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1 **Existence of the Threshold Pressure in the Land-Atmosphere Interaction**

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6

7 **Abstract**

8 Excitation of seismic waves by atmospheric pressure changes is examined from data for
9 two tropical cyclones, Tropical Storm Lee (2011) and Hurricane Isaac (2012). They
10 moved through the Earthscope Transportable Array (USArray) and generated variations
11 in pressure and ground motions that spanned 4-5 orders of magnitude in power spectral
12 density (PSD). For vertical seismic ground velocity PSD (S_V) for frequencies between
13 0.01 and 0.02 Hz, there is a threshold pressure at about pressure PSD (S_P) of 10 (Pa^2s),
14 below which vertical motion is not affected by local atmospheric pressure. Above this
15 threshold pressure, vertical ground motion increases with surface pressure as $S_V \sim S_P^{1.5}$. In
16 order to understand the land-atmosphere interaction, pressure above this threshold is the
17 only useful range. Horizontal-component PSDs are about two orders of magnitude larger
18 than vertical-component PSDs and change with pressure for its entire range. This overall
19 trend is most likely caused by ground tilt.

20

21 **Key Points:**

- 22 • In the excitation of seismic waves by atmospheric pressure changes, there is a
23 critical, threshold pressure.
- 24 • Below the threshold, vertical amplitudes are not affected by local atmospheric
25 pressure changes.
- 26 • Horizontal amplitudes show the effects of tilt for the whole pressure range.

27 **1. Introduction**

28 How seismic signals are generated by the land-atmosphere interaction is an old
29 question [e.g., Tanimoto et al., 2015]. It is a difficult question, mainly because of a lack
30 of good, critical data sets. Good data in this case means a dense network of seismometers
31 and barometers. We have noted that the Earthscope Transportable Array (TA hereafter)
32 could provide unique data sets to address this question, although the principal purpose of
33 TA was to improve our understanding of structure in the solid Earth. TA data became
34 useful for the land-atmosphere interaction study after 2010, because high-quality
35 barometers (SEED channel LDO) and infrasound sensors (SEED channel LDF) were
36 added to this network (<http://www.earthscope.org/science/observatories/usarray>). We use
37 the barometer data in this paper. Consistent results were obtained with the infrasound
38 sensor data. Comparison between barometer and infrasound sensors is shown in Figure
39 S1 (supplement) to support this point. Another pressure sensor, the MEMS pressure
40 sensor (channel LDM), turned out to be inadequate for the frequency range (0.01-0.02
41 Hz) of this study.

42 In this paper, we focus on data for two tropical cyclones, Hurricane Isaac (2012)
43 and Tropical Storm Lee (2011) that moved through the TA after their landfalls. Seismic
44 and barometric data from these cyclones provide us unusual opportunities to observe the
45 response of solid Earth generated by surface atmospheric pressure. Seismic ground
46 motions and surface pressures varied 4-5 orders of magnitude in PSD as these hurricanes
47 passed by.

48 We performed some analyses on Hurricane Isaac (Tanimoto and Lamontagne,
49 2014; Tanimoto and Valovcin, 2015) but in this paper we apply a different approach in

50 order to understand some basic characteristics in the land-atmosphere interactions. In this
51 paper, we only examine the co-located barometer and seismometer data and monitor how
52 they change. The underlying idea is that the largest effects of atmospheric pressure
53 should show up most clearly in the co-located seismic sensors. Despite the simplicity in
54 this approach, we find quite interesting features in the relationships between surface
55 pressure and ground motions. The most important point is the identification of the
56 critical, threshold pressure; below this pressure, vertical ground motions are constant
57 which means that seismic amplitudes are independent of changes in local atmospheric
58 pressure. Above this pressure, ground motions increase with pressure. It [shows](#) that there
59 exists a threshold atmospheric pressure, above which atmospheric pressure overwhelms
60 other sources of seismic noise.

61 We will describe the data and our approach in section 2, three main characteristics
62 in data in section 3 and our interpretations in section 4.

63 **2. Data and Our Approach**

64 Figure 1 shows the tracks of Hurricane Isaac and Tropical Storm Lee in the top
65 panels. Red circles in top panels show the locations of stations (TA and some permanent
66 stations) that had both seismometer and barometer data. Blue circles are stations with
67 seismometers only. Since barometers were installed starting in mid-2010, only the eastern
68 half had barometers at the time of Lee (Figure 1, top-right).

