1.2)$$

$$p_C H + \rho_2 U^2 H = p_B H + \frac{1}{2} g (\rho_1 - \rho_2) h^2 - g (\rho_1 - \rho_2) H h + \rho_2 U_2^2 (H - h) . \quad (1.3)$$

For a given set of values for current thickness, channel height and density ratio, the above relationships represent two equations for the three unknowns U , U_2 and $p_B - p_C$, so that one additional equation is required. To close the problem, Benjamin followed von Kármán's approach and applied Bernoulli's law; however, he did so along the bottom wall $C-B$ of the channel, rather than along the interface as von Karman had done. For a current of fractional height $\alpha = h/H$, Benjamin thus obtained for the Froude number

$$F_{H,b} = \frac{U}{\sqrt{g'H}} = \left[\frac{\alpha(1-\alpha)(2-\alpha)}{\sigma(1+\alpha)} \right]^{1/2} . \quad (1.4)$$

Note that the Froude number F_h based on the current height is related to the Froude number F_H based on the channel height by $F_h = F_H \alpha^{-1/2}$.

For Boussinesq gravity currents, Borden & Meiburg (2013) showed that invoking an energy closure assumption such as Bernoulli's equation in Benjamin's model becomes unnecessary if the conservation of vertical momentum is enforced, along with the conservation of mass and horizontal momentum. This approach bypasses the controversy between Benjamin and von Kármán entirely, as the conservation of energy or head loss arguments are not required. While there is no flow of vertical momentum into or out of the control volume $BCDE$, the importance of vertical momentum conservation inside the control volume is clear. The ambient fluid is first accelerated and then decelerated in the vertical direction, which affects the pressure profiles along the top and bottom walls. In turn, these profiles determine the pressure jump $p_B - p_C$ across the current front, for which the need of an additional equation originally arose. Borden & Meiburg (2013) showed that the conservation of vertical momentum can be accounted for by considering the linear combination of the differential versions of the steady-state, inviscid, horizontal and vertical momentum equations, in the form of the Boussinesq vorticity equation

$$\mathbf{u} \cdot \nabla \omega = -g' \frac{\partial \rho}{\partial x} , \quad (1.5)$$

where $\omega = \frac{\partial v}{\partial x} - \frac{\partial u}{\partial y}$ denotes the vorticity, and x and y represent the horizontal and vertical directions, respectively. By integrating Eq. (1.5) over the control volume, we

obtain a relation governing the total circulation around the control volume

$$\oint \omega \mathbf{u} \cdot \mathbf{n} \, dS = \iint -g' \frac{\partial \rho}{\partial x} \, dA . \quad (1.6)$$

Eq. (1.6) states that for incompressible flows in the Boussinesq limit the flow of vorticity into and out of the control volume is balanced by the baroclinic generation of vorticity inside the control volume. For a sharp interface, the area integral of the baroclinic term becomes $g'h$. Furthermore, no vorticity enters the control volume, and the flow of vorticity out of the control volume is confined to the vortex sheet between the current and the ambient. The vorticity flux carried by this sheet equals the vortex sheet strength, $\gamma = U_2$, multiplied by the sheet's principal velocity, $u_{PV} = U_2/2$ (Pozrikidis (1997); Saffman (1992)). Eq. (1.6) thus reduces to

$$\frac{1}{2}U_2^2 = g'h . \quad (1.7)$$

Combining the vorticity conservation relationship (1.7) with the continuity equation (1.2) produces

$$F_{H,c} = \sqrt{2\alpha}(1 - \alpha) , \quad (1.8)$$

where the subscript 'c' refers to 'circulation model.' Borden & Meiburg (2013) showed that with regard to the vorticity flux of Boussinesq currents this relationship between the Froude number and the current height results in better agreement with DNS simulation results than Benjamin's relationship (1.4). However, even Benjamin's model prediction is found to be quite close to the DNS data, which indicates that his zero-headloss assumption closely approximates the situation in the simulated flow field. We note that in the above analysis, the pressure jump $p_B - p_C$ across the current front has become decoupled from the problem of determining U and U_2 , which were determined from the conservation of mass and vorticity alone. Up to this point, we have used the conservation of horizontal momentum only in linear combination with the conservation of vertical momentum, *i.e.*, as the vorticity equation. Consequently, if desired, the pressure jump $p_B - p_C$ across the current front can now be determined from the horizontal momentum equation, as was shown by Borden & Meiburg (2013). The decoupling of the pressure in the above analysis is analogous to employing the streamfunction-vorticity formulation of the Navier-Stokes equations, which allows for the numerical simulation of incompressible flow fields without having to calculate the pressure explicitly. As explained earlier, by accounting for the conservation of mass, horizontal and vertical momentum, the above analysis did not have to invoke any assumptions about energy conservation. Rather, individual terms in the energy equation can now be evaluated, so that the overall loss of energy can be calculated *a posteriori*, rather than assumed *a priori*.

We remark that both Benjamin's model and the vorticity model assume that the flow is inviscid. However, the role of viscosity in a real flow affects the two models differently. Benjamin invokes the assumption of inviscid flow in order to apply Bernoulli's equation to a streamline along which he expects dissipation to be minimal. We expect that any small amount of viscosity, and hence dissipation, will cause a loss of mechanical energy, so that Bernoulli's equation will no longer hold exactly. The vorticity model, on the other hand, invokes the assumption of inviscid flow in the context of the vorticity equation, so that it can model the vorticity field as an infinitely thin sheet. A small amount of viscous diffusion in the flow will cause the sheet to attain a finite thickness. However, for the parallel flow field far behind the current front a small amount of viscosity will not affect the vorticity flux, which remains the same for a thin but finite vorticity layer as it is for a vortex sheet. Hence we would expect the vorticity model to be less sensitive to small

amounts of viscosity than Benjamin's model. The only caveat concerns the stagnation point O , where even a small amount of viscosity might potentially lead to a diffusive loss of vorticity out of the control volume.

As mentioned above, the investigation by Borden & Meiburg (2013) was limited to Boussinesq gravity currents. In the following, we extend their results to non-Boussinesq liquid gravity currents, such as the ones investigated experimentally by Lowe *et al.* (2005) and computationally by Birman *et al.* (2005). We will investigate in detail the significance of the three key assumptions invoked by all of the above authors, *viz.* steady-state flow, inviscid flow, and gravity current fluid at rest.

2. Non-Boussinesq gravity currents: Theory

In the following, we will present two alternate ways of extending the above analysis to liquid non-Boussinesq flows. The first approach, which more closely follows the work of Borden & Meiburg (2013) by focusing on the vorticity variable, will consider the problem under the standard assumptions of steady-state inviscid flow, with the gravity current fluid at rest. The second, alternative approach starts from the conservative form of the momentum equations for primitive variables. It will be shown that, with this approach, it is possible to relax some of the standard assumptions. The relationship between the two approaches will be discussed briefly towards the end of the section.

2.1. Vorticity approach

In order to extend the modeling approach by Borden & Meiburg (2013) to non-Boussinesq gravity currents, we begin with the steady-state Euler equation

$$\mathbf{u} \cdot \nabla \mathbf{u} = -\frac{1}{\rho} \nabla P + \mathbf{g} . \quad (2.1)$$

By taking the curl, we obtain the steady-state, inviscid, non-Boussinesq vorticity transport equation

$$\mathbf{u} \cdot \nabla \omega = -\nabla \times \left(\frac{1}{\rho} \nabla P \right) . \quad (2.2)$$

Integrating over the control volume and using the divergence theorem on the left hand side yields an expression analogous to the Boussinesq case (1.6)

$$\oint \omega \mathbf{u} \cdot \mathbf{n} \, dS = - \iint \nabla \times \left(\frac{1}{\rho} \nabla P \right) \cdot \mathbf{n} \, dA = - \int \frac{1}{\rho} \nabla P \cdot d\mathbf{l} . \quad (2.3)$$

The final integral is a contour integral along the boundary taken in the positive sense. Unlike the Boussinesq version of the problem, the pressure no longer decouples from the vorticity transport equation. However, since the density is taken to be piecewise constant, in each layer we may take the density out of the integral and reduce the right-hand side of (2.3) to $(\rho_2^{-1} - \rho_1^{-1})(P_O - P_A)$, which depends only on the difference in pressures between O and A . Taking the fluid in the current to be at rest leads to $P_O = P_B = P_A + \rho_1 g h$. We have not assumed anything about the pressure distribution in fluid 2 upstream or downstream.

From (2.3), we thus obtain for non-Boussinesq currents

$$\oint \omega \mathbf{u} \cdot \mathbf{n} \, dS = \frac{g' h}{\sigma} . \quad (2.4)$$

As for the Boussinesq case, there is no vorticity flux entering the control volume and the

vorticity leaving the control volume is confined to a vortex sheet with strength U_2 and principal velocity $U_2/2$. The vorticity balance can then be written as

$$\frac{1}{2}U_2^2 = \frac{g'h}{\sigma} . \quad (2.5)$$

Combining this with the continuity equation produces an expression for the Froude number

$$F_{H,c} = \sqrt{\frac{2\alpha}{\sigma}}(1 - \alpha) . \quad (2.6)$$

In the limit of small density contrasts $\sigma \approx 1$, so that the Boussinesq result is recovered.

