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1 Parkfield Earthquakes: Characteristic or Complementary?

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## Abstract

We model the two most recent  $\sim M_w 6$  Parkfield, California, earthquakes, which occurred in 1966 and 2004, from a non-linear global inversion of near-fault strong-motion seismograms. Our rupture models are characterized by spatially variable slip amplitude and rake, rupture velocity, and rise time. The rupture models indicate that the two earthquakes generated slip in regions of the fault that are not identical, as earlier suggested. Given the sparse seismic dataset available for the 1966 earthquake, we conduct a series of tests to verify our results: (1) we perform synthetic tests in order to study the resolution of the 1966 seismic dataset; (2) we perform an inversion of the 2004 earthquake using a dataset equivalent to the 1966 earthquake; and (3) we model the 1966 dataset under the *a priori* assumption that it was similar to the 2004 earthquake. All of the tests, as well as independent observations, indicate that slip during the 1966 and 2004 Parkfield earthquakes occurred in different regions of the fault. This result implies that regions of a fault that are frictionally locked may remain locked even during a mainshock (moderate-size earthquake). In this scenario, large earthquakes occur when all the locked regions of a fault are "synchronized" and ready to slip at the same time.

## 39 1. Introduction

40 One of the open debates in seismology concerns the existence of fault sections that break  
41 repeatedly in a characteristic manner, producing “characteristic earthquakes”. At a primary level,  
42 characteristic earthquakes are defined as earthquakes that occur in the same location, rupture the  
43 same fault length, and have identical faulting mechanisms and seismic moments (*Bakun and*  
44 *Lindh, 1985; Bakun et al., 2005*). At a secondary level, characteristic earthquakes, in addition to  
45 the previous features, nucleate at the same hypocenter and propagate in the same direction (*Bakun*  
46 *and McEvelly, 1984; Bakun and Lindh, 1985; Bakun et al., 2005*). A more stringent definition of  
47 characteristic earthquakes would further require consecutive events to rupture identical patches of  
48 the fault (*Bakun et al., 2005*). The level to which earthquakes are characteristic has important  
49 implications for earthquake prediction and assessment of seismic hazard.

50 The Parkfield section of the San Andreas Fault in California is a prime example of a  
51 characteristic fault section. In historical times it generated  $\sim M_w 6$  right-lateral strike-slip  
52 earthquakes in 1881, 1901, 1922, 1934, 1966 and most recently in 2004 (earthquakes close to  
53 Parkfield may also have occurred in 1877 and 1908) (*Bakun and McEvelly, 1979; Topozada*  
54 *et al., 2002*). Data permit a common epicenter, beneath Middle Mountain (Figure 1), for the 1922,  
55 1934 and 1966 earthquakes (*Bakun and McEvelly, 1984*). Moreover, the 1934 and 1966  
56 earthquakes generated identical regional and teleseismic waveforms; they are thought to have  
57 ruptured unidirectionally to the southeast over the same fault length; and they were preceded by  
58 co-located  $M_L 5.1$  foreshocks that occurred 17 minutes prior to the respective mainshocks (*Bakun*  
59 *and McEvelly, 1979; Bakun and Lindh, 1985*).

60 The 2004 Parkfield earthquake was a surprise for the seismological community - whereas

61 the earthquake had been anticipated for years, its character was not consistent with the  
62 predictions. Unlike previous events, the 2004 earthquake nucleated close to Gold Hill (Figure 1),  
63 which is located 20 km southeast of Middle Mountain (epicenter of earlier  $M_w6$  events), it  
64 ruptured toward the northwest, and it was not preceded by foreshocks or any other precursors  
65 (*Bakun et al.*, 2005; *Johnston et al.*, 2006). A question that remains open is whether the slip  
66 distribution of the 2004 earthquake was similar to that of previous Parkfield  $M_w6$  events. In other  
67 words, did the 2004 Parkfield earthquake rupture the same fault patches as previous events? The  
68 two most recent Parkfield earthquakes – 1966 and 2004 – were recorded by strong-motion  
69 seismographs located close to the fault. Using these records of ground motion we infer the time-  
70 space slip distributions of the two earthquakes and then analyze the similarities between the two  
71 earthquakes.

72 In order to model the 1966 and 2004 Parkfield earthquakes we apply a non-linear  
73 inversion algorithm (*Liu and Archuleta*, 2004; *Liu et al.*, 2006) to strong-motion seismograms  
74 recorded less than 20 km from the fault. The algorithm generates multiple random space-time slip  
75 distributions, which are used to forward compute synthetic ground motion. From the comparison  
76 of the synthetic ground motion with the observed ground motion, the algorithm chooses the  
77 “best” rupture model (i.e., the space-time slip distribution that leads to synthetic ground motion  
78 that most adequately resembles the observed ground motion). Because the velocity structure at  
79 Parkfield is highly heterogeneous (*Eberhart-Phillips and Michael*, 1993; *Thurber et al.*, 2003,  
80 2006) and the site effects are very strong (*Liu et al.*, 2006), we use site amplification factors and a  
81 weighting scheme to correct the data for site amplification and local resonances.

82 The 2004 earthquake rupture model is based on ground motion recorded at 43 stations  
83 well distributed around the fault plane. In contrast, the 1966 earthquake model is based on data

84 recorded at only five stations, all of them located on the southeast end of the rupture plane. In  
85 order to understand which features of the 1966 model we can be confident about, we perform  
86 three different tests. The first test makes use of synthetic rupture models. We start by generating  
87 simple rupture models, based on which we compute ground-motion. We then add white noise to  
88 the synthetic ground-motion, and try to recover the initial rupture models by inversion of the  
89 noisy synthetic dataset. The degree to which the initial models can be recovered is an indicator of  
90 the resolution of the dataset. In the second test, we compute a model for the 2004 earthquake  
91 using a dataset equivalent to the 1966 dataset. More specifically, we use data from only five  
92 stations, all located close to those that recorded the 1966 earthquake, to infer a rupture model for  
93 the 2004 earthquake. The differences between the 1966 and 2004 rupture models, as obtained  
94 from data recorded at the same stations, will highlight effective differences between the two  
95 earthquakes. In the last test, we compute a model for the 1966 earthquake while constraining the  
96 slip amplitude to be similar to the 2004 earthquake. We then analyze how well this constrained  
97 model can replicate the observed ground motion. Finally, we compare our results with  
98 independent studies.

## 99 **2. Slip model for the 1966 and 2004 earthquakes**

100 The 1966 earthquake was recorded by only five strong ground-motion seismographs  
101 (*Housner and Trifunac, 1967*), all of them located along a line perpendicular to the fault on the  
102 SE end of the rupture zone (Figure 1). The station closest to the fault – CH2W – recorded a large  
103 peak velocity of  $\sim 35$  cm/s in the fault-normal direction. Unfortunately this instrument did not  
104 record the fault-parallel motion. Previous strong-motion studies of the 1966 Parkfield earthquake  
105 can be grossly divided into two groups: (1) those that postulate shallow slip and model only the

106 fault-normal ground motion recorded at station CH2W (e.g. *Aki* (1968)); and (2) those that model  
107 slip over a wider fault plane ( $\sim 40\text{km} \times 8\text{km}$ , in agreement with aftershock locations) and obtain a  
108 good fit to data recorded at all stations except CH2W, where the ground motion predicted from  
109 the models is always less than what was observed (e.g. *Anderson* (1974)). Most of these earlier  
110 studies (1) model the medium around the fault as a homogeneous infinite space; (2) they assume  
111 spatially uniform slip amplitude, rake angle, rupture velocity and rise time over the fault plane  
112 and (3) they model surface displacements (*Aki*, 1968; *Haskell*, 1969; *Boore et al.*, 1971; *Trifunac*  
113 *and Udawadia*, 1974; *Anderson*, 1974, among others). *Trifunac and Udawadia* (1974) were the first  
114 to apply a least-squares inversion scheme to the seismic data in order to obtain slip amplitude.  
115 They divided the fault plane into seven sections, and allowed each section to have different slip  
116 amplitudes. For the whole fault they postulated uniform rupture velocity and rise time. *Archuleta*  
117 *and Day* (1980) presented a quasi-dynamic model for the 1966 Parkfield earthquake assuming  
118 constant rupture velocity. Previous studies using the strong-motion data were done more than 20  
119 years ago, when methods were not as refined and computers were not as efficient as today. Our  
120 kinematic rupture model is the first resulting from a global inversion of the 1966 available  
121 seismic dataset (i.e., our model results from a search for the global minimum of the entire space  
122 of possible models). Our space-time rupture model allows spatially variable slip amplitude and  
123 rake, rupture velocity, and rise-time. Wave propagation is computed assuming a layered velocity  
124 structure that is different for each side of the fault. We fit ground velocities instead of  
125 displacements, which leads to more control on the distribution of kinematic parameters.

126           Because of the prediction that an earthquake would occur in Parkfield between 1983 and  
127 1993 (*Bakun and McEvilly*, 1984), the geophysical instrumentation in Parkfield was strongly  
128 intensified in 1985 (*Bakun and Lindh*, 1985; *Roeloffs and Lagbein*, 1994; *Roeloffs*, 2000). When

129 the earthquake finally occurred in 2004, it generated an excellent dataset useable for inversions.  
130 In particular, 56 near-fault strong-motion seismographs recorded the earthquake (*Shakal et al.*,  
131 2005, 2006).