69 Although the tracks of both tropical cyclones are near the edge of the TA, we
70 could confirm that seismic amplitudes and pressure variations are consistent with
71 (approximate) cylindrical symmetry, at least for available azimuths, and they decreased
72 with distance from the centers. The bottom panels show examples for selected time

73 intervals; we chose UTC 08:00, August 29, 2012 for Isaac (Figure 1 bottom, left) and
74 UTC 10:00, September 3, 2011 for Lee (Figure 1 bottom, right). Each circle is an average
75 PSD for frequencies between 0.01 and 0.02 Hz. Seismic velocity PSDs are shown in blue
76 with scale on the left and pressure PSDs are shown in red with scale on the right.

77 We mainly focus on this low frequency range (0.01-0.02 Hz) because seismic and
78 pressure amplitudes decay systematically with distance from the cyclone centers. Strictly
79 speaking, the amplitudes peak at about 50-100 km from the centers, presumably at the
80 location of the eyewall [Tanimoto and Lamontagne, 2014], and decay outward. Most
81 stations turn out to be outside this (eyewall) peak.

82 Figure 1 shows only vertical-component seismic data (bottom panels). For
83 comparison, we show Figure S2 (supplement) that shows amplitude-distance variations
84 of three component seismic data (0.01-0.02 Hz) at UTC 12:00, August 29, 2012, for
85 Isaac. Similar amplitude decay trends are seen for all components but horizontal data
86 contain much larger scatter.

87 We checked higher frequency signals up to the microseism frequency bands
88 (0.05-0.5 Hz) [Tanimoto and Valovcin, 2015] but these decaying trends with distance
89 were lost in high frequency signals. It appears that higher frequency waves are mostly
90 generated by ocean waves that are excited by hurricane winds, and thus the source area
91 for these high-frequency waves seems quite broad. On the other hand, the amplitudes in
92 our chosen frequency band (0.01-0.02 Hz) show that they decay with distance from the
93 center and support the view that they were generated close to the center of the hurricanes.

94 Two bottom panels in Figure 1 show that the influence zone of hurricanes is
95 mostly within 1000 km from their centers, with particularly large effects confined to the

96 innermost 500 km. Some deviations to this statement can be recognized outside 1000 km
97 as there is a secondary peak of pressure about 1500 km [Figure 1, bottom panels].
98 Associated seismic amplitudes to these pressure variations are quite small and remain
99 within the scatter of short-distance (<1000 km) data [Figure S3, supplement]. We believe
100 these secondary peaks around 1500 km were caused by spiral winds and rain bands that
101 extend outward from the central region. But since they do not bring much information on
102 the land-atmosphere interaction, as evidenced in Figure S3, we focus our analysis on data
103 within 1000 km from the cyclone centers.

104 **3. Pressure PSD vs. Ground Velocity PSD**

105 Figure 2 shows plots of surface pressure PSD (horizontal axis) vs. ground velocity
106 PSD (vertical axis). Three-component ground velocity PSDs are indicated by three
107 colors, vertical (Z) in blue, radial (R) in red and transverse (T) in black. Radial and
108 transverse components were obtained by using the locations of the center of Isaac and
109 Lee, reported in Brown [2011] for Lee and Berg [2013] for Isaac respectively.

110 Each point in Figure 2 represents PSDs computed for a time-series length of 1
111 hour. The entire time interval of data that was used to create Figure 2 was three days
112 (August 29-31, 2012 for Isaac and September 3-5, 2011 for Lee).

113 Vertical-component data (blue) and horizontal-component data (red and black)
114 make two separate clusters in Figure 2 when plotted against surface pressure from the co-
115 located barometers. Horizontal-component PSDs are typically larger than vertical-
116 component PSDs by about 2-3 orders of magnitude. Green dash lines in Figure 2 were
117 determined by the least squares, fitting the formula $\log_{10}(S_V) = A \log_{10}(S_P) + B$ for different
118 pressure ranges. In this formula, S_V is the ground velocity PSD and S_P is the surface

119 pressure PSD. The coefficients determined by this fitting process (A and B) are
120 summarized in Table 1. In total, there are five independent lines in Figures 2 and 3 and
121 each line is denoted by its name (V_g , V_{L1} , V_{L2} , H_g and H_L).