2.2. Primitive variable approach

Alternatively, we can begin with the steady-state, two-dimensional Euler equation in conservative form

$$\nabla \cdot (\rho \mathbf{u}\mathbf{u}) + \nabla P = \rho \mathbf{g} , \quad (2.7)$$

where y is the vertical direction, so that $\mathbf{g} = (0, -g)$ and the velocity vector has components $\mathbf{u} = (u, v)$. We also assume that $\nabla \cdot \mathbf{u} = 0$. Taking the z -component of the curl of this equation gives a scalar equation that can be written as the divergence of a vector field

$$L = \nabla \cdot \mathbf{q} = \nabla \cdot \begin{pmatrix} g\rho + \partial_x(\rho uv) + \frac{1}{2}\partial_y[\rho(v^2 - u^2)] \\ -\partial_y(\rho uv) + \frac{1}{2}\partial_x[\rho(v^2 - u^2)] \end{pmatrix} = 0 . \quad (2.8)$$

After integrating over the control volume $BCDE$ and applying the divergence theorem, we are left with integrals over q_y along the top and bottom walls, and integrals over q_x along the in- and outflow boundaries. Along the top and bottom walls we have $v = 0$, so that

$$q_y = -\rho u \partial_y v - \rho u \partial_x u - \frac{1}{2}u^2 \partial_x \rho = -\frac{1}{2}u^2 \partial_x \rho, \quad (2.9)$$

where the last equality follows from $\nabla \cdot \mathbf{u} = 0$. Along the top there are no density gradients, so that the last term is zero. Along the bottom, if $x = 0$ denotes the front location, the velocity in the vicinity of the front will scale as $u \propto U\sqrt{x/h}$. This was shown first by von Kármán (1940) for the flow in the ambient, assuming that the current was stationary, and later extended by McElwaine (2005), who demonstrated that it also holds in the current. We expect the local density profile near the front to be approximately of error function shape $\rho = (\rho_1 + \rho_2)/2 - (\rho_1 - \rho_2)\text{erf}(x/W)/2$, where W is a (small) width. Multiplying this by the velocity and integrating gives a contribution proportional to $(\rho_1 - \rho_2)U^2W/h^2$ for the right hand side of equation (2.9), which is small provided that W is much less than the current height h .

Along the in- and outflow boundaries we have the q_x -term to consider. When integrating, we can use $v = 0$ along the top and bottom walls to obtain

$$\int_0^H q_x dy = g \int_0^H \rho dy + \int_0^H \partial_x(\rho uv) dy - \frac{1}{2}\rho u^2 \Big|_0^H . \quad (2.10)$$

This result is general in the sense that it holds for any density field, as well as any divergence free velocity field such that v vanishes on the upper surface.

We now limit ourselves to flows in which W/h is small, so that the $\int q_y dx$ contribution discussed above is negligible. Furthermore, we assume that $\partial_x(\rho uv) = 0$ sufficiently far in front of and behind the front. The implications of this assumption will be discussed in more detail below.

The shapes of the inflow and outflow velocity profiles are not important since, when

we integrate, only the top and bottom values contribute. The driving term is then seen as the difference in the integral of density between the inflow and outflow boundaries:

$$g \int_0^H [\rho_{BE}(y) - \rho_{CD}(y)] dy = \frac{1}{2} [u_E^2 \rho_E - u_B^2 \rho_B - u_D^2 \rho_D + u_C^2 \rho_C] . \quad (2.11)$$

In the case considered in detail in this paper, we have $u_E = U_2$, $u_B = 0$ and $u_D = u_C$, so that

$$gh(\rho_1 - \rho_2) = \frac{1}{2} U_2^2 \rho_2. \quad (2.12)$$

For general velocity profiles but piecewise constant density, equation (2.11) yields

$$\rho_2 \left[Hg - \frac{1}{2} u_D^2 + \frac{1}{2} u_C^2 \right] = \rho_2 \left[(H - h)g - \frac{1}{2} u_E^2 \right] + \rho_1 hg, \quad (2.13)$$

$$u_E^2 - u_D^2 + u_C^2 = 2hg(\rho_1/\rho_2 - 1) = 2hg'/\sigma. \quad (2.14)$$

When $u_D = u_C$, this relation gives a Froude number condition

$$F_{H,c} = \sqrt{\frac{2\alpha}{\sigma}} (1 - \alpha), \quad (2.15)$$

which is identical to the result obtained with the vorticity approach (2.6). However, in the case when $u_D \neq u_C$, there is no natural choice for the front velocity to define the Froude number.

The result can be extended to integration along a streamline rather than just $y = 0$ or $y = H$. Integrating from A to B to C and then back along a streamline just outside the current to A gives

$$u_A^2 = 2hg'. \quad (2.16)$$

This suggests that perhaps the best measure of velocity to use is actually the velocity u_A taken just outside the current.

The above analysis holds for general input and output velocity profiles. Details regarding the extension of the primitive variable approach to three-dimensional flows are presented in the appendix.

In the following, we analyze the implications of assuming

$$\int_0^H \frac{\partial}{\partial x} (\rho uv) dy = 0 \quad (2.17)$$

in the above derivation. Consider the inviscid, steady-state, vertical momentum equation in conservative form

$$\frac{\partial}{\partial x} (\rho uv) + \frac{\partial}{\partial y} (\rho v^2) = -\frac{\partial P}{\partial y} - \rho g, \quad (2.18)$$

and integrate from C to D , using $v_C = v_D = 0$

$$\int_C^D \frac{\partial}{\partial x} (\rho uv) dy + (\rho_D v_D^2 - \rho_C v_C^2) = (P_C - P_D) - \rho_2 g H. \quad (2.19)$$

This demonstrates that the assumption (2.17) corresponds to requiring that P_C and P_D are hydrostatic relative to one another, which also had been assumed as part of the vorticity approach in the previous section. Corresponding considerations apply to the outflow boundary, provided that $v_A = 0$, *i.e.*, that the interface at the outflow boundary is flat.

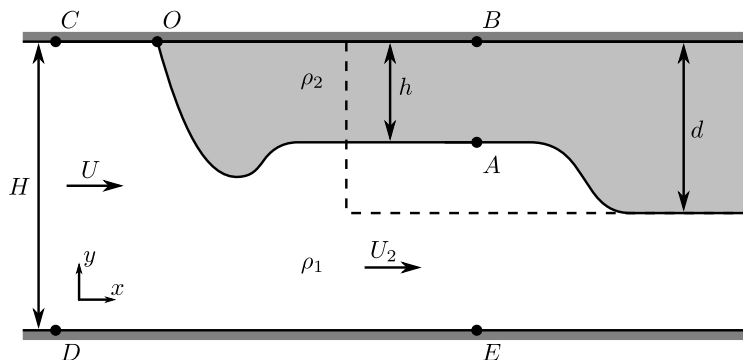


Figure 2: Schematic of a non-Boussinesq lock-exchange gravity current. The Navier-Stokes simulations focus on the buoyant current along the top wall, which more closely corresponds to a quasisteady flow in the moving reference frame than a negatively buoyant bottom current.

3. Numerical simulations

In order to assess the relative accuracy of Benjamin's and the vorticity model, we compare their predictions to two-dimensional Navier-Stokes simulations of lock-exchange gravity currents. The setup of the simulations is shown in figure 2, with the dashed line indicating the initial lock configuration. If the lock depth d is equal to (less than) the height H of the domain, the resulting flow is referred to as a full depth (partial depth) current.

During each simulation, one positively buoyant current is generated that propagates to the left along the top wall, and one negatively buoyant current propagating to the right. For full depth locks, this negatively buoyant current has the form of a gravity current along the bottom wall, whereas for partial depth locks, it is a bore traveling along the density interface. As will be seen below, the light current along the top wall generally can be approximated more accurately by a quasisteady flow in the reference frame moving with the current tip, so that it will be more suitable for assessing the validity of the various models. For light currents, Benjamin's analysis yields

$$F_{H,b} = \frac{U}{\sqrt{g'H}} = \left[\frac{\alpha(1-\alpha)(2-\alpha)}{1+\alpha} \right]^{1/2}, \quad (3.1)$$

instead of equation (1.4) for dense currents, while the vorticity model results in

$$F_{H,c} = \sqrt{2\alpha}(1-\alpha), \quad (3.2)$$

rather than the corresponding relationship (2.6) for dense currents.

3.1. Governing equations

We follow the simulation approach of Birman *et al.* (2005) and employ the incompressible, non-Boussinesq Navier-Stokes equations in two dimensions. As long as there is minimal diffusion, the velocity field can be considered divergence free, as the flow consists of two separate incompressible fluids. For a discussion of the effects of diffusion on the continuity equation and their quantitative assessment, we refer the reader to the appendix and to Chen & Meiburg (2002). The dynamic viscosities of the two fluids are taken to be equal, and the density field evolves based on a convection-diffusion equation. To minimize mixing, we employ small diffusivities. Referring to figure 2 and letting a star

symbol denote a dimensionless quantity, we nondimensionalize the equations with the lock height d , the buoyancy velocity $U_b = \sqrt{g'd}$, where $g' = g(\rho_1 - \rho_2)/\rho_1$ is the reduced gravity, the dynamic pressure $\rho_1 U_b^2$ and the ambient fluid density ρ_1 to obtain

$$\nabla \cdot \mathbf{u}^* = 0 \quad (3.3)$$

$$\frac{D\mathbf{u}^*}{Dt^*} = -\frac{1}{\rho^*} \nabla P^* + \frac{1}{\rho^* Re} \nabla^2 \mathbf{u}^* + \frac{1}{1-\sigma} \mathbf{e}_g \quad (3.4)$$

$$\frac{D\rho^*}{Dt^*} = \frac{1}{ReSc} \nabla^2 \rho^* \quad (3.5)$$

Here D/Dt^* denotes the material derivative and \mathbf{e}_g is the unit vector in the direction of gravity. The nondimensional parameters are then

$$Re = \frac{\rho_1 U_b d}{\mu}, \quad Sc = \frac{\mu}{\rho_1 \kappa}, \quad \sigma = \frac{\rho_2}{\rho_1}, \quad (3.6)$$

where μ represents the dynamic viscosity and κ indicates the molecular diffusivity of the density field. Alternatively, we can employ the Péclet number $Pe = Re Sc$. We recast the momentum equation (3.4) into the vorticity form

$$\frac{D\omega^*}{Dt^*} = \frac{\rho_y^*}{\rho^*} \frac{Du^*}{Dt^*} - \frac{\rho_x^*}{\rho^*} \frac{Dv^*}{Dt^*} + \frac{1}{\rho^* Re} \nabla^2 \omega^* - \frac{\rho_x^*}{(1-\sigma)\rho^*}, \quad (3.7)$$

where the velocity is defined as

$$\mathbf{u}^* = \begin{pmatrix} u^* \\ v^* \end{pmatrix}. \quad (3.8)$$

We employ free-slip and no-flux conditions along all walls, so that the vorticity vanishes along the boundaries. We emphasize that this does not necessarily translate into a symmetry boundary condition for the vorticity field. To clarify this issue, consider the flow along the top wall in the vicinity of the stagnation point. Applying the boundary conditions $\omega^* = 0$ and $\rho_y^* = 0$ yields

$$\omega_{yy}^* = \frac{Re}{1-\sigma} \rho_x^*, \quad (3.9)$$

so that $\omega_{yy}^* \neq 0$ in regions with horizontal density gradients.