132 In a previous paper *Liu et al.* (2006) obtained a rupture model for the 2004 earthquake  
133 from inversion of the strong-motion dataset. Here, we derive a model for the 1966 earthquake  
134 following the same procedure that they used. The data processing and modeling is explained in  
135 depth in *Liu et al.* (2006), thus we will not repeat it here. Instead, we will review the most  
136 important aspects of their modeling and refer the reader to their paper for further details.

## 137 **2.1. Data Processing**

138 We model the 1966 earthquake based on ground motion recorded at five stations –  
139 CH2W, CH5W, CH8W, CH12W and TEMB (Figure 1). The 2004 earthquake is modeled from  
140 ground motion recorded at 43 stations – CH1E, CH2E, CH2W, CH3E, CH3W, CH4AW, CH4W,  
141 COAL, DFU, EFU, FFU, FZ1, FZ3, FZ4, FZ6, FZ7, FZ8, FZ9, FZ11, FZ12, FZ15, GFU, GH1W,  
142 GH2E, GH3E, GH3W, GH5W, JFU, KFU, MFU, PHOB, SC1E, SC2E, SC3E, RFU, TEMB,  
143 VC1W, VC2E, VC2W, VC3W, VC4W, VC5W, VFU. These 43 stations are chosen from the  
144 available 56 based on data quality (*Liu et al.*, 2006). The strong-motion stations record ground  
145 accelerations that we integrate to obtain ground velocity. The velocity waveforms are filtered in  
146 the passbands 0.25-1 Hz (1966 earthquake) and 0.16-1 Hz (2004 earthquake) with a 4-pole zero-  
147 phase (forward and backward) Butterworth filter. The lower limit of the passband is based on the  
148 quality of the ground motion record; the upper limit is chosen based on the quality of the velocity  
149 structure approximation we use.



150 We use an approximation to site effects in order to correct for amplifications and resonances  
151 (*Liu et al.*, 2006). This correction is based on data recorded by the Parkfield array during the 1983  
152  $M_w$ 6.5 Coalinga earthquake, which occurred 25 km NE of Parkfield. The correction has two  
153 components:

154 1. Amplification. This correction is applied when the ground motion is amplified over all the  
155 frequency range of interest to us. Based on observations of the Coalinga earthquake, we  
156 compute the factor by which ground motion is amplified at each station due to site  
157 conditions. We then divide each observed waveform by the corresponding station  
158 amplification factor. This correction is frequency-independent.

159 2. Resonance. We account for frequency resonances (i.e., amplification of ground motion at  
160 particular frequencies) by using a weighting scheme; stations strongly affected by  
161 resonances are downweighted in our inversion.

162 For the rest of the paper we will use the word “observation” to refer to the velocity  
163 waveforms obtained through this process.

## 164 **2.2. Non-Linear Global Inversion**

165 We use the non-linear simulated annealing inversion scheme of *Liu and Archuleta* (2004)  
166 to infer a rupture model from the observed ground motion. The rupture model is defined by five  
167 source parameters: slip amplitude (amount of slip), rake angle (direction of slip), average rupture  
168 velocity (related to the time it takes for the rupture to propagate from the hypocenter to a given  
169 point on the fault) and rise time (time interval during which a point on the fault slips). The rise  
170 time is defined as the sum of two independent time parameters: the time during which slip

171 accelerates and the time during which slip decelerates. The simulated annealing algorithm starts  
172 by generating very dissimilar random rupture models, for which synthetic ground motion is  
173 computed. By comparing synthetic and observed ground motion, the algorithm proceeds with the  
174 most adequate rupture model. In the next step, the algorithm generates random rupture models  
175 that resemble the best rupture model from the previous iteration. As the number of iterations  
176 increases, the random models become more similar to the previous best model (the amplitude of  
177 the random variations around the preferred model is decreased), allowing for fine-tuning of the  
178 rupture model. Instead of evaluating all possible rupture models, the simulated annealing  
179 algorithm relies on an adequate sampling of the parameter space. The sampled models are  
180 generated in a random way; it is thus possible to obtain multiple preferred rupture models by  
181 making the inversion follow different random paths. In other words, it is possible to arrive at  
182 different final rupture models by sampling different random rupture models as we iterate through.  
183 If the inversion algorithm works correctly, all final rupture models look alike, independently of  
184 the random path that the algorithm took to arrive at the final model.

185         For each different trial rupture model, synthetic ground motion is computed and compared  
186 with data. The goodness of fit between recorded and synthetic waveforms is quantified with a  
187 correlative misfit function (*Spudich and Miller, 1990; Liu et al., 2006, equation 2*). Synthetic  
188 ground motion is obtained by summing the slip contributions from a grid of points on the fault  
189 (167 m x 167 m). Green's functions and the five source parameters are computed on coarser grids  
190 of 0.5 km x 0.5 km and 2 km x 2 km, respectively. Green's functions are computed with the  
191 frequency-wavenumber method of *Zhu and Rivera (2002)* assuming a layered velocity structure  
192 (*Liu et al., 2006*) based on 3D velocity models (*Thurber et al., 2003, 2006*). The layered velocity  
193 structure is different for each side of the fault, thus accounting for material differences between

194 the Franciscan and Salinian blocks on opposite sides of the San Andreas Fault. Both Green's  
195 functions and source parameters are interpolated to the finer grid of 167 m x 167 m, where the  
196 fault slip (as defined by the five source parameters) is convolved with Green's functions in order  
197 to obtain surface velocities.

198         Because little or no co-seismic surface break was observed during either mainshock (1966  
199 (*Smith and Wyss*, 1968) and 2004 (*Rymer et al.*, 2006)), the fault planes that we model are buried  
200 500 m below the surface. In our models, the fault strikes 140°SE and dips 87°SW (*Liu et al.*,  
201 2006). For both mainshocks we model a rupture plane that is 10 km deep and 40 km long, in  
202 accordance with aftershock locations. In order to obtain a better fit between the locations of the  
203 aftershocks and the rupture area of each earthquake, we offset the 1966 rupture plane by 5 km  
204 along-strike to the southeast with respect to the 2004 plane (Figure 1). The coordinates of the  
205 1966 and 2004 hypocenters are, respectively, 35.951°N, 120.507°W, 8.5 km deep and 35.8185°N,  
206 120.3706°W, 8.26 km deep. These locations are very close to previously reported hypocenters  
207 (*McEvelly*, 1966; *Thurber et al.*, 2006), and they are adjusted so that both epicenters fall on the  
208 same previously chosen fault plane (which fits the microseismicity).

209         Following *Liu et al.* (2006) we impose smoothness constraints on the slip to avoid  
210 physically unrealistic abrupt variations in the rupture model. Constraints are also imposed on the  
211 seismic moment; otherwise the large areas of the fault that slip by small amounts and are not well  
212 resolved lead to a spuriously high seismic moment (*Liu et al.*, 2006).

### 213 **2.3. Rupture Models**

214         For both the 1966 and 2004 earthquakes, we compute 10 different preferred rupture

215 models that are obtained by sampling different random models throughout the inversion (E-supp:  
216 1966 mainshock – Figure fs01 and Tables ts01-ts10; 2004 mainshock – Figure fs02 and Tables  
217 ts11-ts20). All 10 rupture models are equally adequate, in the sense that they all have the ability  
218 to reproduce equally well the observed ground motion according to a goodness of fit for all data  
219 criterion. For each mainshock we show: (A) the rupture model that generates synthetic ground-  
220 motion that best fits the observations; and (B) the average, (C) the standard deviation, and (D) the  
221 coefficient of variation of the 10 preferred models. Given that all 10 preferred models are  
222 similarly good in fitting the observations, we find that the most robust slip distribution is given by  
223 their average. Despite showing a more diffuse image of slip on the fault, the average slip model  
224 (B) only retains the coherent features of the 10 preferred models. The standard deviation (C) is a  
225 measure of the variability between the 10 preferred models, and the coefficient of variation (D)  
226 indicates how variable the preferred models are with respect to the average.

227         Our rupture model for the 1966 earthquake (Figure 2) indicates that co-seismic slip  
228 occurred primarily at shallow depth and toward the SE end of the fault plane. In contrast, in 2004  
229 peak slip occurred in a small region beneath Gold Hill surrounding the hypocenter, and further  
230 significant slip occurred 10-25 km NW of Gold Hill, at a depth between 2 and 10 km (Figure 3).  
231 We consider these features robust, as they appear in all 10 preferred rupture models (E-supp:  
232 Figures fs01 and fs02). As expected, the variability between the ten 1966 preferred models, as  
233 given by the standard deviation, is much larger than the variability between the 2004 preferred  
234 models. This is a reflection of the differences in the seismic datasets available to study the two  
235 mainshocks. The few stations that recorded the 1966 earthquake do not provide a good azimuthal  
236 coverage of the fault. It is thus important to understand which features of co-seismic slip the 1966  
237 dataset can resolve.

### 238 **3. Discussion of the 1966 Rupture Model**

239 Our rupture models for the 1966 and 2004 earthquakes indicate that the two most recent  
240 Parkfield events ruptured fault patches that do not overlap but rather complement each other.  
241 However, the quality of the rupture models for the two earthquakes is very different – based on a  
242 better dataset, the 2004 model is better resolved than the 1966 model. How confident can we be  
243 that that the 1966 and 2004 slip distributions are in fact different?