122 Both vertical and horizontal data were fit separately below and above the
123 threshold pressure (PSD) $S_p=10$ (Pa^2s). This threshold pressure was first chosen from
124 vertical-component data that show clear a break in the data. We overlay the vertical PSDs
125 from two cyclones in Figure 3 (top). Because Isaac was much stronger than Lee, we can
126 see more points in higher pressure ranges for Isaac but the threshold pressure seems to
127 agree between the two cyclones.

128 By fitting data from both cyclones above $S_p=10$, the dash line denoted by V_g was
129 obtained. For the vertical-component data below this threshold value, we obtained V_{L2} .
130 The latter is constant as the coefficient A was set to zero. There is a slight difference on
131 this constant value between Isaac and Lee. In order to indicate this difference, we denote
132 the value for Isaac by V_{L1} (Table 1) but it is not significantly different from V_{L2} that was
133 determined from the combined vertical-component data. But this difference indicates that
134 the background noise level, created by other noise sources, varies seasonally and
135 sometimes year to year. If we took into account the differences between these flat noise
136 levels from two cyclones, the threshold value ($S_p=10$) can vary from $S_p=5$ to 20
137 approximately.

138 Existence of a threshold value is not so obvious in horizontal-component data in
139 Figure 2. It is partly because an overall trend in horizontal data shows a large gradient for
140 the entire pressure range (Figures 2 and 3). We believe this overall trend in gradient is
141 caused by the well-known ground tilt. Tilt causes the same effect with horizontal

142 acceleration and is particularly large in low-frequency bands below 0.02 Hz [e.g., Aki
143 and Richards, 2002; Farrell, 1969; Rodgers, 1968].

144 There is an additional feature in horizontal data; if we overlay data from two
145 tropical cyclones (Figure 3, bottom), there is a hint that the gradient becomes steeper as
146 pressure increases. The least squares fits below and above $S_p=10$ (lines H_g and H_L)
147 clearly show a steepening trend in gradient. Although we used two lines to fit horizontal
148 data in Figure 3, in terms of underlying physical processes, it is hard to imagine a
149 threshold pressure for horizontal data that causes a sudden change. We interpret that this
150 gradient increase occurs gradually.

151 But why does the gradient in horizontal data increase with pressure? We speculate
152 that there exists a direct wind effect for high pressure ranges. In general, pressure
153 fluctuation for a frequency range 0.01-0.02 Hz is controlled by winds and is nearly
154 proportional to the square of wind velocity. Therefore, some effects of wind are already
155 included in pressure changes. But when the wind becomes strong, it can exert forces
156 directly on nearby trees and observational facilities and generate additional ground tilt.
157 This should be in addition to surface pressure changes and thus could be a cause for an
158 increase in gradients in Figure 3. However, this is a speculation and details are hard to
159 verify with current data sets.

160 In Figure 4, we show similar seismic amplitudes vs. pressure plot for Tropical
161 Storm Lee for four different frequencies, 0.01-0.02 Hz (top left), 0.04-0.05 Hz (top right),
162 0.09-0.10 Hz (bottom left) and 0.14-0.15 Hz (bottom, right). Amplitude differences
163 between horizontal-component data and vertical-component data are the largest for 0.01-
164 0.02 Hz and quite large for 0.04-0.05 Hz. Both panels at top show that horizontal

165 amplitudes increase with pressure amplitudes (PSD). The differences in vertical and
166 horizontal amplitudes decrease in higher frequency plots and the correlation between
167 horizontal amplitudes and pressure amplitudes also becomes smaller. In the panel for
168 0.09-0.10 Hz, there may still be a weak correlation for pressure above 1-10 (Pa^2s) but in
169 the 0.14-0.15 Hz plot, seismic amplitudes change little with local surface pressure.
170 Clearly the dominance of local atmospheric effects is confined to low frequencies below
171 about 0.05 Hz. It should also be noted that these higher-frequency signals in the bottom
172 panels are mostly the secondary microseism (seismic noise) that are generated in the
173 oceans [Longuet-Higgins, 1950; Hasselman, 1963].