3.2. Computational approach

The unsteady simulations are performed in a streamfunction-vorticity formulation, by integrating equations (3.7) and (3.5) with an explicit third order, low storage Runge-Kutta scheme (Williamson 1980). The time derivatives $\frac{\partial u^*}{\partial t^*}$ and $\frac{\partial v^*}{\partial t^*}$ appearing on the right hand side of (3.7) are evaluated iteratively at each Runge-Kutta substep. A pseudospectral method in the x -direction and a sixth order compact finite difference scheme in the y -direction are employed for the spatial discretization. As mentioned above, symmetry boundary conditions cannot be applied along the top and bottom walls, so that we instead employ right at the boundary a one-sided third order scheme for the concentration and a fourth-order scheme for the vorticity, along with a centered fourth order scheme one point away from the boundary.

An equation for pressure can be found by taking the divergence of (3.4)

$$\nabla^2 P^* = -2\rho^* \left[\left(\frac{\partial u^*}{\partial x^*} \right)^2 + \frac{\partial u^*}{\partial y^*} \frac{\partial v^*}{\partial x^*} \right] - \frac{\partial \rho^*}{\partial x^*} \frac{Du^*}{Dt^*} - \frac{\partial \rho^*}{\partial y^*} \left(\frac{Dv^*}{Dt^*} + \frac{1}{1-\sigma} \right). \quad (3.10)$$

Since this pressure relation is decoupled from the vorticity and density equations, the

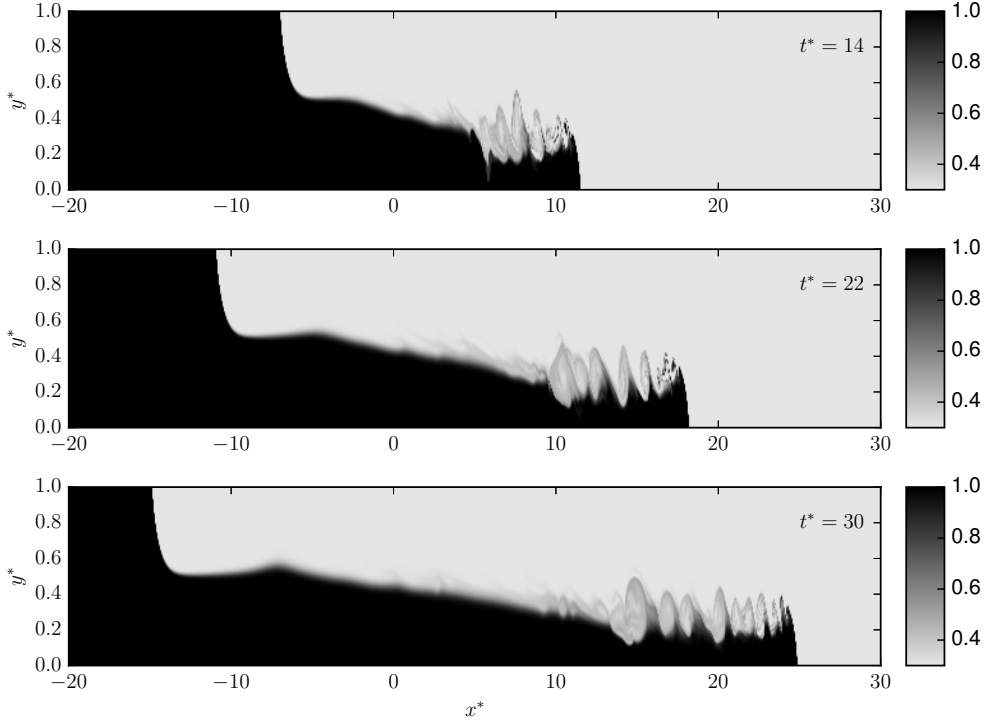


Figure 3: Simulation results for the density field of a full depth, non-Boussinesq flow with $Re = 5,000$, $Pe = 50,000$ and $\sigma = 0.3$.

pressure field can be evaluated during a postprocessing step after the simulation has finished.

3.3. Diagnostic tools

Figure 3 shows a representative full depth simulation at various times. The computational grid employs $16,384 \times 512$ points, with a time step of $O(5 \times 10^{-4})$, although its exact size varies according to the CFL condition. The figure confirms that the buoyant current propagating to the left along the top wall is more amenable to quasisteady modeling than the bottom current. Nevertheless, below we will discuss comparisons between DNS simulation results and model predictions for both the upper and the lower current.

The simulation is performed in the laboratory frame, and the results are then shifted to the reference frame moving with the current front during postprocessing. Towards this end, we employ linear interpolation to find the tip of the upper current as the location where $\rho^* = \frac{\sigma+1}{2}$ along the top wall. The front velocity U^* is then determined via linear regression on the front location *vs.* time data, (*cf.* Figure 4). To shift the results to the moving reference frame, U^* is subtracted from the laboratory frame velocity field.

The height $h^*(x^*, t^*)$ of the top current is defined as

$$h^*(x^*, t^*) = \frac{H}{d} - \int_0^{H/d} \frac{\rho^*(x^*, y^*, t^*) - \sigma}{1 - \sigma} dy^*. \quad (3.11)$$

For the flow of Figure 3, the current height is shown as a function of the distance be-

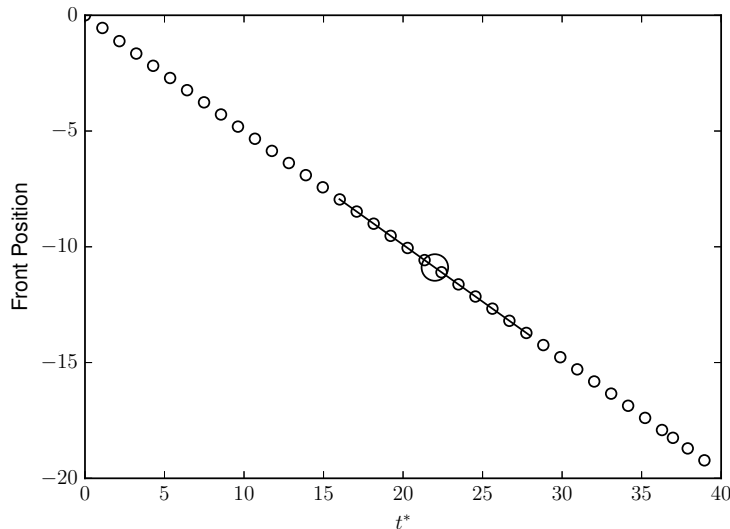


Figure 4: Calculation of the quasisteady front velocity U for the full depth top current with $Re = 5,000$, $Pe = 50,000$ and $\sigma = 0.3$. The small circles represent the tip location at every 2,000th time step. In order to evaluate the front velocity at a given time, e.g., the large circle, we employ a local linear best fit of the front locations, as indicated by the line.

hind the current tip in Figure 5, at selected times. This confirms that the steady-state approximation holds with good accuracy near the front of the buoyant current.

In order to assess the validity of Benjamin’s and the current model, we will primarily compare their predictions for the vorticity flux as a function of location with corresponding simulation results. Borden & Meiburg (2013) discuss the reasons for focusing on the vorticity flux, rather than the front velocity, due primarily to the difficulty in identifying a single representative value for the current height to use in (1.4) and (2.6). In the past, different authors have employed such measures as the first maximum in the current height behind the front, the current height at the gate location, a spatially averaged value for this purpose or the center of mass Anjum *et al.* (2013). Depending on which value is selected to represent the current height, the predicted front velocities can vary appreciably, so that the front velocity is ill-suited for determining which model is more accurate.