#### 244 **3.1. Synthetic Tests**

245 In order to understand the resolution of the 1966 dataset, we perform the following synthetic  
246 test:

- 247 1. We generate synthetic rupture models, which we will refer to as input rupture models (they  
248 will serve as input to the synthetic tests) – Figures 4A and 5A.
- 249 2. We generate ground motion at the 5 stations that recorded the 1966 earthquake, based on  
250 the input rupture models we created in Step 1. We will refer to this synthetic dataset as  
251 input dataset.
- 252 3. We add white noise (arbitrarily we choose the amplitude of the white noise to be 20% of  
253 the maximum amplitude in the synthetic waveform) to the input dataset created above; thus  
254 we simulate more realistic Earth-like conditions where data is contaminated by noise.
- 255 4. We use the non-linear inversion algorithm to infer a rupture model (output model) from the  
256 noisy input dataset created in Step 3.

257 5. For each input model we compute 10 output rupture models, as we do for the inversions of  
258 real data. The 10 models are equally adequate, in the sense that the ground motion  
259 generated by all models matches the input noisy synthetic records similarly well. Our final  
260 output model is the average of the 10 preferred models – Figures 4B and 5B.

261 The comparison between input and output test models suggests which features of the 1966 co-  
262 seismic slip we can recover given the available dataset. According to the synthetic tests, if slip  
263 occurs only in one small patch of the fault plane, then we are able to retrieve the correct slip  
264 distribution, independent of where slip occurs (Figure 4). However, if more than one patch of the  
265 fault slips during the earthquake, then slip that occurs on the NW end of the fault plane becomes  
266 almost unrecoverable (Figure 5). For this reason, we will limit our analysis to the SE portion of  
267 the fault plane. Note that these synthetic tests do not show a tendency for shallow slip in the  
268 output models.

### 269 **3.2. Inversion of the 2004 Earthquake Using a Limited Dataset**

270 The results of the synthetic tests are encouraging. However, the simple input models we  
271 used in the previous section (with geometrically simple regions of slip and uniform rake angle,  
272 rupture velocity and rise time) lack the complexity of a real earthquake rupture. Also,  
273 uncertainties in the velocity structure are left out in the synthetic tests. In order to assess the  
274 resolution of the 1966 earthquake given a real earthquake rupture and uncertainties in the  
275 inversion process (e.g., uncertainties in the Green's functions), we perform a second test where  
276 we invert the 2004 earthquake using only a limited dataset. For this purpose, we choose five  
277 stations (CH2W, CH4AW, CH6W, CH12W, TEMB) that recorded the 2004 earthquake, co-  
278 located or close to the stations that recorded the 1966 earthquake (Figure 1). Because station

279 CH8W, which recorded the 1966 earthquake, did not record the 2004 earthquake, we replaced it  
280 by station CH6W. In order to keep the stations in this limited dataset reasonably equidistant, we  
281 then also replaced station CH5W by station CH4AW. We do not use the fault-parallel component  
282 of ground motion recorded at station CH2W, in order to better reproduce the 1966 dataset. Station  
283 CH2W was moved to a new location (the two locations of station CH2W differ by only a few  
284 hundred meters) in the time interval between the two mainshocks (A. Shakal, personal  
285 communication, 2006); thus records of the two earthquakes at this station were not obtained under  
286 the exact same conditions. Because in 1966 geodetic measurements were not as accurate as today,  
287 the exact coordinates of station CH2W at the time are unknown. For the purpose of rupture  
288 modeling, it is reasonable to ignore the station relocation and assume that the station has always  
289 been at its present position.

290 Figure 6 shows the slip model obtained for the 2004 earthquake by inversion of the five  
291 stations. This rupture model is very different from the one obtained by inversion of the complete  
292 2004 seismic dataset (Figure 3). The 2004 complete seismic dataset leads to a robust result; the  
293 slip distribution shown in Figure 3 is similar to those obtained by inversion of subsets of seismic  
294 data (*Custódio et al.*, 2005) and compares well with models obtained by other authors using  
295 independent datasets and inversion methods (*Liu et al.*, 2006). The striking differences between  
296 the two 2004 models, inferred from the complete dataset (Figure 3) and from the limited dataset  
297 (Figure 6), indicate that the 1966 dataset has a poor ability to resolve a complex space-time slip  
298 distribution.

299 In spite of the poor resolution of the sparse 1966 dataset, we can gain insight on the  
300 differences between the 1966 and 2004 earthquakes from the comparison of the two mainshock  
301 rupture models as obtained from the similar Cholame Valley datasets (Figure 7). In particular, it

302 is worthwhile noticing that the ground-motion recorded at the five Cholame Valley stations  
303 confirms that the region of the fault that radiated most energy in 1966 is shallower and further to  
304 the southeast than in 2004. We emphasize that the available 1966 seismic dataset cannot discern a  
305 clear slip pattern; the rupture model for the 1966 earthquake (Figure 2) must be understood and  
306 interpreted within its limitations. Therefore, we will not pursue further statistical analysis  
307 regarding the correlation between the 1966 and the 2004 Parkfield earthquakes. Rather, we will  
308 investigate the hypothesis that these two earthquakes ruptured identical fault patches.

### 309 **3.3. Constrained Inversion of the 1966 Earthquake**

310 Do the data permit similar slip distributions for the 1966 and 2004 Parkfield earthquakes?  
311 In order to answer this question, we performed an inversion of the 1966 earthquake where we  
312 constrained the slip amplitude distribution to be within 20% of the 2004 slip amplitude, which is  
313 well resolved. In fact, it is possible to find an adequate model for the 1966 earthquake where the  
314 slip amplitude distribution resembles that of the 2004 earthquake. Figure 8 shows the comparison  
315 between data recorded during the 1966 earthquake and synthetic ground motion generated both  
316 from the constrained model (constrained by the 2004 slip amplitude distribution) and the  
317 unconstrained 1966 model. The synthetic waveforms generated by the two models are very  
318 similar to the observed ground motion at stations CH8W, CH12W and TEMB. At all stations, the  
319 velocity pulses arrive at the correct time and have the correct duration – our algorithm is able to  
320 find the correct phase for the waveforms by adjusting the time source parameters (rupture velocity  
321 and rise time). However, the large amplitude of the velocity pulses recorded at stations CH2W  
322 and CH5W – the stations closest to the fault – are fit only in our unconstrained model. The large  
323 velocity pulse recorded close to the fault during the 1966 event, which quickly attenuates with



324 distance from the fault, requires a significant amount of shallow slip around Gold Hill. Site  
325 effects can account only for a portion of the large velocities recorded close to the fault in 1966.  
326 Indeed, we included in this study an approximation to site effects (*Liu et al.*, 2006) in order to  
327 account for amplification and frequency resonance at the different stations. Due to this  
328 amplification correction, we only try to fit a maximum peak velocity of 20 cm/s at station CH2W,  
329 whereas the recorded peak velocity at this station was 35 cm/s.

### 330 **3.4. Comparison with Independent Data**

331 Inversions of geodetic data (*Segall and Du*, 1993; *Murray and Langbein*, 2006) indicate  
332 that slip in the 1934 and 2004 Parkfield earthquakes occurred mainly on the NW Parkfield fault  
333 section, close to Middle Mountain, whereas in 1966 slip occurred predominantly to the SE,  
334 around Gold Hill. Thus the geodetic studies agree well with our results inferred from strong-  
335 motion data. Like the seismic data, the geodetic data clearly do not permit identical slip  
336 distribution for the 1966 and 2004 earthquakes (*Murray and Langbein*, 2006), and further exclude  
337 the hypothesis that the 1934 and 1966 earthquakes were identical (*Segall and Du*, 1993).

338 Figure 9 shows aftershocks of both the 1966 and 2004 earthquakes (*Thurber et al.*, 2006).  
339 Due to disparities in dataset quality, the 2004 aftershocks are more precisely located (double  
340 difference locations) than the 1966 aftershocks (absolute locations with respect to the assumed  
341 velocity structure). Also, the stations that recorded the 1966 aftershocks were deployed around  
342 Gold Hill, which explains the reduced number of located aftershocks to the NW of the fault  
343 section. Toward the SE, where the 1966 catalog is more complete, 1966 aftershocks extend  
344 further SE than 2004 aftershocks. Some of the 1966 aftershocks are shallow, surrounding the  
345 region of high slip in our model (Figure 9). Aftershocks of the 2004 earthquake overlap

346 background microseismicity and 1966 aftershocks, but specially illuminate seismic spots toward  
347 the NW of the Parkfield fault section. This aftershock pattern agrees with a larger region of slip  
348 toward the NW and with the effect of stress loading due to directivity of a northwestward  
349 propagating rupture in 2004.

350         Teleseismic waves due to the 1922, 1934, 1966 and 2004 were recorded in De Bilt,  
351 Netherlands (*Bakun and McEvilly, 1984; Dost and Haak, 2006*). The seismograms of the four  
352 events are similar, as expected for earthquakes of similar sizes that occurred in the same location  
353 with a similar focal mechanism. From the four Parkfield events, the 1966 earthquake generated  
354 the most dissimilar records of the set (*Dost and Haak, 2006*), indicating first order differences  
355 between the 1966 and 2004 ruptures.