174 **4. Discussion and Summary**

175 One of the most robust features in our observation is the existence of a threshold
176 pressure in vertical-component data at a pressure PSD of about $S_p=10$ (Pa^2s). Because of
177 scatter in data, this value contains some uncertainties and can vary from $S_p=5$ to 20.
178 Below this threshold pressure, vertical amplitudes do not change with pressure. This lack
179 of correlation means that the local atmospheric pressure is not the main source of seismic
180 ground motion (noise) at the site. These signals below the threshold pressure were
181 generated by processes other than the local atmospheric pressure, such as ocean waves
182 away from the station. The threshold pressure can be viewed as the pressure when the
183 effects of the local atmospheric pressure exceed those of other seismic-noise sources. In
184 order to understand the land-atmosphere interaction in more details, we must focus on the
185 pressure range above this threshold.

186 We take a view that atmospheric pressure acts as an excitation source at Earth's
187 surface for seismic waves. In the whole, coupled Earth system, this view may not apply if

188 phase velocity of atmospheric waves were close to phase velocity of seismic waves in the
 189 solid Earth as the transmission of waves become very efficient between the atmosphere
 190 and the solid Earth. But such a match in phase velocity is not likely to occur as
 191 atmospheric waves have velocities of a few hundred meters per second and seismic
 192 waves have velocities of 3-4 km/s for surface waves and faster body waves. It was
 193 pointed out previously [Sorrells, 1971; Sorrells and Goforth, 1973] that atmospheric
 194 pressure acts almost as a surface load under such a condition. Seismic data show such
 195 amplitude behaviors to first order, although they should also contain some smaller-
 196 amplitude propagating surface waves. But those seismic data are in the near-field and
 197 seem to be dominated by pressure loading effects.

198 The proportionality constant (A in the log-log formula) between S_V and S_P in
 199 Figure 2 (and 3) is not 1 above the threshold pressure. Instead, it is about 1.5 (V_g in Table
 200 1). We interpret this observation as follows; the excitation of seismic ground motion by
 201 atmospheric pressure occurs by a force that can be considered to be a random force. This
 202 is because atmospheric pressure has very short correlation distance on Earth's surface
 203 (about 100 m or less). It changes its sign with short wavelengths of the order of 10-100 m
 204 [e.g., Herron et al., 1969; McDonald et al., 1971]. On the other hand, the pressure source
 205 is spread out over many kilometers. In essence, we have a rapidly fluctuating source that
 206 extends over a large area. In such a case, one can approximate that the excited seismic
 207 ground motion PSDs become proportional to pressure PSD by $S_V \propto L^2 S_P$ where L is the
 208 correlation length in the surface pressure field [e.g., Kobayashi and Nishida, 1998; Fukao
 209 et al., 2002; Tanimoto, 2005; Tanimoto and Valocin, 2015]. In such a model, if the
 210 correlation length L is proportional to $S_p^{0.25}$, the gradient of 1.5 can be explained. This

211 means that the correlation length changes with pressure. Physically, one would expect
212 that larger pressure is related to stronger wind. If strong lateral wind exists, one can
213 imagine that the correlation length in the surface pressure field should become larger as
214 pressure at a location can be transported to nearby location by winds. However, why the
215 exponent becomes 0.25 is left unexplained. Understanding it requires a careful theoretical
216 study.

217

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222

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265

266

267 **Table 1:** Least squares fit by the formula $\log_{10}(Sv)=A \log_{10}(Sp)+B$ for various ranges.
 268 IDs are the same in Figures 2, 3 and 4. Ranges of barometer (pressure) PSD are in the
 269 second column. V_{L1} is for Isaac only but all others were derived for the combined data of
 270 Isaac and Lee.

271

| ID | Range (S_p) | A | B |
|----------|-----------------|-------------------|-------------------|
| V_g | $S_p > 10$ | 1.501 ± 0.001 | -17.20 ± 0.08 |
| V_{L1} | $S_p < 10$ | 0.0 | -15.70 ± 0.02 |
| V_{L2} | $S_p < 10$ | 0.0 | -15.52 ± 0.02 |
| H_g | $S_p > 10$ | 1.261 ± 0.020 | -13.71 ± 0.02 |
| H_L | $S_p < 10$ | 0.618 ± 0.031 | -13.26 ± 0.03 |

272

273 **Figure Captions**

274 **Figure 1:** (top, left) Track of Hurricane Isaac (August, 2012) and seismic stations from
 275 Earthscope. Black circles are the locations of its center at every six hours. Green circles
 276 indicate the midnight of each day. Red circles indicate stations had barometer and
 277 seismometer. Blue circles indicate stations with seismometer only. (top, right) Track of
 278 Tropical cyclone Lee (September, 2011). (bottom, left) Seismic vertical PSD and
 279 pressure PSD plotted against distance from the center of Isaac. (bottom, right) Seismic
 280 vertical PSDs and pressure PSDs for Lee.