The vorticity flux Ω_B predicted by the Benjamin model can be found by using (3.1), along with the conservation of mass

$$\frac{\Omega_B}{g'd} = \Omega_B^* = \frac{h}{d} \frac{2 - \alpha}{2 - 2\alpha^2} \quad (3.12)$$

The corresponding vorticity flux predicted by the current model is

$$\frac{\Omega_C}{g'd} = \Omega_C^* = \frac{h}{d}, \quad (3.13)$$

cf. also equations (16) and (18) in Borden & Meiburg (2013). Both models predict identical fluxes for $\alpha = \frac{1}{2}$ and in the limit $\alpha \rightarrow 0$, *i.e.*, for currents that either occupy half the channel height or are much smaller than the channel height. The ratio be-

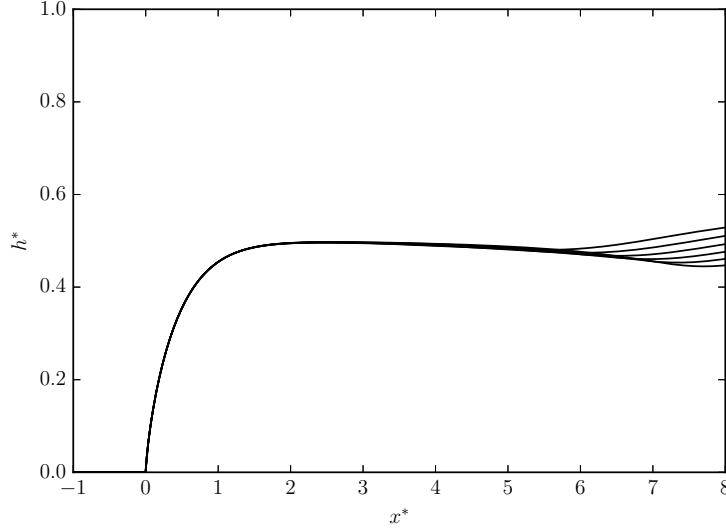


Figure 5: Current height as a function of distance behind the front for a full depth top current with $Re = 5,000$, $Pe = 50,000$ and $\sigma = 0.3$, at $t^* = 20, 22, 24, 26, 28$ and 30 . The steady-state approximation is seen to be valid in the vicinity of the current tip.

tween the two predicted vorticity fluxes reaches a maximum of approximately 1.07 at $\alpha = 2 - \sqrt{3} \approx 0.268$.

The origin of vorticity flux discrepancies between simulation results and theoretical predictions will be discussed here for the vorticity approach, with a corresponding discussion for the primitive variable approach given in the appendix. If we had kept the viscous and unsteady terms when deriving (2.3), we would have obtained

$$\Omega^* = \Omega_C^* + E_P^* - E_t^* - E_\mu^* , \quad (3.14)$$

where Ω^* represents the instantaneous dimensionless vorticity flux out of the domain. Ω_C^* indicates the dimensionless vorticity flux predicted by the vorticity model for the steady, inviscid case in which the gravity current fluid is at rest. E_P^* , E_t^* and E_μ^* denote the deviations from this idealized model due to, respectively, fluid motion within the gravity current, unsteadiness and viscous effects

$$E_P^* = \iint -\nabla \times \left(\frac{1}{\rho^*} \nabla P^* \right) dA^* - \Omega_C^* , \quad (3.15)$$

$$E_t^* = \iint \nabla \times \left(\frac{\partial \mathbf{u}^*}{\partial t^*} \right) dA^* , \quad (3.16)$$

$$E_\mu^* = -\frac{1}{Re} \iint \nabla \times \left(\frac{1}{\rho^*} \nabla^2 \mathbf{u}^* \right) dA^* , \quad (3.17)$$

where the integration is carried out over the control volume $BCDE$. The discrepancies derived in the appendix for the primitive variable approach are closely related to (3.15), (3.16) and (3.17). For this reason, we will in the following section limit our discussion of the discrepancies between theoretical model predictions and simulation results to terms (3.15), (3.16) and (3.17).

We furthermore remark that, if we assume a hydrostatic pressure profile along the

downstream boundary B-A-E of the control volume, E_P^* can alternatively be evaluated as

$$E_P^* = (P_O^* - P_B^*) \frac{\sigma - 1}{\sigma}. \quad (3.18)$$

The difference between evaluating E_P^* via (3.15) and (3.18) thus provides information on how close to hydrostatic the pressure profile is along B-A-E. The pressure difference $P_O^* - P_B^*$ can be found by integrating the x -momentum equation from O to B in the simulation. The x -momentum equation yields

$$P_O^* - P_B^* = \frac{\sigma U_B^{*2}}{2} - \sigma E_{\mu,B}^* - \sigma E_{t,B}^* \quad (3.19)$$

$$E_{\mu,B}^* = \frac{1}{\sigma Re} \int_O^B \nabla^2 u^* dx^* \quad (3.20)$$

$$E_{t,B}^* = - \int_O^B \frac{\partial u^*}{\partial t^*} dx^* \quad (3.21)$$

Note that $E_{\mu,B}^*$ and $E_{t,B}^*$ can be thought of as partial evaluations of E_μ^* and E_t^* after using Stokes' theorem. By substituting (3.19) into (3.18), one obtains

$$E_P^* = (1 - \sigma) \left(-\frac{U_B^{*2}}{2} + E_{\mu,B}^* + E_{t,B}^* \right) \quad (3.22)$$

For $E_{\mu,B}^* \approx E_\mu^*$ and $E_{t,B}^* \approx E_t^*$, E_P^* , E_μ^* and E_t^* will tend to cancel each other out partially in (3.14). This effect will be greatest when σ and the fluid motion inside the current U_B^* are both small. As $\sigma \rightarrow 1$, $E_P^* \rightarrow 0$, which is consistent with the Boussinesq vorticity model, which did not require any assumptions regarding the pressure profile inside the current.

4. Simulation results and discussion

4.1. Full Depth Lock Releases

Figure 6 compares the model predictions to the vorticity flux in the simulation for the full depth current with $Re = 5,000$, $Pe = 50,000$ and $\sigma = 0.3$, as a function of the distance behind the current tip. We note that both model predictions are very close to the simulation result, and also to each other. This is perhaps not unexpected in light of the fact that for a full depth current $\alpha \approx 0.5$, and that for $\alpha = 0.5$ the vorticity model (3.13) and Benjamin's model (3.12) predict identical vorticity flux values.

We now analyze the magnitude of the terms that account for the deviation between the simulation result and the prediction by the vorticity model, i.e., E_P^* , E_t^* and E_μ^* . Figure 7a shows the values of the integrals in (3.15), (3.16) and (3.17) as functions of the distance x^* of the control volume boundary B-A-E behind the current front. The figure indicates that close to the current tip the assumption of steady flow is very accurate. Further downstream, the influence of the unsteadiness increases, which is consistent with the graphs of the current heights at various times shown in figure 5. The influence of viscous diffusion is significant near the tip of the current, but very small further downstream. The fluid motion inside the gravity current plays a significant role near the current tip, and farther downstream where the current height varies more strongly with x^* . Figure 7b confirms that if the vorticity model prediction is augmented by the three terms E_P^* , E_t^* and E_μ^* , the correct simulation result for the vorticity flux is recovered.

Figure 8 shows the magnitude of the pressure term E_P^* as a function of the distance x^* of the downstream control volume boundary B-A-E behind the current tip. The open

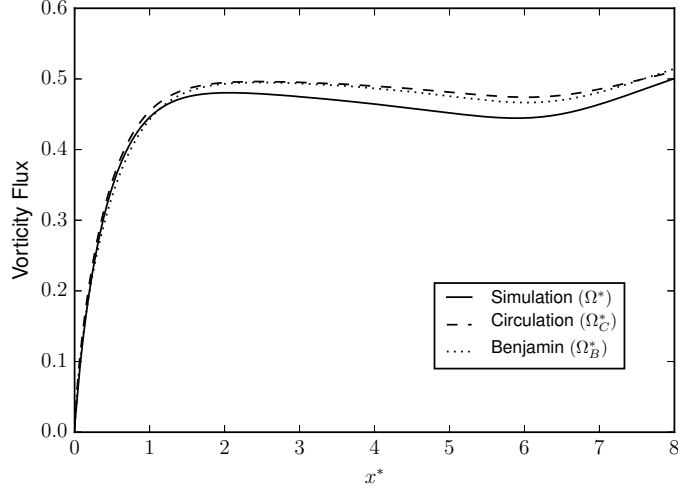


Figure 6: Vorticity flux normalized by $g'd$ vs. distance x^* behind the current head for the full depth current with $Re = 5,000$, $Pe = 50,000$ and $\sigma = 0.3$ at $t^* = 22$. The values predicted by Benjamin's model (3.12) and the circulation model (3.13) are close to each other, and to the simulation results.

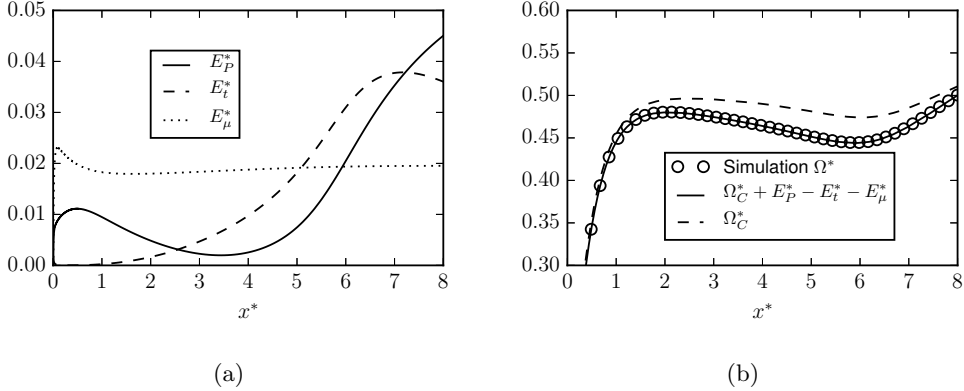


Figure 7: (a) Components of the difference between the vorticity flux predicted by the circulation model and the flux observed in the simulation, stemming from the three assumptions of motionless fluid inside the current (E_P^*), steady state (E_t^*) and inviscid flow (E_μ^*). (b) Simulation vorticity flux Ω^* along with Ω_C^* and $\Omega_C^* + E_P^* - E_t^* - E_\mu^*$ as functions of x^* , for the full depth current with $Re = 5,000$, $Pe = 50,000$ and $\sigma = 0.3$ at $t^* = 22$. All quantities are evaluated directly from the simulation data, and made dimensionless by $g'd$. The discrepancy between the vorticity flux Ω_C^* predicted by the vorticity model and the simulation result Ω^* is due to the quantities E_P^* , E_t^* and E_μ^* . Here E_P^* is evaluated using (3.15).