356         *Wu* (1968) studied the 1966 Parkfield earthquake from local strong-motion seismometer  
357 and seismoscope records, regional short-period seismometer records and long-period teleseismic  
358 records. He concluded that the magnitude of the 1966 earthquake as inferred from surface waves  
359 ( $M_S=6.5$ ) is much larger than inferred from other measurements ( $M_b=5.9$ ,  $M_L=5.5$ ). The efficient  
360 excitation of surface waves supports the existence of shallow slip in 1966 that we infer from our  
361 modeling. *Aki* (1968) also inferred that in 1966 most energy was radiated from a shallow depth.  
362 He arrived at this conclusion by combining the dislocation estimated from strong-motion near-  
363 fault records with the seismic moment obtained from long-period surface waves (*Tsai and Aki,*  
364 1969)

#### 365 **4. Implications for Earthquake Prediction**

366         Our results, obtained from the analysis of seismic strong-motion data, strongly suggest

367 that the 1966 and 2004 Parkfield earthquakes did not rupture identical regions of the Parkfield  
368 fault section. Our interpretation that at least to some extent the 1966 and 2004 ruptures  
369 complement each other is supported by geodetic studies, aftershock locations, comparison of  
370 teleseismic waveforms and surface wave observations. An identical behavior of earthquakes that  
371 rupture complementary zones of the same fault plane is observed in Papua New Guinea (*Park and*  
372 *Mori, in press*). *Okada et al. (2005)* studied the most recent (post 1930's) moderate and large  
373 earthquakes in Miyagi-Oki, Japan, and proposed that "*several asperities exist offshore of Miyagi*  
374 *Prefecture, and (...) those asperities rupture repeatedly at sometime simultaneously and other*  
375 *time separately*". On a larger scale, similar patterns are observed in subduction zones (e.g., Alaska  
376 (*Sykes, 1971*), Chile (*Comte et al., 1986; Campos et al., 2002*)). Here, different fault sections  
377 rupture individually or together, complementing previous ruptures, so that eventually the entire  
378 subduction zone has slipped co-seismically.

379         The persistence of microseismicity in Parkfield suggests that intrinsic fault properties (e.g.  
380 rheology, geometry) control the seismicity. This observation implies the existence of asperities  
381 (regions of the fault that are strongly coupled, i.e., frictionally locked) that persist over multiple  
382 earthquake cycles. To reconcile the existence of persistent asperities with our observation of  
383 different slip distributions in consecutive Parkfield earthquakes, we propose a scenario where the  
384 fault contains different asperities, each of which has its own frictional properties and thus ruptures  
385 at a different time. Within this framework we hypothesize that if all asperities are synchronized,  
386 i.e. simultaneously loaded, then a major earthquake will happen. We can speculate that this was  
387 the case in 1857, when two Parkfield earthquakes (~M5.6 and ~M6.1) shortly preceded the great  
388 Fort Tejon earthquake by a few hours (*Sieh, 1978b*). The ~M8 Fort Tejon earthquake was the last  
389 large earthquake to rupture the south-central San Andreas Fault. It nucleated in Parkfield

390 (Cholame Valley) and ruptured ~400 km south until Wrightwood (*Sieh*, 1978a).

391           The Parkfield Experiment yielded a dataset of exceptional quality for the 2004 earthquake;  
392 we expect that the next Parkfield earthquakes will generate at least equally good datasets.  
393 Parkfield remains as a very good location to trap ~ $M_w$ 6 earthquakes, and given the knowledge we  
394 already have of past Parkfield events, a thorough recording and study of future events will  
395 certainly contribute significantly to our understanding of seismicity and loading patterns. The  
396 Parkfield Experiment remains of great value to the physical understanding of earthquake  
397 interaction, and therefore to earthquake forecast.

## 398 **5. Conclusions**

399           We used strong ground-motion recorded by five near-field seismographs to infer a  
400 kinematic rupture model for the 1966  $M_w$ 6 Parkfield earthquake. Because this earthquake was  
401 recorded by a very limited number of instruments, we conducted a number of tests to assess the  
402 resolution of our rupture model. In particular, we focused on the hypothesis of similar slip  
403 distributions in the 1966 and 2004 earthquakes. We concluded that slip in the two most recent  
404 Parkfield events (1966 and 2004) did not occur in a characteristic manner, but rather occurred in a  
405 complementary way. The most robust result of our analysis is that slip in the 1966 earthquake  
406 occurred further SE than in the 2004 earthquake. In order to explain the complementary character  
407 of the 1966 and 2004 slip distributions, while taking in account the characteristic behavior of  
408 microseismicity in Parkfield, we propose the existence of several asperities along the Parkfield  
409 section of the San Andreas Fault, each of which having its own frictional properties. In this  
410 scenario, characteristic earthquakes may occur if the same asperities reach their yield stress  
411 simultaneously. If all the asperities on the fault reach their yield stress at approximately the same

412 time, then a larger earthquake may result.

413

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420 1008 and ICS contribution number 0734.

421

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523

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528 (blue arrow). The fault planes modeled for the 1966 and 2004 earthquakes correspond to the red  
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533 plane used to model the 2004 earthquake with a subset of data equivalent to the 1966 data set.  
534 MM - Middle Mountain; GH - Gold Hill; CH - Cholame.

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537 earthquake (1966 and 2004), the fault plane is buried 500 m below the surface. The red star marks  
538 the 1966 hypocenter. A) Best model (model that best fits the data); B) Average of 10 best models;  
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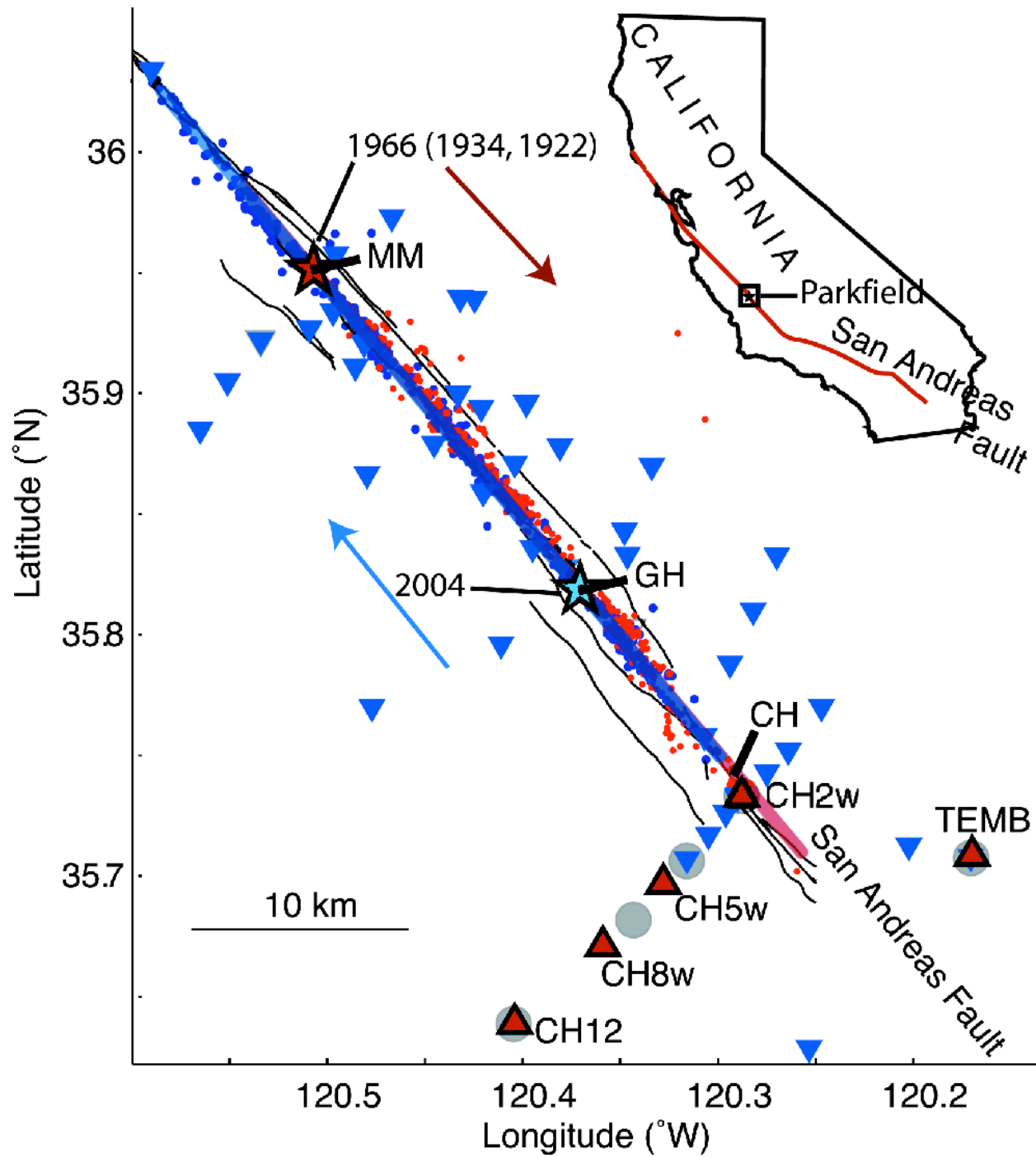
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588 **Figure 9.** Comparison between the rupture models for the 1966 and 2004 Parkfield earthquakes  
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591 aftershocks of the 2004 earthquake (blue circles) and background seismicity from 1984 to the  
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593 computed assuming a 3 MPa stress drop in a circular region. In the absence of information on the  
594 magnitudes of the 1966 aftershocks, we cannot compute their size; these aftershocks are  
595 represented by crosses. The rectangle indicates the position of the fault plane modeled for the  
596 2004 earthquake. The red and blue stars mark the 1966 and 2004 hypocenters, respectively. MM -  
597 Middle Mountain; GH - Gold Hill; CH - Cholame.