281 **Figure 2:** Seismic amplitudes (PSD) plotted against pressure PSD for every 1-hour
 282 interval. Top is for Hurricane Isaac and bottom is for Tropical Storm Lee. Vertical PSDs
 283 are denoted by blue circles, radial by red and transverse by black. Lines by the least-
 284 squares fit are shown by green dashes. Except for V_{L1} , they were derived from the
 285 combined data set for Isaac and Lee. H_g is for horizontal component data above the
 286 threshold value $S_p=10$. H_L is for horizontal component data below this threshold pressure.
 287 V_g is for vertical component data above the threshold pressure, determined from the
 288 combined data from both tropical cyclones. V_{L1} is for below the threshold for Isaac only.
 289 V_{L2} is for the combined data of Isaac and Lee. The coefficients are in Table 1.

290 **Figure 3:** Same data as in Figure 2 but the data from Isaac and Lee were overlaid. Top is
 291 the vertical component data and bottom is the horizontal component data. Lines are the
 292 same with those in Figure 2.

293 **Figure 4:** Seismic amplitudes (PSD) vs. pressure PSD for four frequency ranges, 0.01-
 294 0.02 Hz (top, left), 0.04-0.05 Hz (top, right), 0.09-0.10 Hz (bottom, left) and 0.14-0.15
 295 Hz (bottom, right). Because of tilt, horizontal component data have much larger

296 amplitudes than vertical component data for lower frequency ranges (0.01-0.02 and 0.04-
297 0.05 Hz) and have good correlation with local pressure data. In higher frequency ranges
298 (0.09-0.10 and 0.14-0.15 Hz), tilt effects are much smaller and vertical and horizontal
299 components have similar amplitudes. In the 0.14-0.15 Hz plot, signals are generated in
300 the ocean and do not show much correlation with local atmospheric pressure.

301

Figure 1.