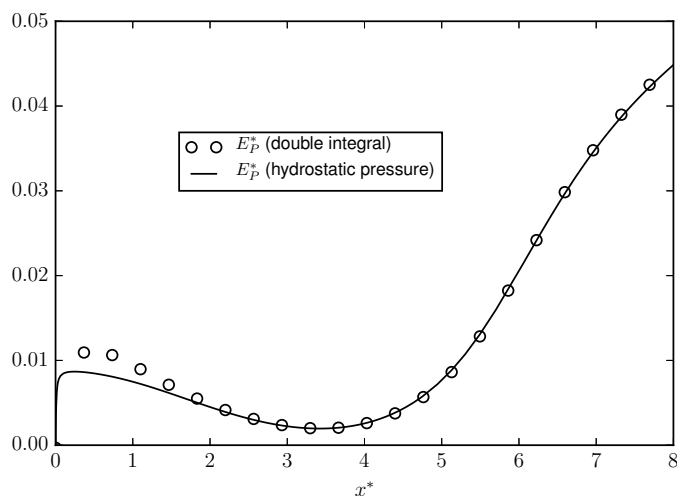


Figure 8: Pressure-related deviation E_P^* evaluated using (3.15) and (3.18), for the full depth current with $Re = 5,000$, $Pe = 50,000$ and $\sigma = 0.3$ at $t^* = 22$. The difference near the tip reflects the non-hydrostatic nature of the pressure in this region, since (3.18) implied a hydrostatic pressure distribution. Farther behind the current tip, the assumption of hydrostatic pressure is very accurate. All quantities are made dimensionless by $g'd$.

symbols are obtained by direct integration of the integral in (3.15) from the simulation pressure field, while the solid line assumes a hydrostatic pressure profile along B-A-E and evaluates (3.18). The results are shown to be in good agreement everywhere except near the current tip, which reflects the non-hydrostatic nature of the pressure field in this region. Recall that the non-Boussinesq vorticity model made two assumptions about the pressure: (a) it assumed that the pressure distribution at the downstream boundary of the control volume is hydrostatic; and (b) it assumed that as a result of the current fluid being at rest, $P_O - P_B = 0$. Figure 8 indicates that far behind the current front (a) is very accurate, so that (b) is largely responsible for the discrepancy between simulation results and model predictions for the vorticity flux.

Figure 9 analyzes the dependence of E_P^* , E_t^* and E_μ^* on the density ratio σ and on Re . We observe that increases in σ or Re tend to reduce the magnitude of all three of these terms, which indicates that predictions by the vorticity model become more accurate as the flow is less viscous and closer to Boussinesq. The decrease in E_P^* for larger σ is consistent with (3.22) and reflects the fact that the pressure profile inside the current becomes less influential as the flow approaches Boussinesq conditions. In order to understand the decrease in E_P^* for larger Re -values, it is important to realize that for higher Re the shear layer between the current and the ambient becomes thinner, so that the ambient stream drags less current fluid with it. Consequently, the counterflow along the top wall inside the current required to replenish the loss of current fluid in the mixing layer is reduced in strength for larger Re , which is confirmed by figure 10. Hence, the streamwise pressure gradient inside the current is weaker for higher Re , so that E_P^* is reduced.

The weaker flow inside the current for larger σ and higher Re also lowers any unsteady

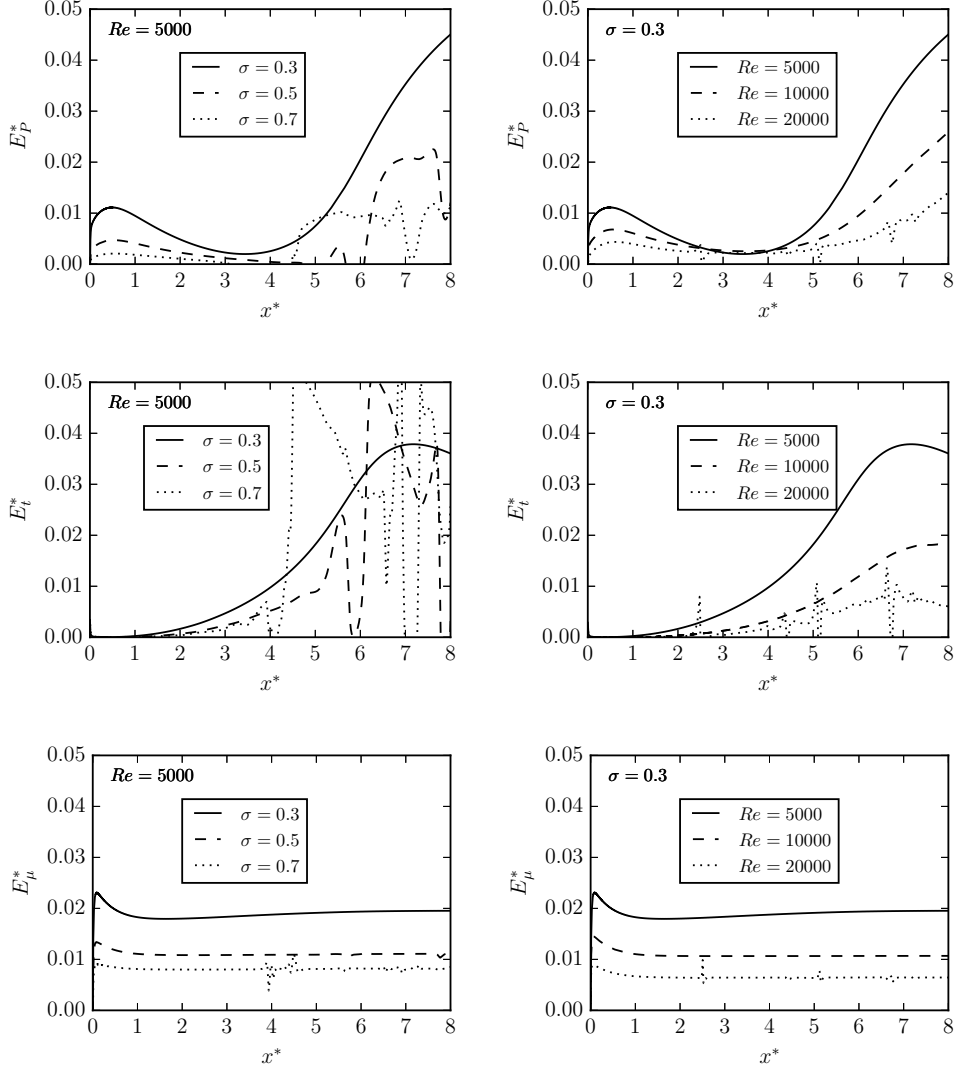


Figure 9: Effect of σ (left) and Re (right) on E_P^* , E_t^* and E_μ^* . All three get smaller for lower viscosity and reduced density contrast, as explained in the text.

effects, thereby reducing E_t^* . Finally, eqn. (3.20) indicates that E_μ^* scales with $\frac{1}{\sigma Re}$, so that it should decrease for larger values of σ and Re , which is confirmed by figure 9.

4.2. Partial Depth Lock Releases

Figure 11 shows the evolution of a partial depth gravity current from a lock with $d/H = \frac{1}{2}$. The front of the buoyant current is not as smooth as that of the corresponding full depth current discussed earlier, as a result of instabilities that emerge along the interface. Nevertheless, figure 12 indicates that for both values of σ tested, the vorticity model predicts the vorticity flux accurately near the front. Benjamin's model, while not quite as close to the DNS results as the vorticity model, nevertheless shows good quantitative agreement with the simulation data, which indicates that his zero-headloss assumption

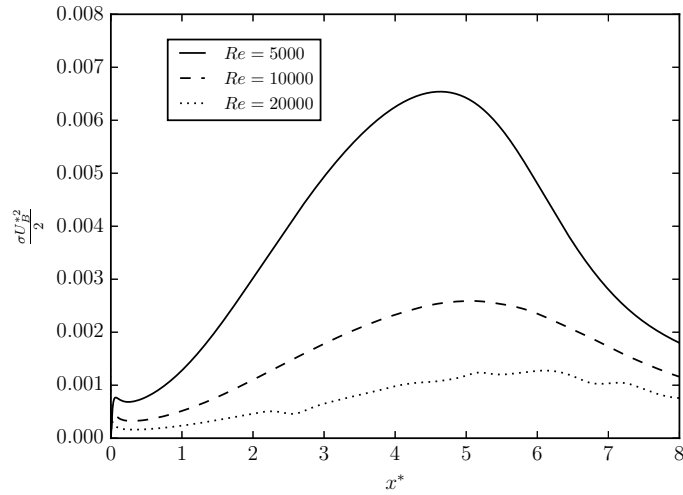


Figure 10: Streamwise velocity along the top wall inside the current, as a function of the distance x^* behind the current front, for $\sigma = 0.3$. For increasing Re -values, the flow inside the current is reduced.