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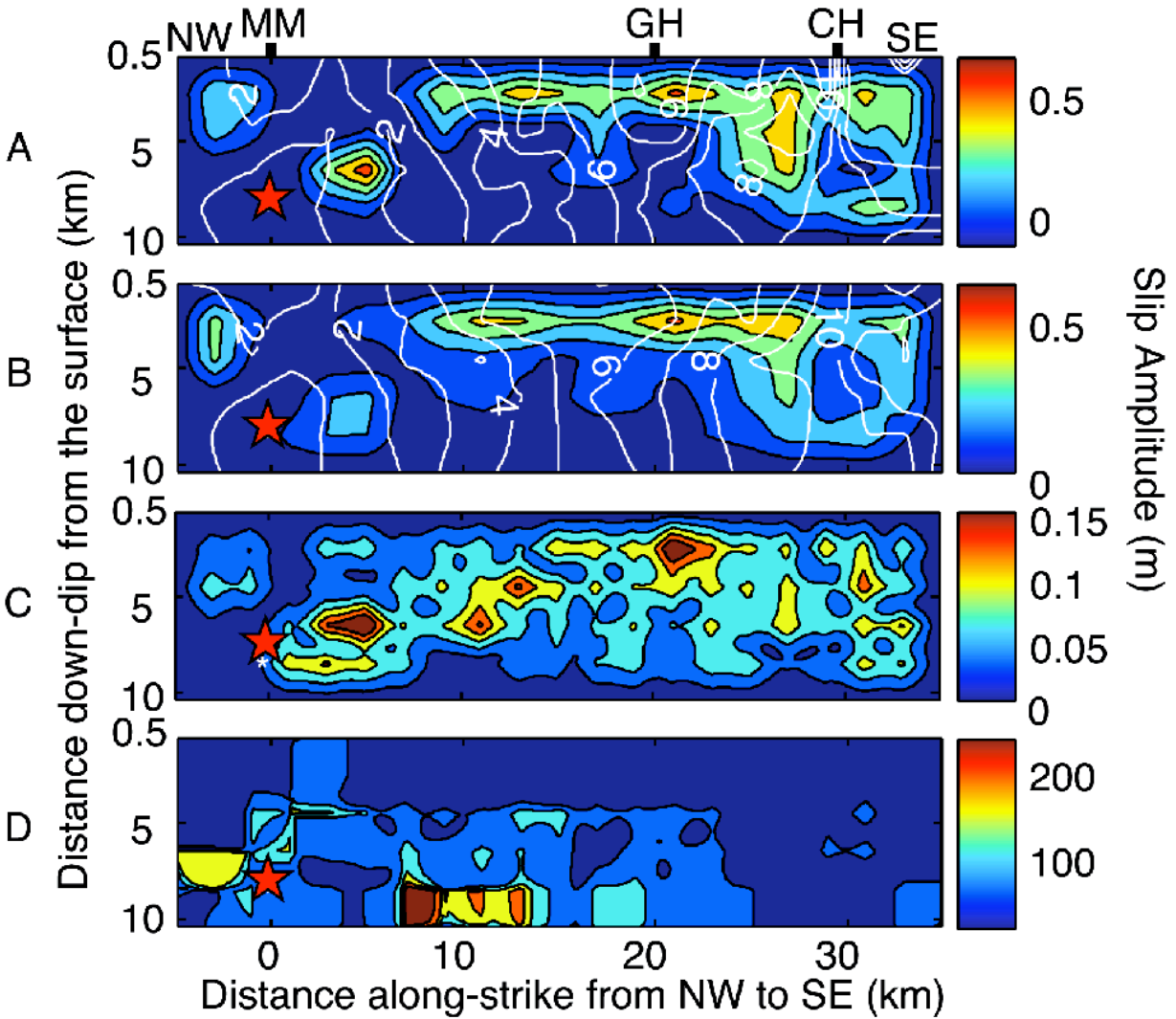


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614 either earthquake (1966 and 2004), the fault plane is buried 500 m below the surface. The red star

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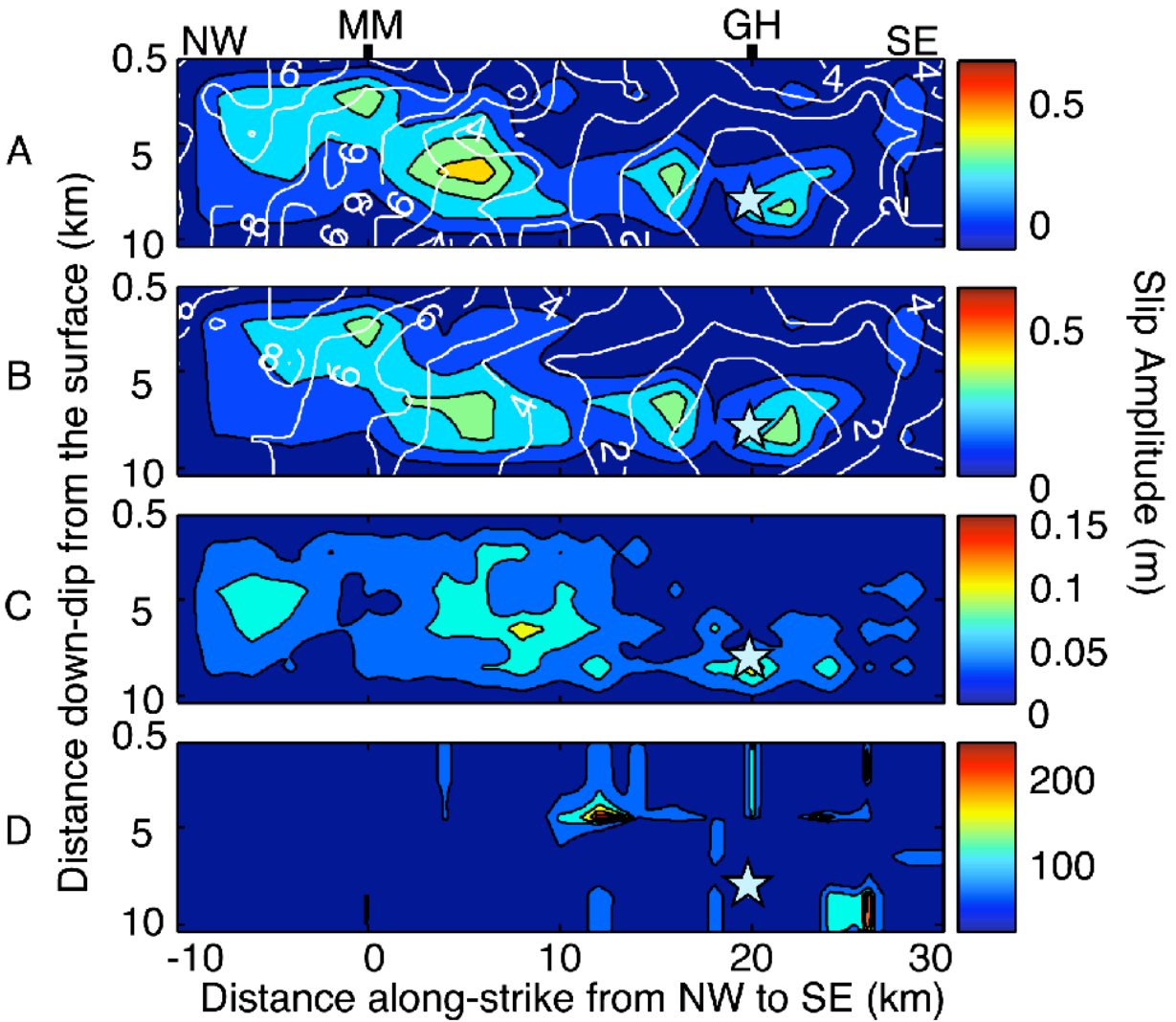
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621

622 **Figure 3.** (1-column) Rupture model for the 2004 Parkfield earthquake – slip amplitude (color