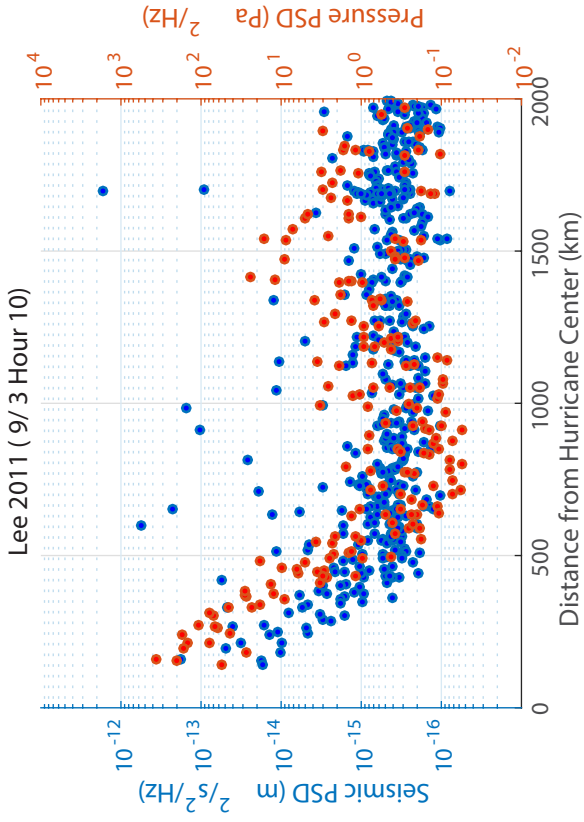
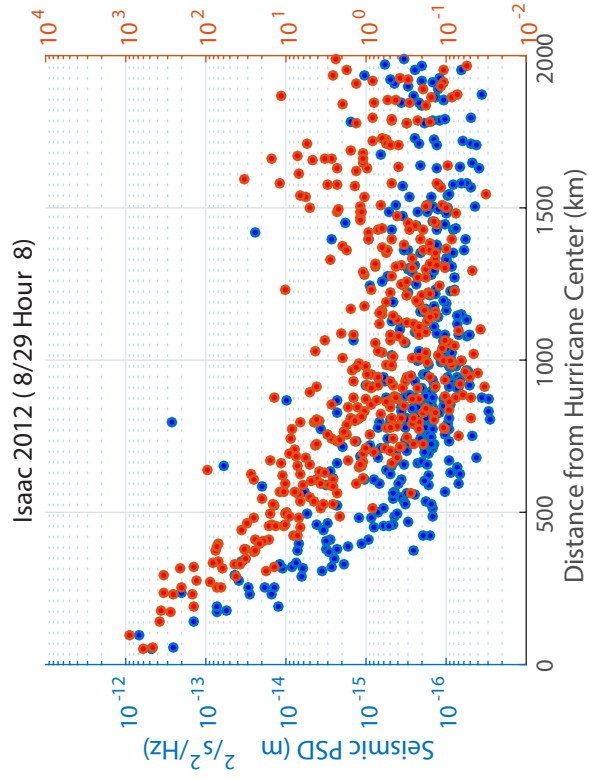
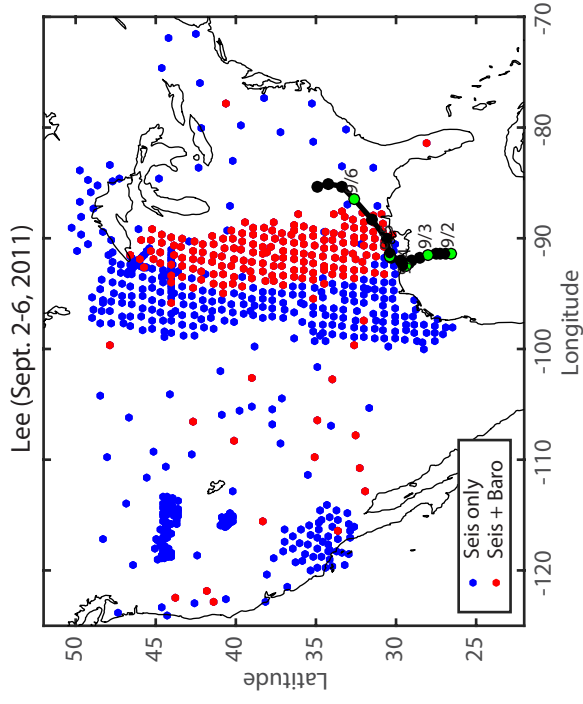
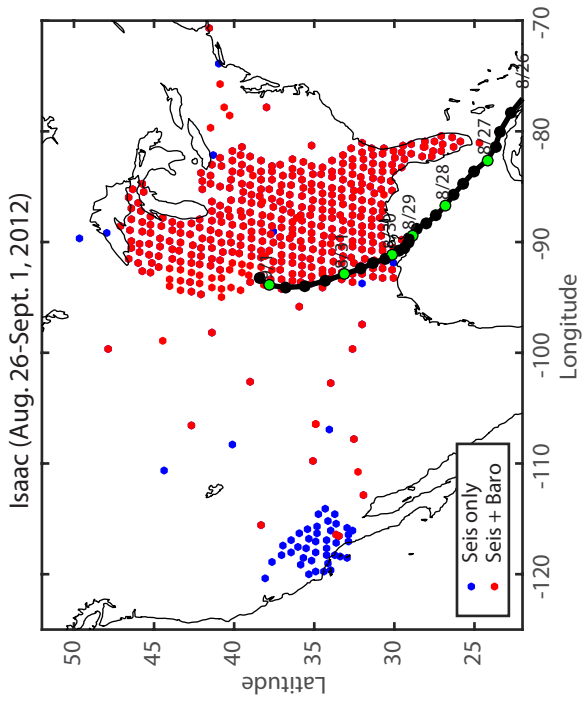


Figure 2.