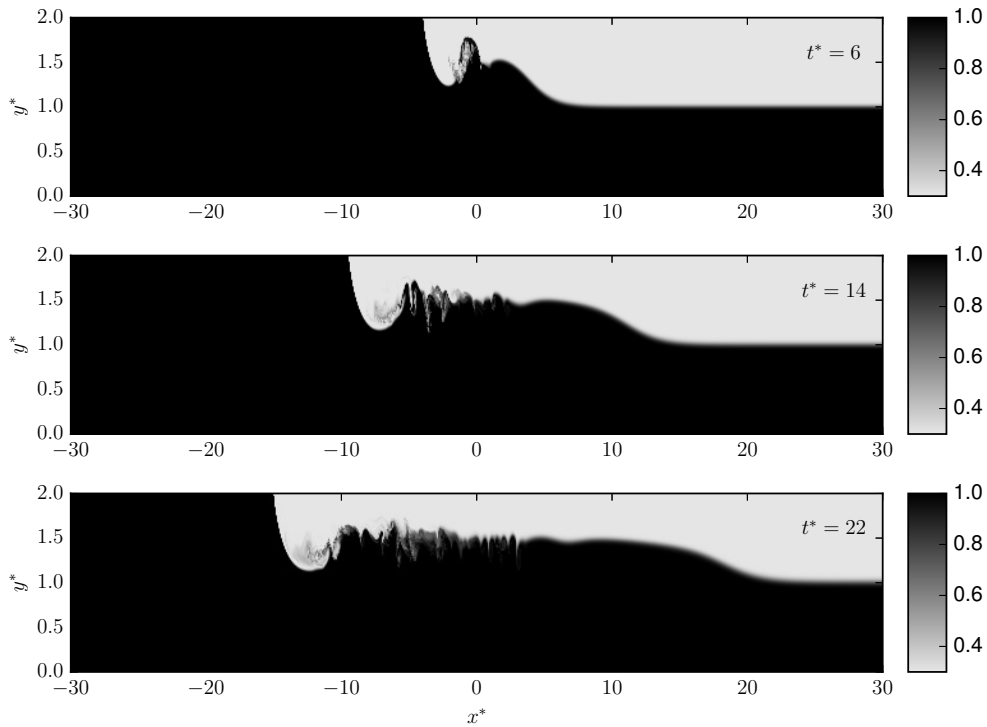


Figure 11: Density field of a partial depth, non-Boussinesq gravity current with $Re = 5,000$, $Pe = 50,000$, $\sigma = 0.3$ and $d/H = 0.5$.

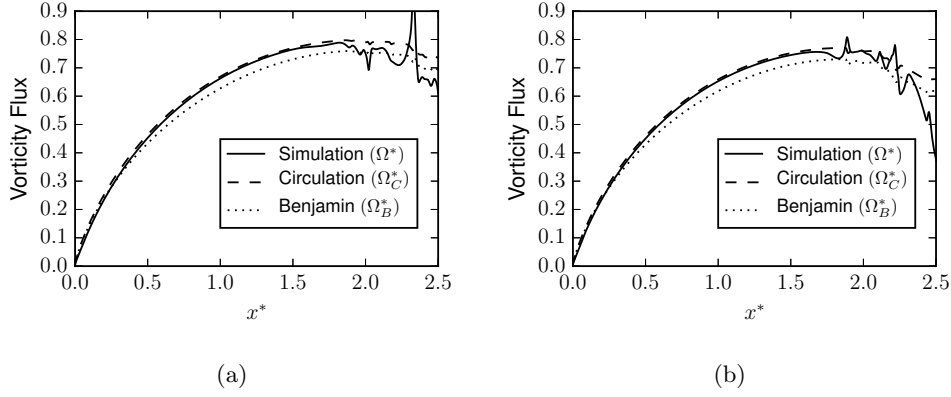


Figure 12: Vorticity flux *vs.* distance behind the current tip for the partial depth current with $Re = 5,000$, $Pe = 50,000$, $d/H = 0.5$ and $\sigma = 0.2$ (a) and $\sigma = 0.3$ (b) at $t^* = 6$. For both density ratios, the circulation model (3.13) is seen to agree very closely with the simulation results. Benjamin's model (3.12), while not quite as close, nevertheless also yields good quantitative agreement.

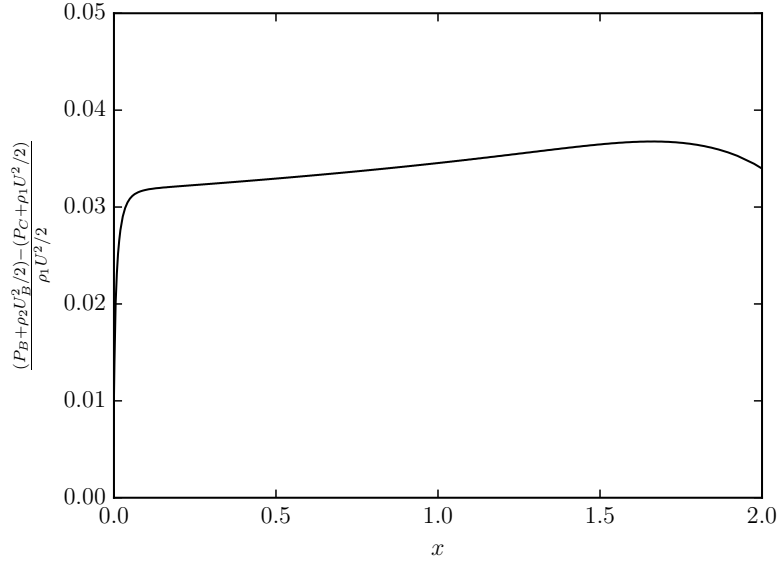


Figure 13: Headloss inside the current along the top wall, for the partial depth current with $Re = 5,000$, $Pe = 50,000$, $d/H = 0.5$ and $\sigma = 0.3$ at $t^* = 6$. The headloss is limited to about 3-4% of the free stream kinetic energy, which explains the good quantitative agreement between Benjamin's model predictions and the simulation results.

closely approximates the situation in the simulated flow. This is confirmed by figure 13, which demonstrates that (for the present case of slip walls) the headloss along the wall is limited to about 3-4% of the free stream kinetic energy.

The reasons for the good agreement between the vorticity model predictions and the simulation data become clear from figure 14, which shows the fluid velocity along the

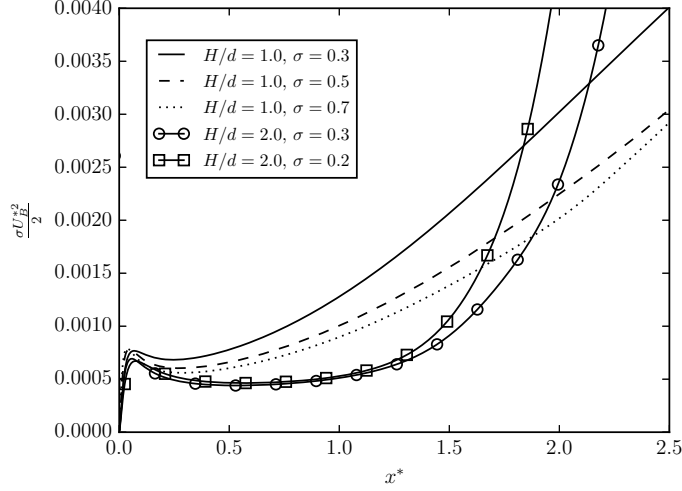


Figure 14: Fluid velocity inside the current along the top wall. Near the current tip, partial depth currents (shown at $t^* = 6$) exhibit smaller velocities than full depth currents (shown at $t^* = 22$).

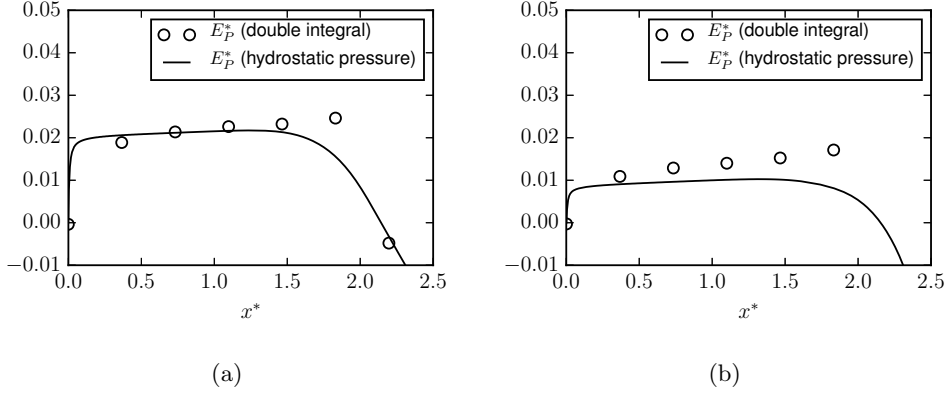


Figure 15: Deviation E_P^* due to the fluid motion inside the gravity current, evaluated from (3.15) (circles) and (3.18) (solid line), respectively. The simulations are the same as in figure 11 at $t^* = 6$ with $\sigma = 0.2$ (a) and $\sigma = 0.3$ (b), and all quantities are made dimensionless by $g'd$.

top wall inside the current, for several full and half depth currents. The partial depth currents generally give rise to smaller velocities inside the current. This is a consequence of the weaker acceleration of the ambient fluid around the tip of partial depth currents, so that partial depth currents experience less shear and a lower momentum transfer. Equation (3.22) indicates that the weaker values of U_B^* associated with half depth currents enhance the partial cancellation of E_P^* by E_μ^* and E_t^* , thereby resulting in improved model predictions.

Figure 15 shows the deviation E_P^* due to the fluid motion inside the gravity current, evaluated from (3.15) and (3.18), respectively. Both results agree closely with each other

in the vicinity of the current tip, which demonstrates that the pressure is approximately hydrostatic there, despite the interfacial instabilities. Consistent with our earlier observations for full depth currents, the deviation decreases for larger σ . Furthermore, the values of E_P^* and E_μ^* as a fraction of Ω_C^* are only about half as large as for the full depth current. This also contributes to the good agreement observed in figure 12.