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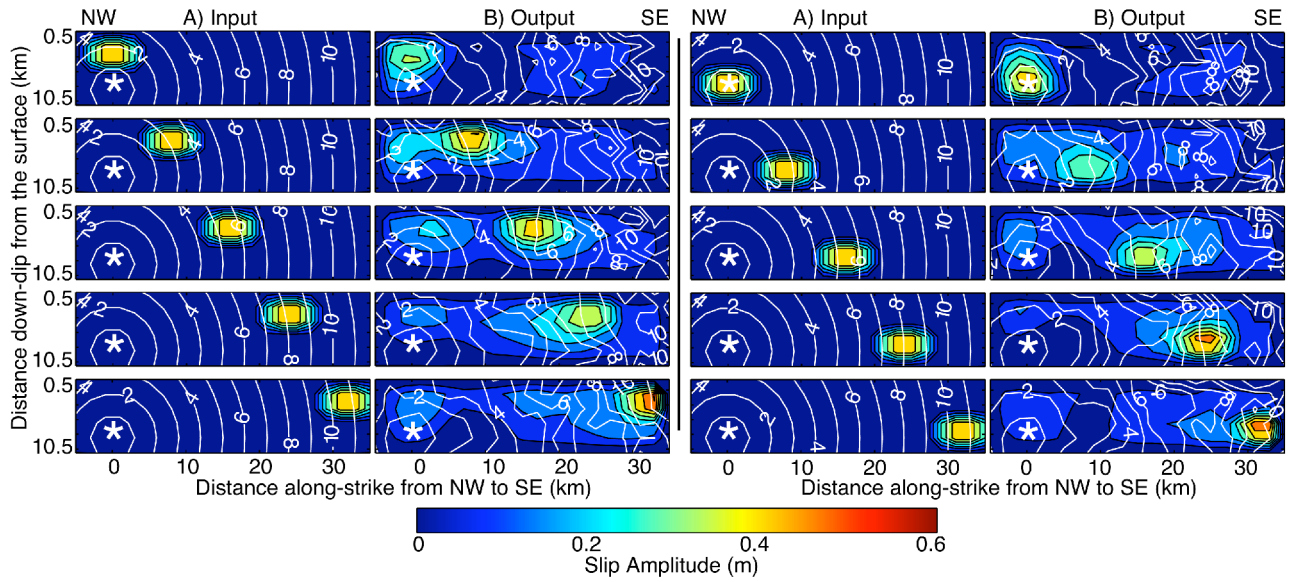
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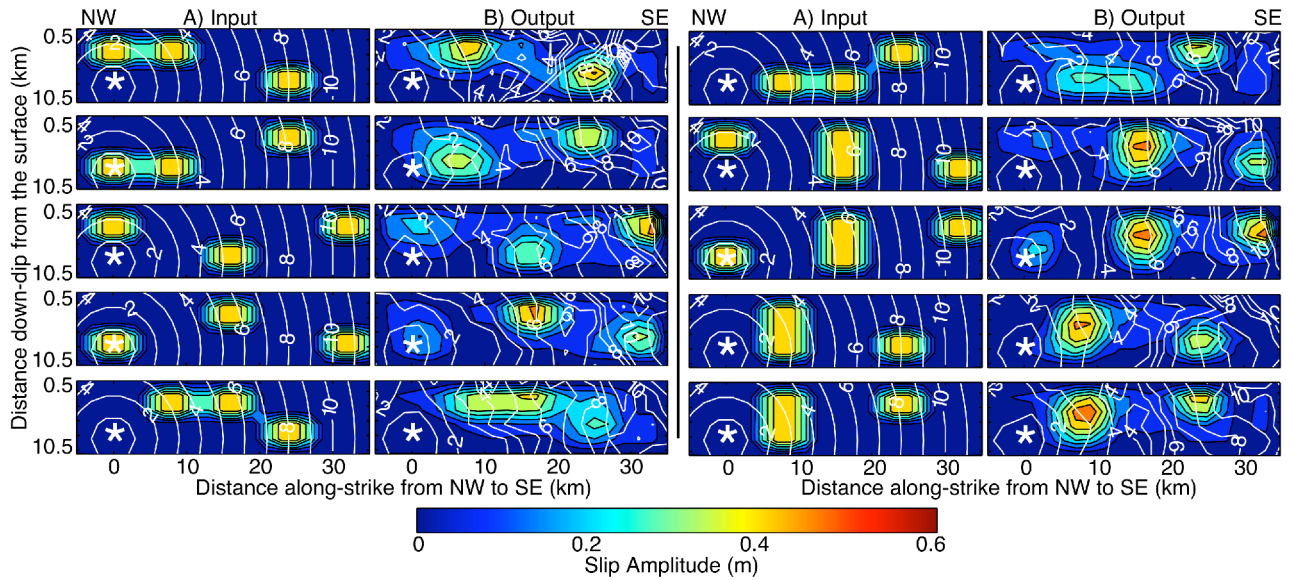
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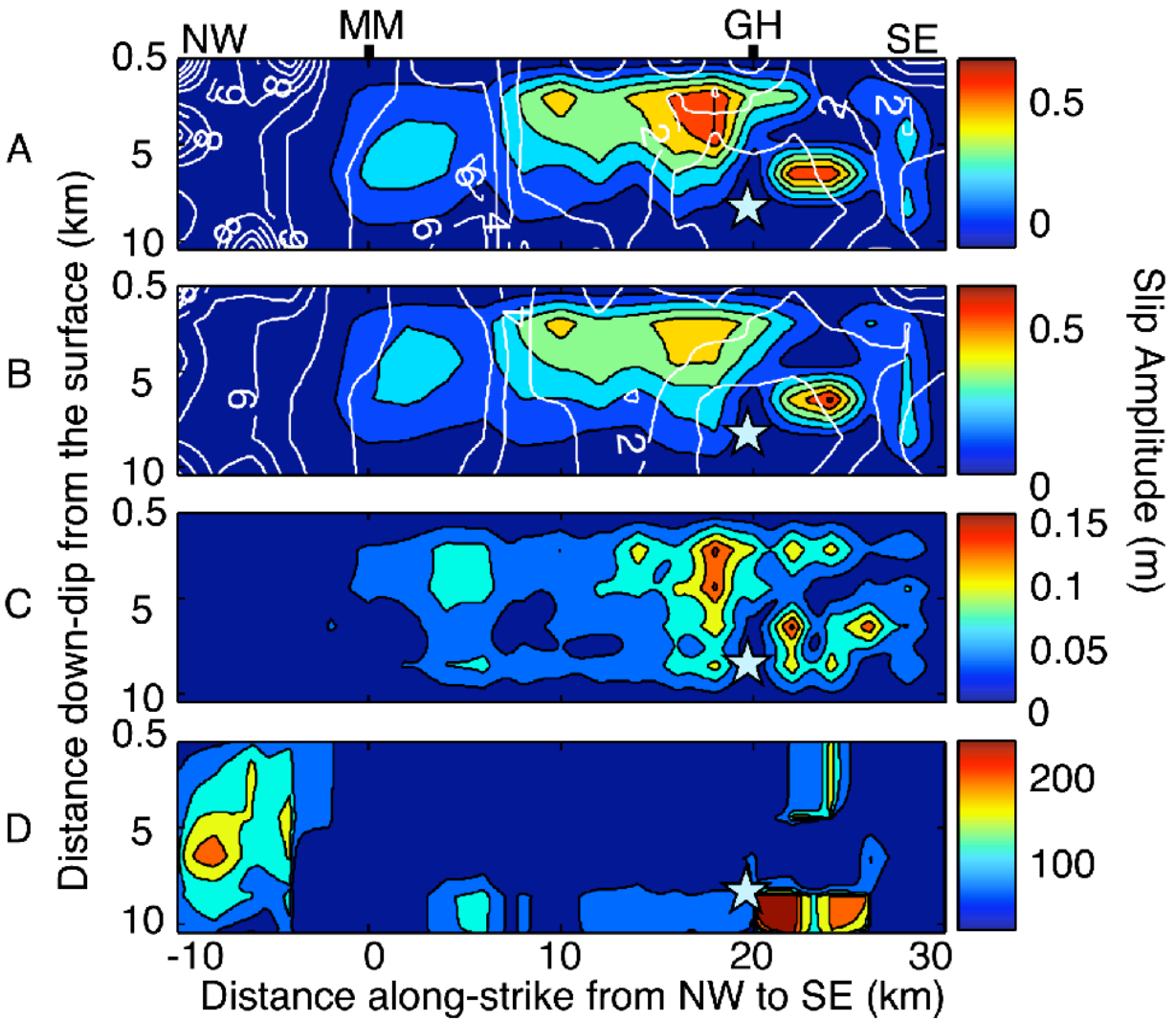
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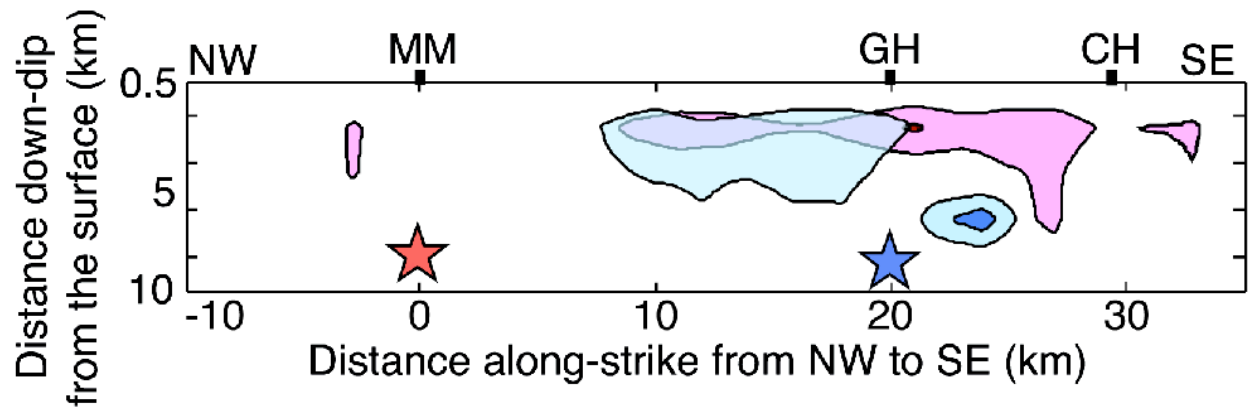
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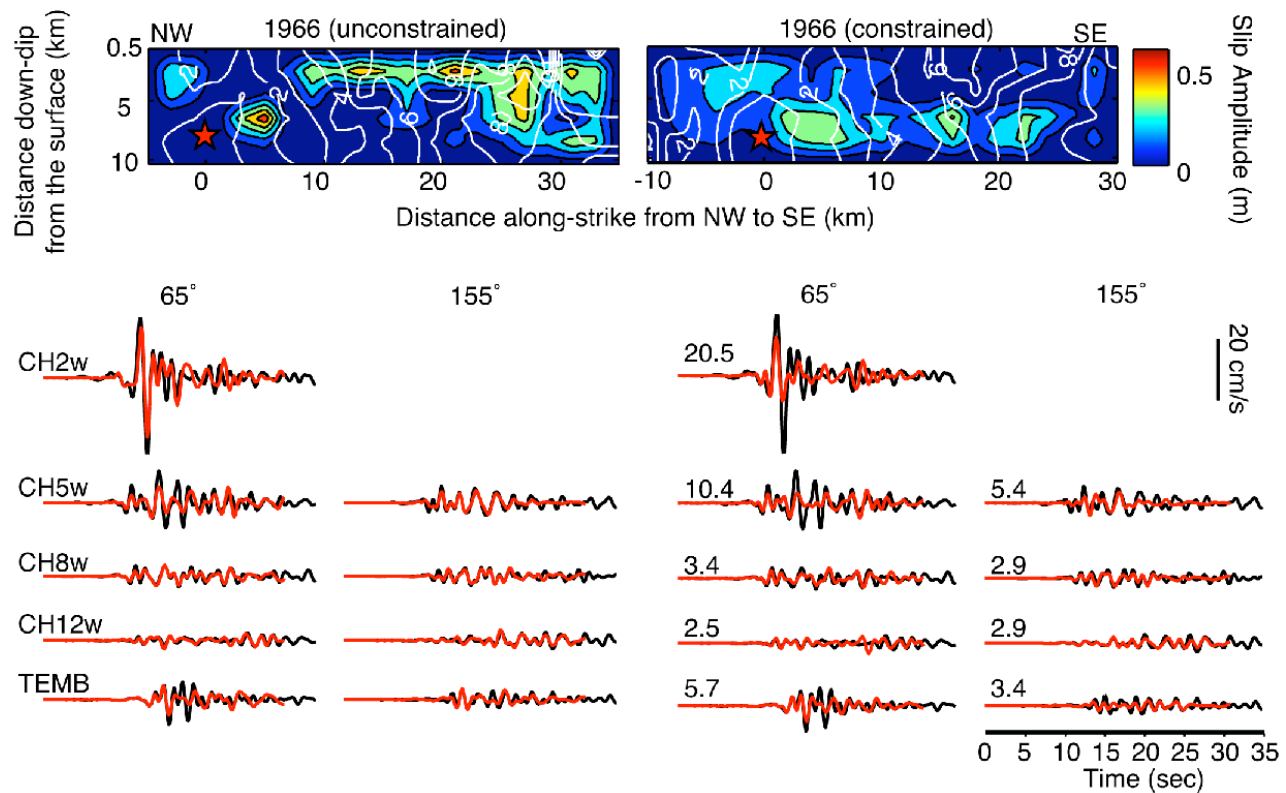
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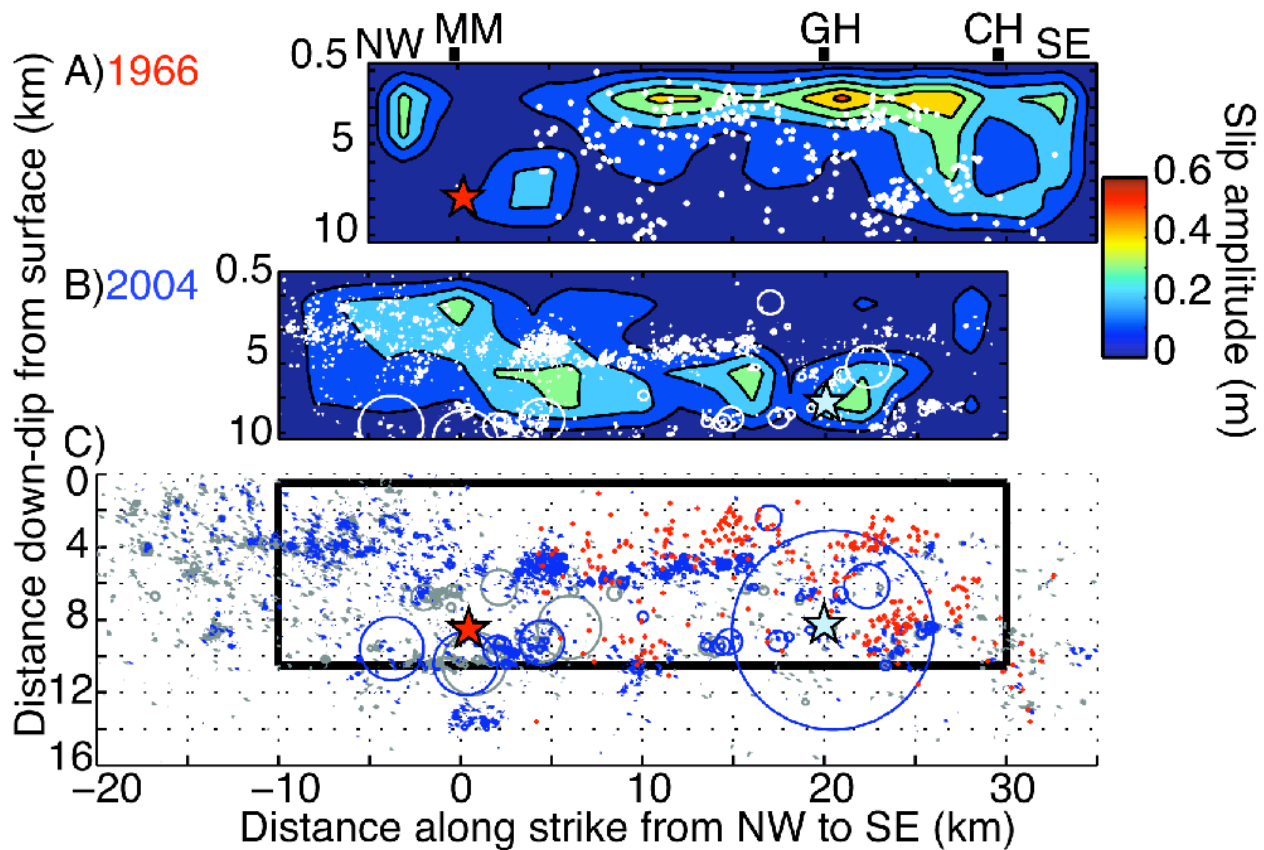