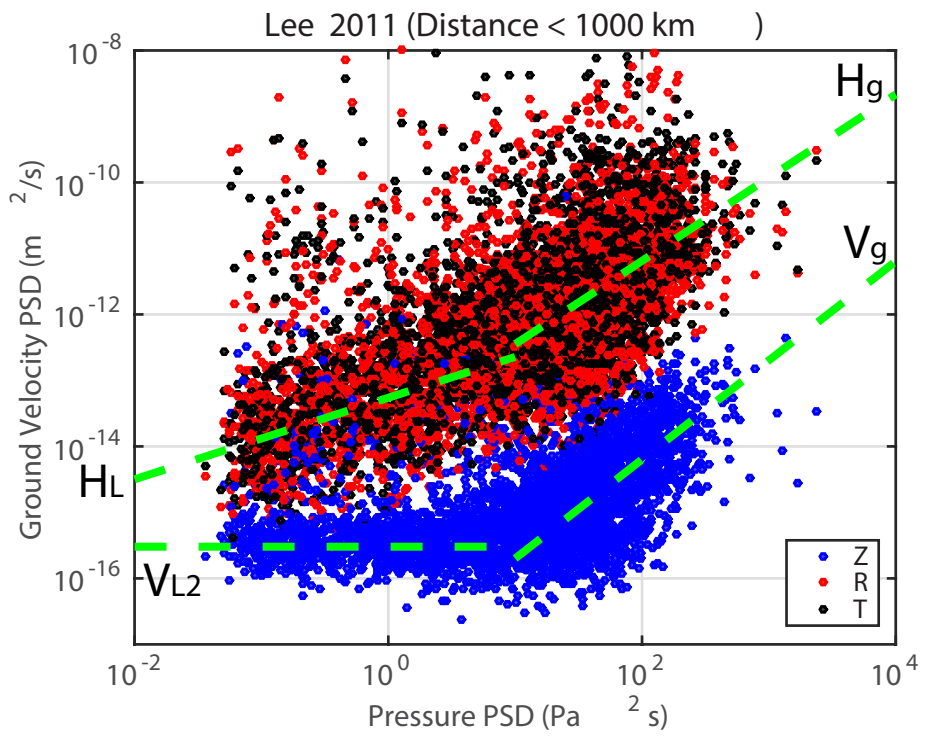
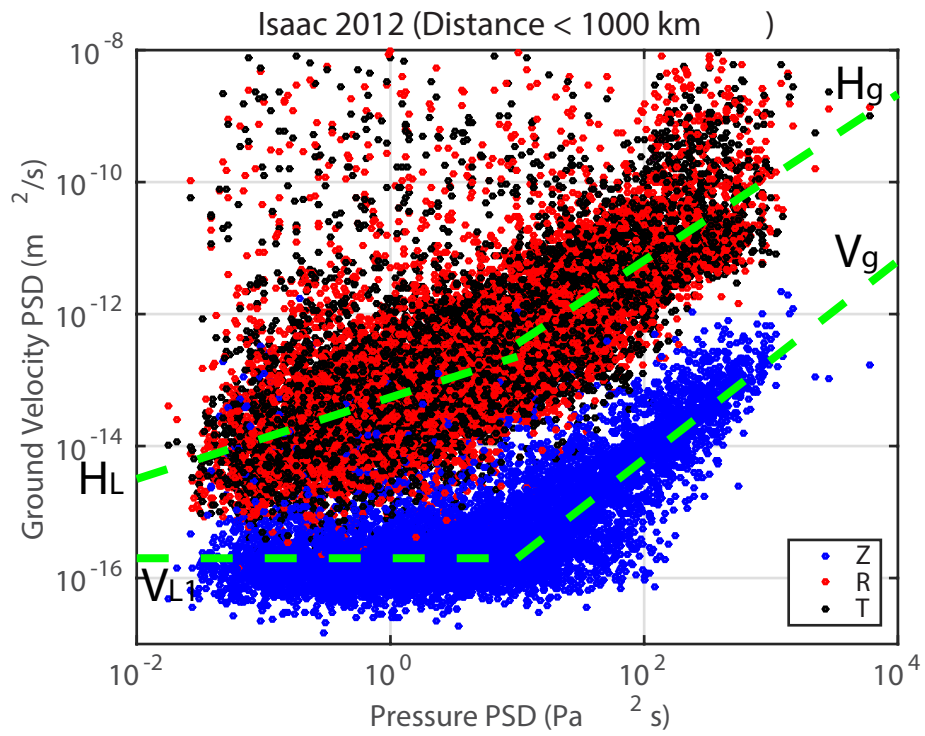
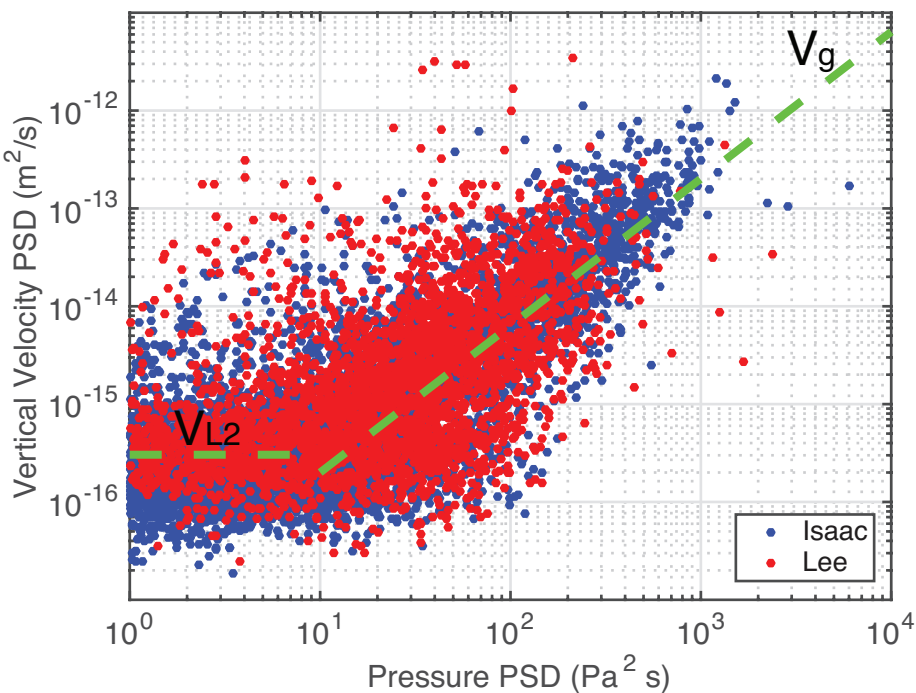