4.3. Dense Currents

We now focus on the dense current moving towards the right along the bottom wall in figure 3. Figure 16 indicates that this current also has a steady front velocity. Figure 17 shows the bottom current heights for several times, corresponding to figure 5 for the top current. While a steady-state region again develops near the tip, it is much shorter than that for the top current, as a result of the turbulent billows. Figure 18 compares DNS results and model predictions for the vorticity flux in this region. The model predictions are given in (4.1) and (4.2).

$$\frac{\Omega_B}{g'd} = \Omega_B^* = \frac{h}{d} \frac{2 - \alpha}{2\sigma(1 - \alpha^2)} \quad (4.1)$$

The corresponding vorticity flux predicted by the current model is

$$\frac{\Omega_C}{g'd} = \Omega_C^* = \frac{h}{d} \frac{1}{\sigma} \quad (4.2)$$

These predictions differ by (3.12) and (3.13) by a factor of $1/\sigma$. Good agreement is seen for both models, in spite of the fact the hydrostatic pressure assumption may not be very accurate so close to the tip.

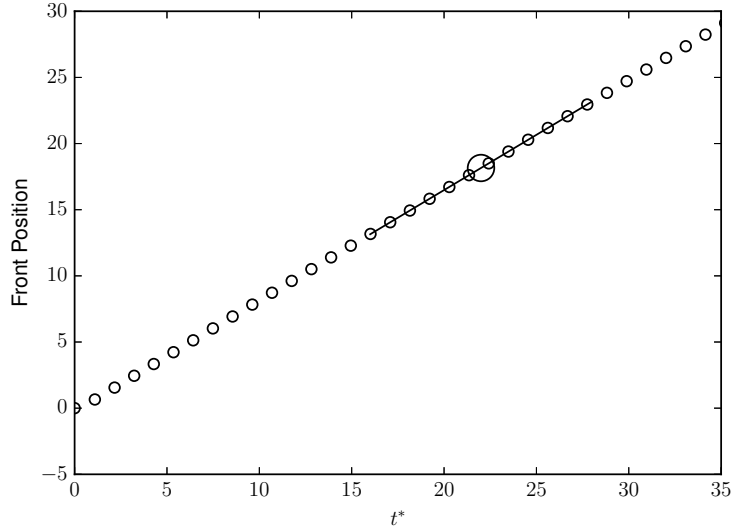


Figure 16: Calculation of the quasisteady front velocity U for the full depth bottom current with $Re = 5,000$, $Pe = 50,000$ and $\sigma = 0.3$. The small circles represent the tip location at every 2,000th time step. In order to evaluate the front velocity at a given time, e.g., the large circle, we employ a local linear best fit of the front locations, as indicated by the line.

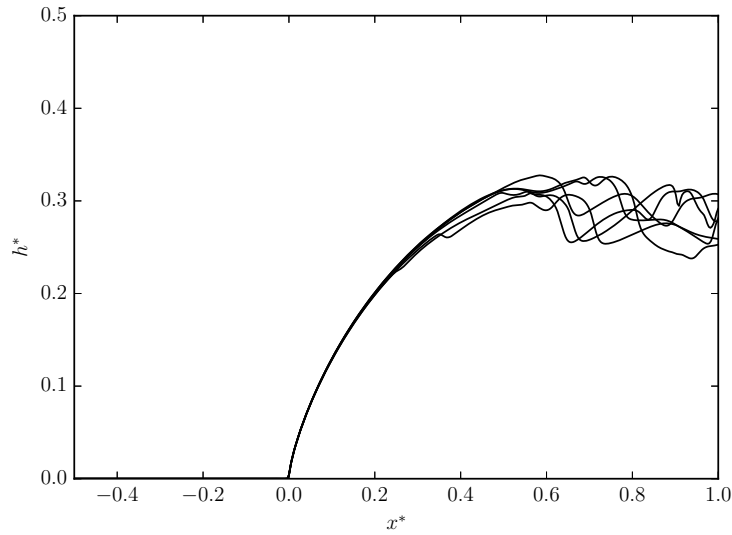


Figure 17: Current height as a function of distance behind the front for a full depth bottom current with $Re = 5,000$, $Pe = 50,000$ and $\sigma = 0.3$, at $t^* = 20, 22, 24, 26, 28$ and 30 . The steady-state approximation is seen to be valid in the vicinity of the current tip, although this region steady-state region is significantly shorter than that for the top current.

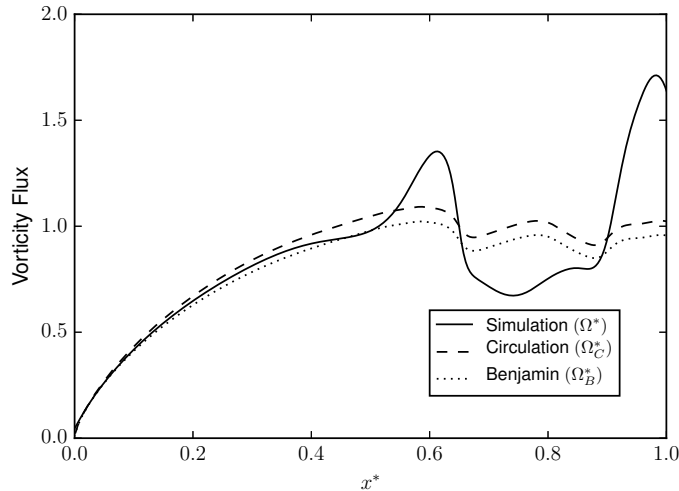


Figure 18: Vorticity flux normalized by $g'd$ vs. distance x^* behind the current head, for the full depth bottom current with $Re = 5,000$, $Pe = 50,000$ and $\sigma = 0.3$ at $t^* = 22$. In the steady-state region near the tip, the values predicted by Benjamin's model (4.1) and the circulation model (4.2) are close to each other, and to the simulation results.

5. Summary and conclusions

In the present investigation we have extended the vorticity-based modeling approach by Borden & Meiburg (2013) to non-Boussinesq gravity currents. This approach enables us to arrive at a closed form solution for the Froude number without having to invoke an energy-based closure assumption, such as had been required in the analyses by von Kármán (1940) and Benjamin (1968). Hence the vorticity approach bypasses the discussion among those authors as to which energy closure provides the optimal fit with experimental and simulation data.

In the Boussinesq limit, it had been possible to decouple the pressure entirely from the conservation equations for mass and vorticity, so that no assumptions whatsoever had been required regarding the pressure. For non-Boussinesq currents, on the other hand, the pressure does not decouple from the vorticity transport equation, so that a certain amount of information regarding the pressure is needed for the exact integration of the vorticity equation over a finite control volume. Towards this end, we stipulate that the pressure distributions inside the current is hydrostatic. Furthermore, we assume the pressure inside the current to be constant along the wall, since the current fluid is considered to be at rest. On this basis we obtain a closed-form solution for the Froude number of non-Boussinesq gravity currents that reduces to the correct expression derived for the Boussinesq limit.

In order to assess the accuracy of the predictions by the various models for non-Boussinesq flows, we analyze the rate at which vorticity is convected out of the control volume. For full-depth currents, the prediction by the vorticity model is close to that of Benjamin's model, and both are very close to corresponding high-resolution simulation data. For partial depth currents, the vorticity model agrees closely with simulation data. We show that Benjamin's model predictions also reproduce the DNS results with good accuracy, which indicates that the simulated flow satisfies Benjamin's assumption of vanishing headloss to a good approximation. Hence, the key contribution of the vorticity

model should be seen in its ability to predict the front velocity without any energy-based closure assumptions, rather than in its improved accuracy.

We furthermore discuss the influence of the three main assumptions underlying all of the above models, including the present vorticity-based model, regarding the nature of the flow, *viz.* i) the flow is steady in the reference frame moving with the current front; ii) the flow is inviscid; and iii) the fluid inside the current is at rest. We find the quasisteady flow assumption to be very accurate in the neighborhood of the front of the top current, although unsteady effects increase farther downstream. The influence of viscosity is significant near the front, but very small further downstream. The effects of the fluid motion inside the current are small at an intermediate distance of a few current heights behind the tip, but they increase both farther downstream and in the immediate neighborhood of the tip. For a constant density ratio, the model predictions generally improve with increasing Reynolds number, while for a constant Reynolds number they improve for weaker density contrasts. We furthermore show that the effects of the above three assumptions partially cancel each other out with regard to the predicted vorticity flux, which explains the good agreement with simulation data across the entire range of Reynolds numbers and density ratios investigated.

Acknowledgements

This work emerged from a collaboration between EM, SGLS and James Rottman at the 2013 Woods Hole summer program on Geophysical Fluid Dynamics. EM and SGLS are grateful to James Rottman and Paul Linden for a number of helpful discussions during this program. EM acknowledges support through NSF grant CBET-1335148.

Appendix A. Extension to three-dimensional flows

The steady-state Euler equation can be written in conservative form as

$$\nabla \cdot (\rho \mathbf{u} \mathbf{u}) + \nabla p = \rho \mathbf{g} . \quad (\text{A } 1)$$

We define L as the z -component of the curl of (A 1). We take y as the vertical direction so that $\mathbf{g} = (0, -g, 0)$, and denote the velocity components by $\mathbf{u} = (u, v, w)$. Then we can write $L = \nabla \cdot \mathbf{q}$, where

$$\mathbf{q} = \begin{pmatrix} q_x \\ q_y \\ q_z \end{pmatrix} = \begin{pmatrix} g\rho + \partial_x(uv\rho) + \partial_z(vw\rho) + \frac{1}{2}\partial_y[\rho(v^2 - u^2)] \\ -\partial_y(uv\rho) - \partial_z(uw\rho) + \frac{1}{2}\partial_x[\rho(v^2 - u^2)] \\ 0 \end{pmatrix} . \quad (\text{A } 2)$$

Here \mathbf{q} is arbitrary up to a gauge transformation, so that we can add the curl of any vector field to \mathbf{q} and still have $L = \nabla \cdot \mathbf{q}$. We have used this choice so that $q_z = 0$.