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677 **Figure 9.** (1-column) Comparison between the rupture models for the 1966 and 2004 Parkfield  
 678 earthquakes and microseismicity. A) Slip amplitude and aftershocks of the 1966 earthquake. B)  
 679 Slip amplitude and aftershocks of the 2004 earthquake. C) Aftershocks of the 1966 earthquake  
 680 (red crosses), aftershocks of the 2004 earthquake (blue circles) and background seismicity from  
 681 1984 to the 2004 earthquake (gray circles) (*Thurber et al.*, in press). The size of the aftershocks  
 682 (circles) is computed assuming a 3 MPa stress drop in a circular region. In the absence of  
 683 information on the magnitudes of the 1966 aftershocks, we cannot compute their size; these  
 684 aftershocks are represented by crosses. The rectangle indicates the position of the fault plane  
 685 modeled for the 2004 earthquake. The red and blue stars mark the 1966 and 2004 hypocenters,  
 686 respectively. MM - Middle Mountain; GH - Gold Hill; CH - Cholame.

687



711 1. fs01.tif (Figure S1) Ten preferred rupture models for the 1966  
712 Parkfield earthquake. The color scale indicates slip amplitude  
713 (m), and the white lines show 1-sec rupture time contours. The m-  
714 value above each model indicates the numerical misfit between  
715 observed ground motion and synthetic ground motion generated by  
716 each rupture model. All models are equally adequate, in the sense  
717 that they produce very similar misfits.

718  
719 2. ts01.txt-ts10.txt Ten preferred rupture models for the 1966  
720 Parkfield earthquake, given at a grid spacing of 2km x 2km (grid  
721 spacing at which we invert for source parameters). 2.1 Column  
722 "x", km, position of the grid node along-strike from NW to SE  
723 from the NW end of the fault plane. 2.2 Column "y", km, position  
724 of the grid node down-dip from the surface. 2.3 Column "amp", m,  
725 slip amplitude. 2.4 Column "rake", degrees, slip rake. 2.5 Column  
726 "rup\_vel", km/s, average rupture velocity. 2.6 Column "rise1",  
727 sec, rise time during which slip accelerates. 2.7 Column "rise2",  
728 sec, rise time during which slip decelerates.

729  
730 3. ts11.txt-ts20.txt Ten preferred rupture models for the 1966  
731 Parkfield earthquake, given at a grid spacing of 167m x 167m  
732 (grid spacing at which source parameters are convolved with  
733 Green's functions in order to generate synthetic ground motion).  
734 2.1 Column "x", km, position of the grid node along-strike from  
735 NW to SE from the NW end of the fault plane. 2.2 Column "y", km,

736 position of the grid node down-dip from the surface. 2.3 Column  
737 "amp", m, slip amplitude. 2.4 Column "rake", degrees, slip rake.  
738 2.5 Column "rup\_vel", km/s, average rupture velocity. 2.6 Column  
739 "rise1", sec, rise time during which slip accelerates. 2.7 Column  
740 "rise2", sec, rise time during which slip decelerates.