Figure 3.

Vertical PSD (Distance < 1000 km)



Horizontal PSD (Distance < 1000 km)

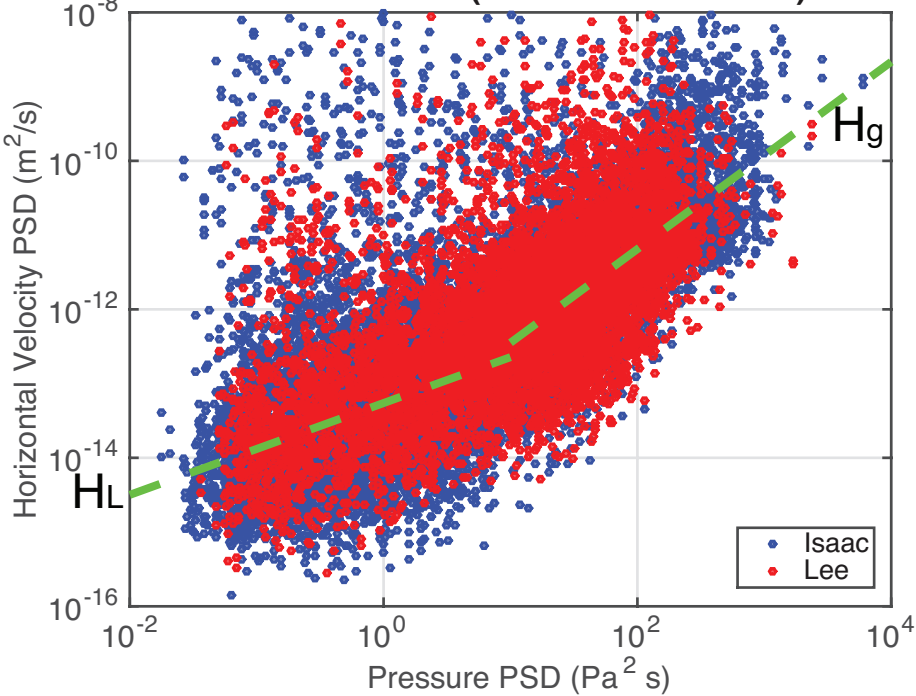


Figure 4.