If the system is two-dimensional or periodic in the spanwise direction, then on applying the divergence theorem there is no boundary in the z -direction, so that there will be no contribution from q_z . If side walls are present, we need to integrate q_z on the walls. However, the integral is zero, so that again there is no contribution.

Now consider q_y on the basal and top surfaces where $v = 0$. Using the continuity equation,

$$q_y = -u\rho\partial_y v - \partial_z(uw\rho) - \frac{1}{2}\partial_x(\rho u^2) = u\rho\partial_z w - \partial_z(uw\rho) - \frac{1}{2}u^2\partial_x\rho . \quad (\text{A } 3)$$

When we integrate with respect to z , the $\partial_z(uw\rho)$ term will be zero whether the system is periodic or has sidewalls. All of these terms are likely to be small in the limit of time

averaging, low diffusion and high Reynolds number. This follows for the $u\rho\partial_z w$ term in the periodic case through symmetry arguments, although in the presence of sidewalls there may be some mean contribution.

Finally, we have the q_x term to consider. The y -integration of $\frac{1}{2}\partial_y(\rho v^2)$ will be zero, since v vanishes on the top and basal surfaces. Integration of $\partial_z(vw\rho)$ in z results in zero for both periodic boundaries and sidewalls ($w = 0$). With these simplifications, q_x can be written as

$$q_x = g\rho + \partial_x(uv\rho) - \frac{1}{2}\partial_y(\rho u^2) . \quad (\text{A } 4)$$

In the same manner as in section 2.2, we assume that the integration in y and z of $\partial_x(uv\rho)$ is zero. Then

$$q_x = g\rho - \frac{1}{2}\partial_y(\rho u^2) , \quad (\text{A } 5)$$

so that we recover the same result for three dimensions as we had in two dimensions.

Appendix B. Vorticity flux deviations for the primitive variable approach

An equation corresponding to (3.14) for the vorticity approach will now be derived for the primitive variable approach in two dimensions. We start with the conservative equations

$$\partial_t(\rho u) + \partial_x(\rho u^2) + \partial_y(\rho uv) + \partial_x P - \mu\nabla^2 u = 0 , \quad (\text{B } 1)$$

$$\partial_t(\rho v) + \partial_x(\rho uv) + \partial_y(\rho v^2) + \partial_y P - \mu\nabla^2 v + \rho g = 0 . \quad (\text{B } 2)$$

Taking the z -component of the curl, dividing by $\rho_1 U_b^2$ and integrating over BCDE gives

$$E_t^* + \frac{1}{\rho_1 U_b^2} \oint \mathbf{q} \cdot \hat{\mathbf{n}} \, dl + E_\mu^* = 0 , \quad (\text{B } 3)$$

where

$$\mathbf{q} = \begin{pmatrix} gp + \partial_x(\rho uv) + \frac{1}{2}\partial_y(\rho(v^2 - u^2)) \\ -\partial_y(\rho uv) + \frac{1}{2}\partial_x(\rho(v^2 - u^2)) \end{pmatrix} , \quad (\text{B } 4)$$

and $U_b = \sqrt{g(1-\sigma)H}$ is the buoyancy velocity. The terms E_t^* and E_μ^* are given by

$$E_t^* = \iint \nabla^* \times (\partial_t(\rho^* \mathbf{u}^*)) \, dA^* , \quad (\text{B } 5)$$

$$E_\mu^* = -\frac{1}{Re} \iint \nabla^* \times (\nabla^{*2} \mathbf{u}^*) \, dA^* . \quad (\text{B } 6)$$

Integrating the second term in (B 3), and using the fact that $v = 0$ on the top and bottom walls, as well as $u_C = u_D$, gives

$$\begin{aligned} \oint \mathbf{q} \cdot \hat{\mathbf{n}} \, dl &= \int_{DE+BC} \frac{1}{2} u^2 \partial_x \rho \, dx + \int_{EB+CD} g\rho \, dy + \int_{EB+CD} \partial_x(\rho uv) \, dy \\ &\quad - \frac{1}{2} \rho_B u_B^2 + \frac{1}{2} \rho_E u_E^2 . \end{aligned} \quad (\text{B } 7)$$

The vorticity flux is defined as

$$\Omega = \int_{EB} u(\partial_x v - \partial_y u) \, dy = \int_{EB} u \partial_x v \, dy - \int_{EB} \frac{1}{2} \partial_y(u^2) \, dy = \int_{EB} u \partial_x v \, dy - \frac{1}{2}(u_B^2 - u_E^2) . \quad (\text{B } 8)$$

Consequently

$$\begin{aligned} -\frac{1}{2}\rho_B u_B^2 + \frac{1}{2}\rho_E u_E^2 &= \frac{1}{2}\rho_E(u_E^2 - u_B^2) + \frac{1}{2}u_B^2(\rho_E - \rho_B) \\ &= \rho_E \Omega - \rho_E \int_{EB} u \partial_x v \, dy + \frac{1}{2}u_B^2(\rho_E - \rho_B). \end{aligned} \quad (\text{B 9})$$

Substituting (B 9) into (B 7) gives

$$\begin{aligned} \oint \mathbf{q} \cdot \hat{\mathbf{n}} \, dl &= \int_{DE+BC} \frac{1}{2}u^2 \partial_x \rho \, dx + \int_{EB+CD} g \rho \, dy + \int_{EB+CD} \partial_x(\rho uv) \, dy \\ &\quad + \rho_E \Omega - \rho_E \int_{EB} u \partial_x v \, dy + \frac{1}{2}u_B^2(\rho_E - \rho_B). \end{aligned} \quad (\text{B 10})$$

We can divide by $\rho_1 U_b^2$ and substitute this back into (B 3). Since $\rho_E = \rho_1$ for a top current

$$\begin{aligned} \frac{1}{\rho_1 U_b^2} \oint \mathbf{q} \cdot \hat{\mathbf{n}} \, dl &= \int_{DE+BC} \frac{1}{2}u^{*2} \partial_x \rho^* \, dx^* + \int_{EB+CD} \frac{\rho^*}{1-\sigma} \, dy^* + \int_{EB+CD} \partial_x(\rho^* u^* v^*) \, dy^* \\ &\quad + \Omega^* - \int_{EB} u^* \partial_x v^* \, dy^* + \frac{1}{2}u_B^{*2}(1 - \rho_B^*). \end{aligned} \quad (\text{B 11})$$

The second integral on the RHS can be evaluated piecewise

$$\int_{EB+CD} \frac{\rho^*}{1-\sigma} \, dy^* = \frac{1}{1-\sigma} ((1-h^*) + \sigma h^* - 1) = \frac{1}{1-\sigma} (h^*(\sigma - 1)) = -h^* = -\Omega_C^*. \quad (\text{B 12})$$

We can define the error in this piecewise evaluation such that

$$\int_{EB+CD} \frac{\rho^*}{1-\sigma} \, dy^* = E_C^* - \Omega_C^*. \quad (\text{B 13})$$

Substituting (B 13) into (B 11) and then employing this in (B 3) gives an equation for Ω^*

$$\begin{aligned} \Omega^* &= \Omega_C^* - E_C^* - E_t^* - E_\mu^* - E_a^* \\ &\quad - \frac{1}{2}u_B^{*2}(1 - \rho_B^*) - \int_{DE+BC} \frac{1}{2}u^{*2} \partial_x \rho^* \, dx^* + \int_{EB} u^* \partial_x v^* \, dy^*, \end{aligned} \quad (\text{B 14})$$

where

$$E_a^* = \int_{EB+CD} \partial_x(\rho^* u^* v^*) \, dy^*. \quad (\text{B 15})$$

Although equation (B 14) contains some terms that are very similar to those in (3.14), it is generally more complicated, so that in the main body of this work we chose to evaluate the deviations from the model using the vorticity approach.

REFERENCES

- ANJUM, H.J., MCELWAIN, J.N. & CAULFIELD, C.P. 2013 The instantaneous froude number and depth of unsteady gravity currents. *J. Hydr. Res.* **51** (4), 432–445.
- BENJAMIN, T.B. 1968 Gravity currents and related phenomena. *J. Fluid Mech.* **31**, 209–248.
- BIRMAN, V.K., MARTIN, J.E. & MEIBURG, E. 2005 The non-Boussinesq lock-exchange problem. part 2. high-resolution simulations. *J. Fluid Mech.* **537**, 125–144.
- BORDEN, Z. & MEIBURG, E. 2013 Circulation based models for Boussinesq gravity currents. *Phys. Fluids* **25** (10).
- CHEN, C. & MEIBURG, E. 2002 Miscible displacements in capillary tubes: Influence of Korteweg stresses and divergence effects. *Phys. Fluids* **14** (7), 2052–2058.

- VON KÁRMÁN, T. 1940 The engineer grapples with nonlinear problems. *Bulletin of the American Mathematical Society* **46** (8), 615–683.
- LOWE, R.J., ROTTMAN, J.W. & LINDEN, P.F. 2005 The non-boussinesq lock-exchange problem. part 1. theory and experiments. *J. Fluid Mech.* **537**, 101–124.
- MCELWAIN, J.N. 2005 Rotational flow in gravity current heads. *Phil. Trans. Royal Society London A* **363** (1832), 1603–1623.
- POZRIKIDIS, C. 1997 *Introduction to theoretical and computational fluid dynamics*. Oxford University Press.
- SAFFMAN, P.G. 1992 *Vortex dynamics*. Cambridge University Press.
- WILLIAMSON, J.H. 1980 Low-storage Runge-Kutta schemes. *J. Comp. Phys.* **35** (1), 48–56.