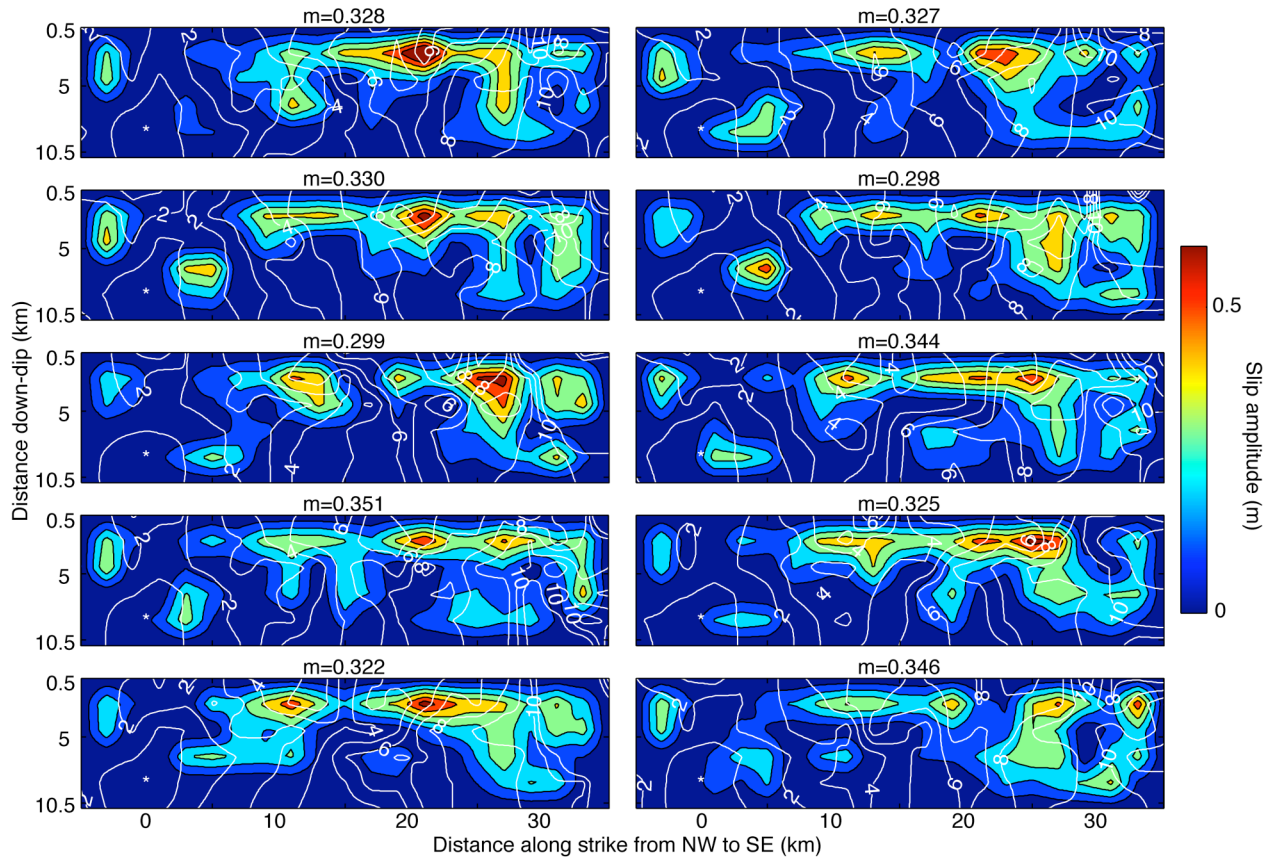
741  
742 4. fs02.tif (Figure S2) Ten preferred rupture models for the 2004  
743 Parkfield earthquake. The color scale indicates slip amplitude  
744 (m), and the white lines show 1-sec rupture time contours. The m-  
745 value above each model indicates the numerical misfit between  
746 observed ground motion and synthetic ground motion generated by  
747 each rupture model. All models are equally adequate, in the sense  
748 that they produce very similar misfits.

749  
750 5. ts21.txt-ts30.txt Ten preferred rupture models for the 2004  
751 Parkfield earthquake, given at a grid spacing of 2km x 2km (grid  
752 spacing at which we invert for source parameters). 2.1 Column  
753 "x", km, position of the grid node along-strike from NW to SE  
754 from the NW end of the fault plane. 2.2 Column "y", km, position  
755 of the grid node down-dip from the surface. 2.3 Column "amp", m,  
756 slip amplitude. 2.4 Column "rake", degrees, slip rake. 2.5 Column  
757 "rup\_vel", km/s, average rupture velocity. 2.6 Column "rise1",  
758 sec, rise time during which slip accelerates. 2.7 Column "rise2",  
759 sec, rise time during which slip decelerates.

760



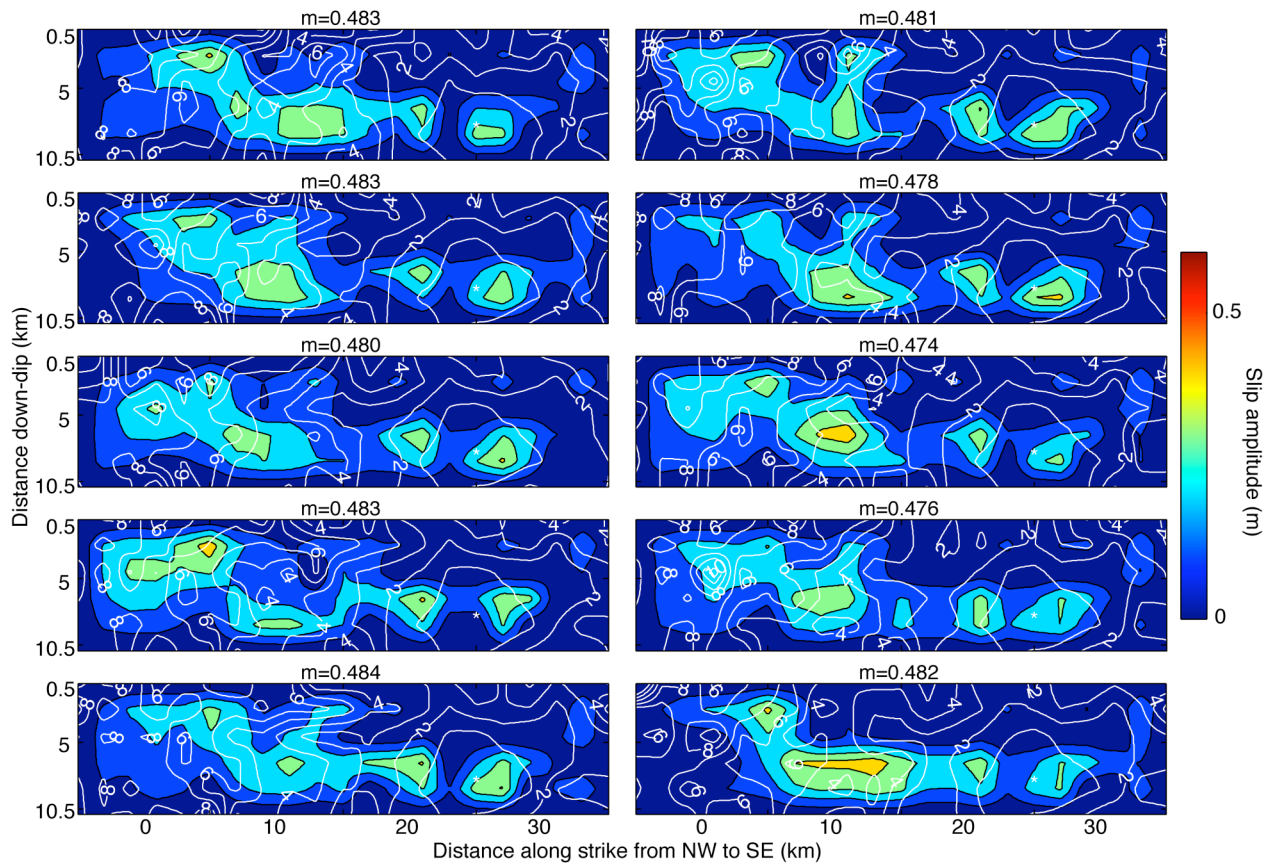
761 6. ts31.txt-ts40.txt Ten preferred rupture models for the 2004  
762 Parkfield earthquake, given at a grid spacing of 167m x 167m  
763 (grid spacing at which source parameters are convolved with  
764 Green's functions in order to generate synthetic ground motion).  
765 2.1 Column "x", km, position of the grid node along-strike from  
766 NW to SE from the NW end of the fault plane. 2.2 Column "y", km,  
767 position of the grid node down-dip from the surface. 2.3 Column  
768 "amp", m, slip amplitude. 2.4 Column "rake", degrees, slip rake.  
769 2.5 Column "rup\_vel", km/s, average rupture velocity. 2.6 Column  
770 "rise1", sec, rise time during which slip accelerates. 2.7 Column  
771 "rise2", sec, rise time during which slip decelerates.  
772  
773



773

774 Figure fs01 (fs01.tif)

775



775

776 Figure fs02 (fs02.tif)

777